

European
Innovation
Council



UK Research
and Innovation

C L U D



deepening further into reactor neutrino physics



**TWENTY-SECOND LOMONOSOV
CONFERENCE** August, 21-27, 2025
ON ELEMENTARY PARTICLE PHYSICS
MOSCOW STATE UNIVERSITY

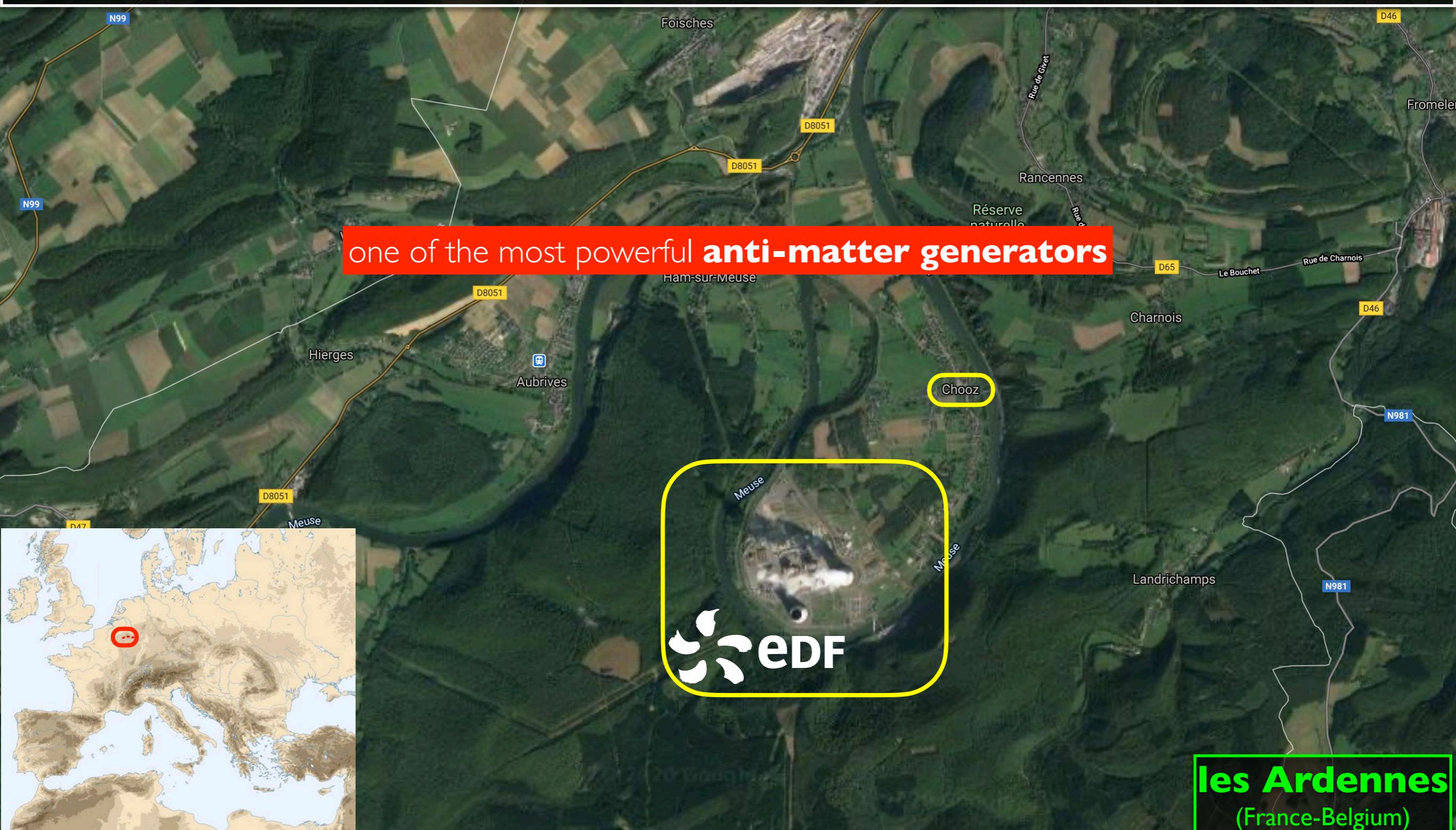
Lomonosov Conference
27 August 2025 — Moscow (Russia)

Anatael Cabrera

Université Paris-Saclay / CNRS — IJCLab / LNCA
Orsay/Chooz, France

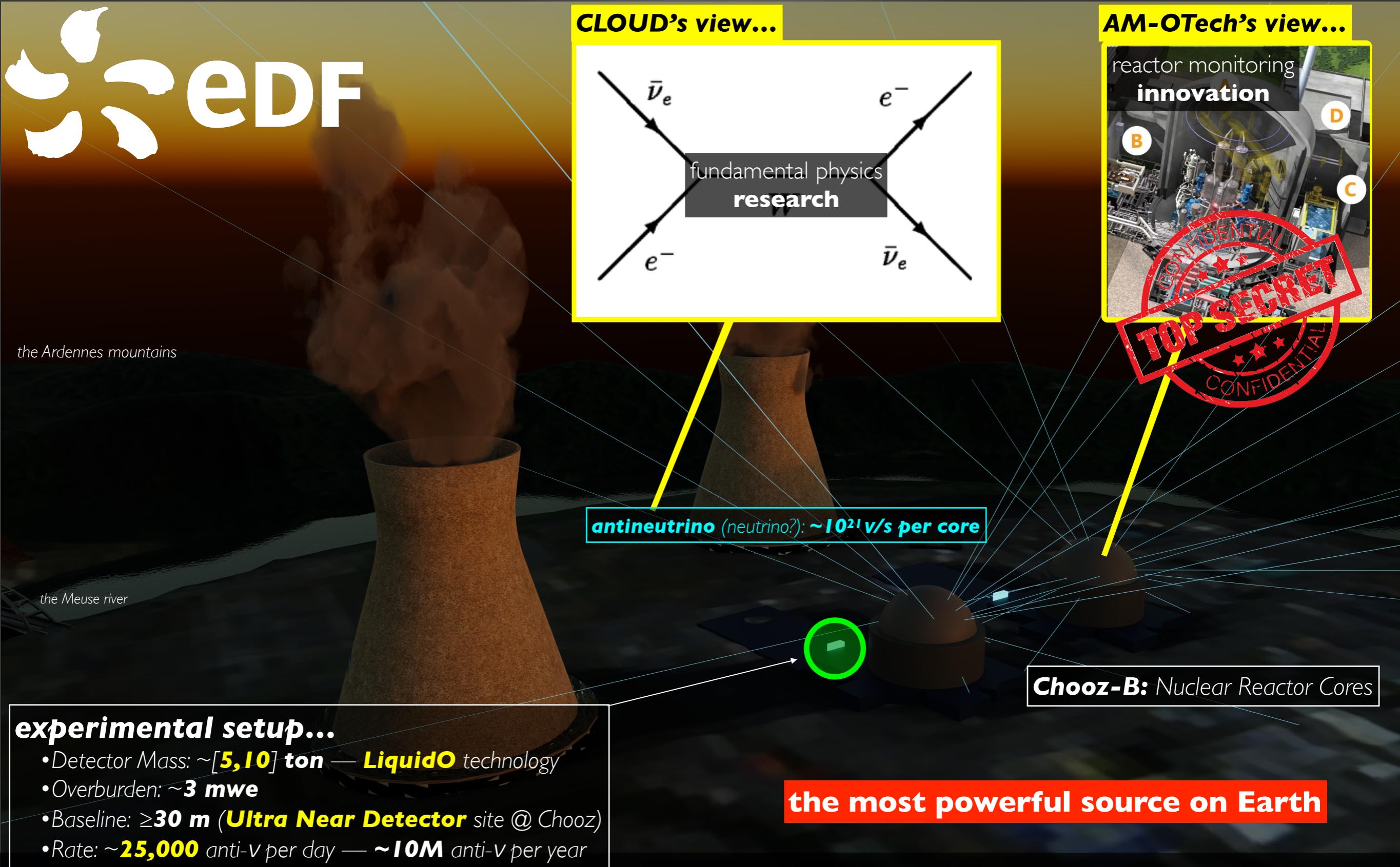
(co)spokesperson:
•DoubleChooz
•LiquidO
•CLOUD — AM-OTech (EIC)
•SuperChooz Pathfinder

Gimnée
in the **middle of central Europe** (between France-Belgium): **Chooz** [meeting point with Germany, Luxembourg, Netherlands]



Europe's most powerful reactor site...

3rd generation of reactor neutrino experiments @ Chooz



CLOUD vs AntiMatter-OTech...

C L U D

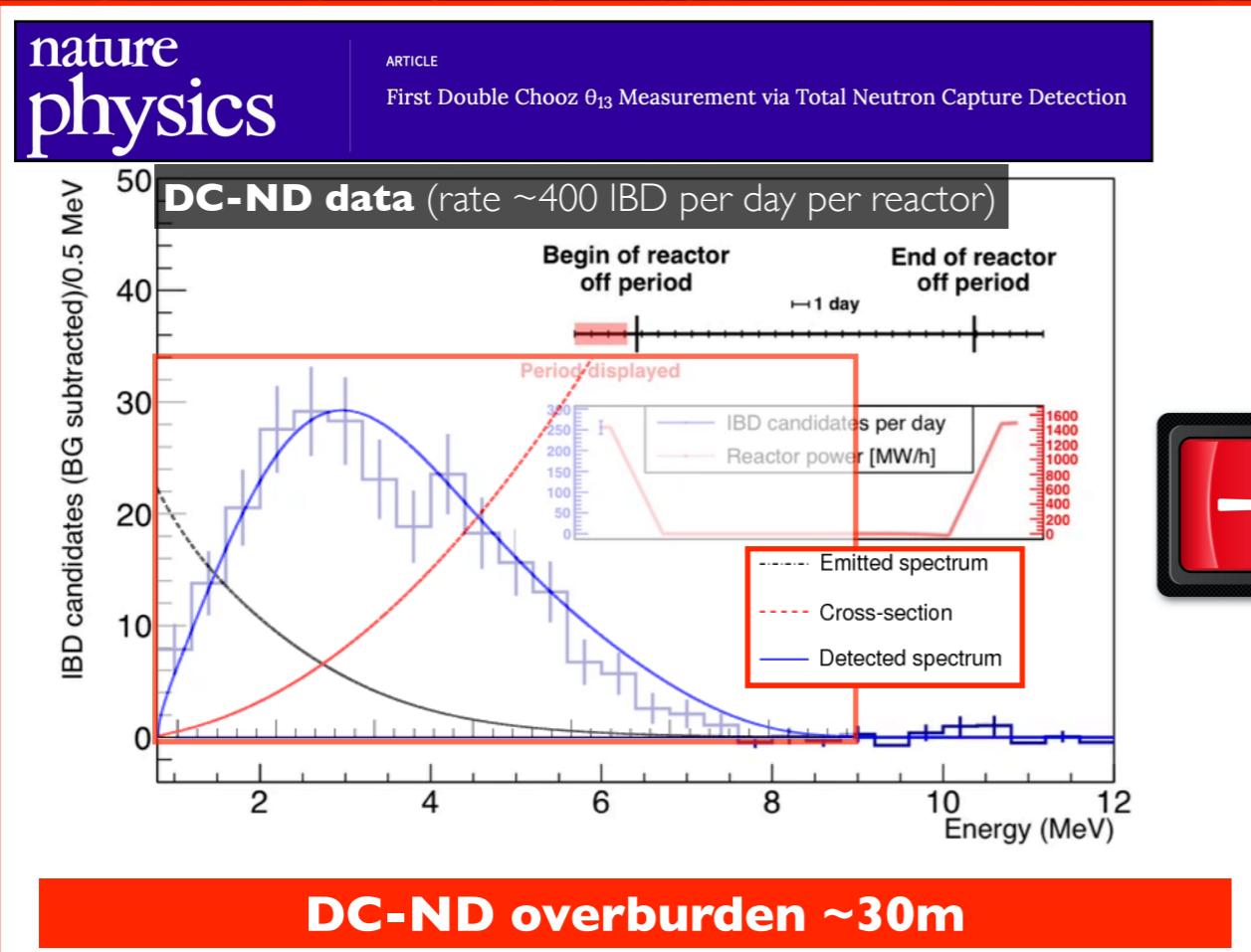


reactor neutrinos **experimental methodology largely similar** for the last **~70 years** (Reines et al.)

powerful framework so far, but **good enough for discoveries $\geq 2025?$**

on the shoulders of giants...

reactor neutrinos...



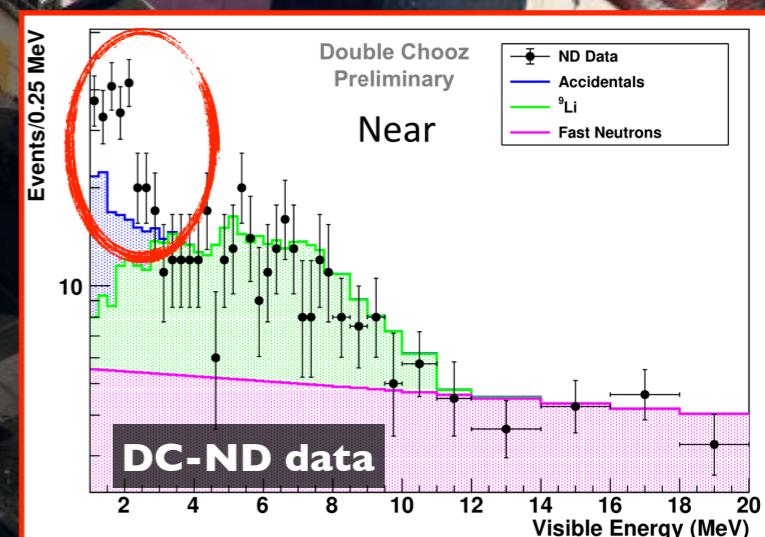
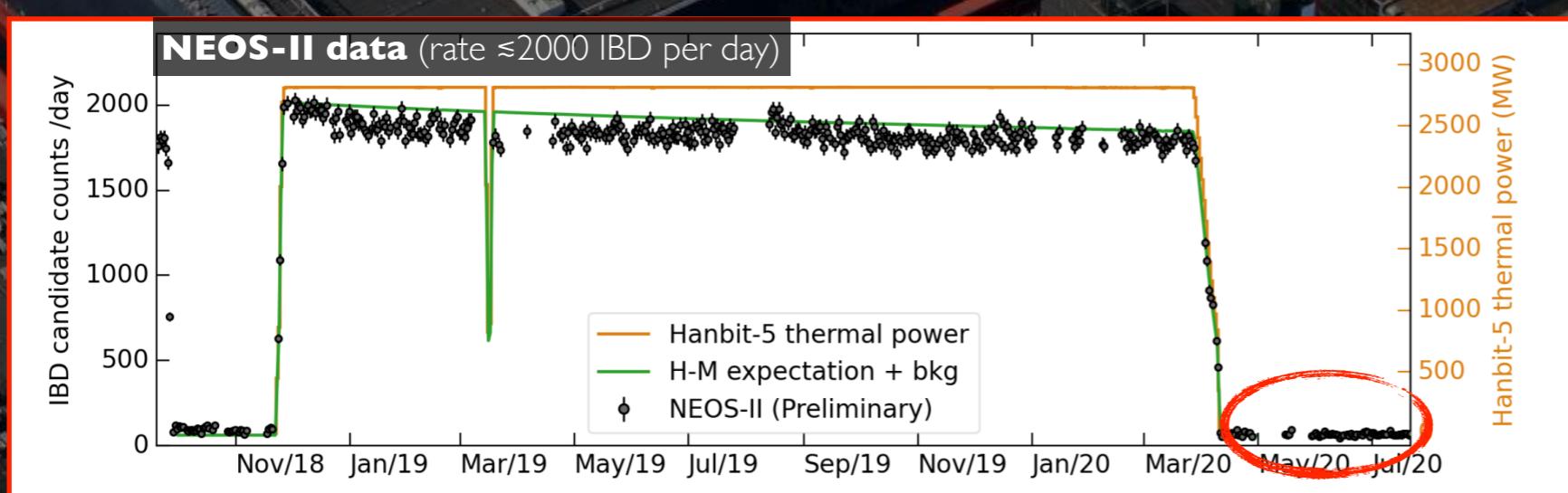
powerful experimental framework

- IBD(p): one interaction (CC)

→ NC still very difficult

- flux: several issues!

status: issues vs physics?



future: discoveries?

today's BG: standard neutrino oscillation

neutrino \oplus weak-interaction remains **bizarre** (Majorana, etc)...

- new neutrino **phenomenology**? [ex. mixing and masses]
- new neutrino **interactions**?
- new neutrino **states**? [assume: “3+1 sterile” is largely ruled out]

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a probe to the future?
(let's try)

the experiment...

today, **most experiments** bypass (whenever possible) the **absolute flux knowledge** — complex!
relative knowledge (ex. multi-detector; etc.) well suited to **extract “known model” parameters**

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confront **absolute flux knowledge** for **new neutrino physics** via “flux bias explorations”

- **extreme signal** (to BG) rates — unprecedented
- must: the **best-known cross-section**(s) today ($\leq 1\%$)
- **extreme energy control** ($\leq 1\%$) — avoid spectral distortions (\rightarrow flux biases)
- **much redundancy** — as much as Nature kindly allows...

Chooz-B Power Station

- facility: EDF CNPE
- location: Chooz (France)
- reactor cores: 2x PWR AREVA-N4
- thermal power: 8.4GW (total)

Double Chooz
Near Detector

LNCA-Hall (CNRS)

Ultra Near Detector (UND) sites



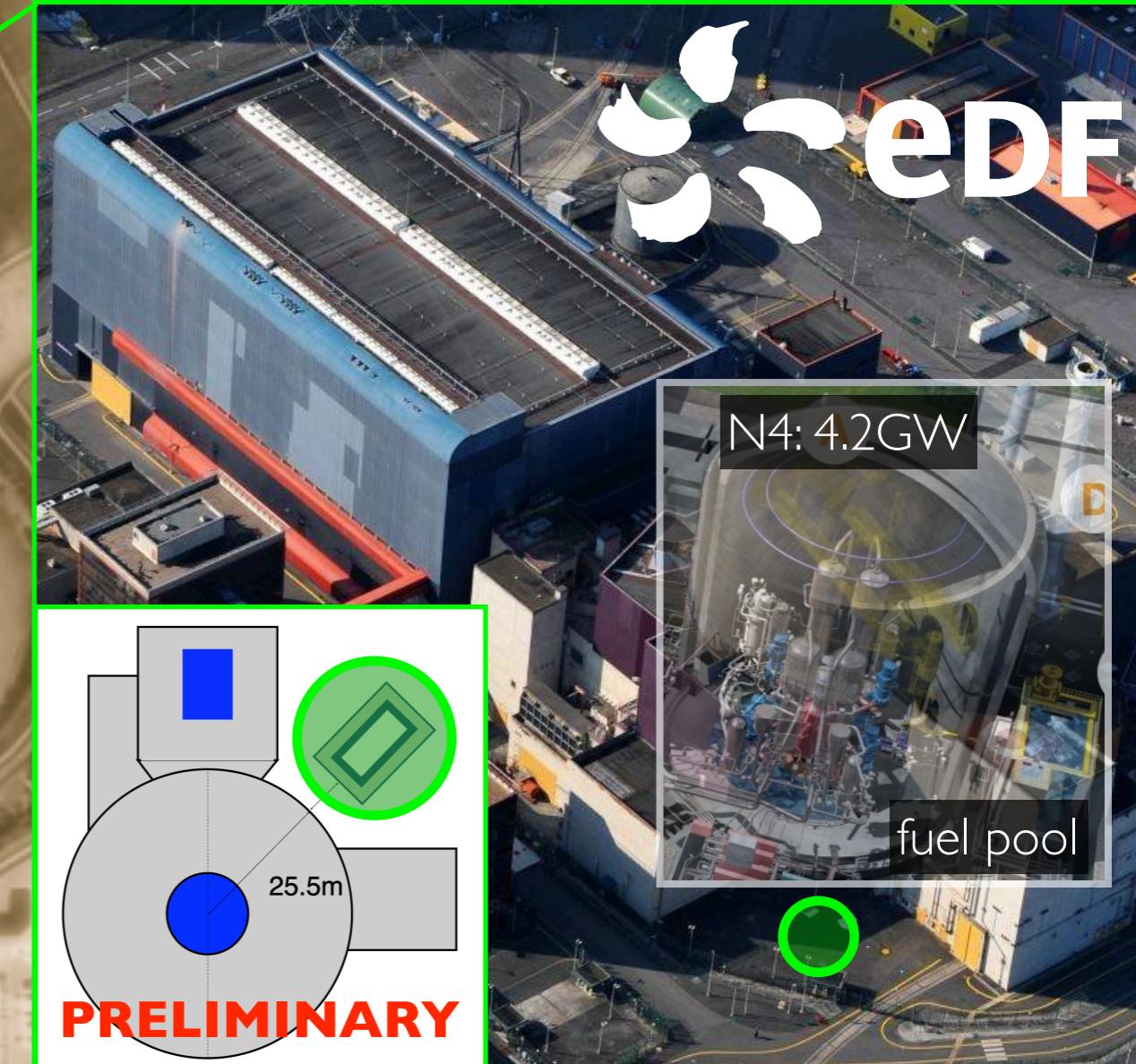
OFF

ON

due to global warm → more frequent reactor-OFF (2022: several months)

CLOUD = “Chooz Liquido Ultranear Detector”

Double Chooz
Far Detector

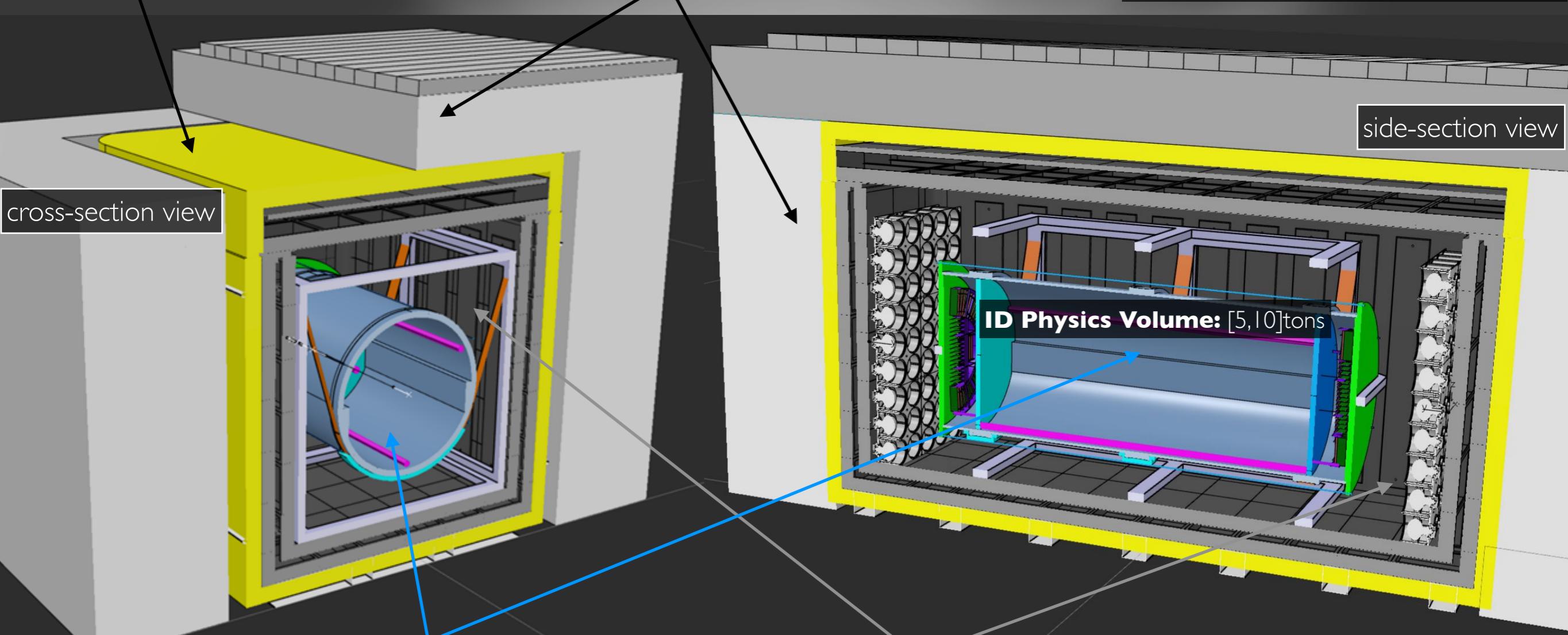


Europe's best reactor-V site...

Water Pool [20,40]cm thick
 • 4π shield & neutron moderator
 •controllable thermal-bath

IGLOO [~ 3 mwe]
 •concrete **bunker** (with boron?)
 •DC's iron steel shield (15cm thick)

Redundant “surface neutron” layers...
 •**IGLOO** (absorption) — passive
 •**Water** (moderator⊕absorption) — passive
 •**Armour** (veto⊕moderator⊕absorption) — **OD**
 •**Tracker** (PID⊕moderator) — **ID**



LiquidO-Tracker (or inner-detector) [≤ 10 tons fiducial]
 •opaque scintillator — new formulation(s) [more on this soon]
 • $\sim 10,000$ fibres⊕SiPM readout channels (GHz waveforms)
 •designed light level: ≥ 200 pe/MeV

ARMOUR (or outer-detector) [~ 0.5 m thickness]
 •transparent scintillator (LAB⊕PPO⊕Bis-MSB)
 • ≤ 180 DC-PMTs & highly reflecting walls
 •designed light yield ≥ 400 pe/MeV

CLOUD detector...

experimental setup...
 •Detector Mass: $\sim [5,10]$ ton — **LiquidO** technology
 •Overburden: ~ 3 mwe
 •Baseline: ≥ 30 m (**Ultra Near Detector** site @ Chooz)
 •Rate: $\sim 25,000$ anti- ν per day — $\sim 10M$ anti- ν per year

CLOUD is powered by...

L | Q U | D



nature communications physics

Article | [Open access](#) | Published: 21 December 2021

Neutrino physics with an opaque detector

[LiquidO Consortium](#)

[Communications Physics](#) 4, Article number: 273 (2021) | [Cite this article](#)

8831 Accesses | 27 Altmetric | [Metrics](#)

Abstract

Publication “LiquidO-Zero”

In 1956 Reines & Cowan discovered the neutrino using a liquid scintillator detector. The neutrinos interacted with the scintillator, producing light that propagated across transparent volumes to surrounding photo-sensors. This approach has remained one of the most widespread and successful neutrino detection technologies used since. This article introduces a concept that breaks with the conventional paradigm of transparency by confining and collecting light near its creation point with an opaque scintillator and a dense array of optical fibres. This technique, called LiquidO, can provide high-resolution imaging to enable efficient identification of individual particles event-by-event. A natural affinity for adding dopants at high concentrations is provided by the use of an opaque medium. With these and other capabilities, the potential of our detector concept to unlock opportunities in neutrino physics is presented here, alongside the results of the first experimental validation.

www.nature.com/articles/s42005-021-00763-5

LiquidO Official WEB: <https://liquidoo.ijclab.in2p3.fr/>

L I Q U I D O

XXX Neutrino Conference
June 2022 — Seoul, South Korea



Anatael Cabrera
CNRS/IN2P3
IJCLab/Université Paris-Saclay
(Orsay)



<https://zenodo.org/record/6697273>

FNAL Seminar 2023

(May 2023)

thanks to the LiquidO consortium...

L I Q U I D O

Detection and Imaging in Opaque Media

Neutrino Seminar @ FNAL
4th May 2023 — Chicago, USA



Anatael Cabrera
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(Orsay)

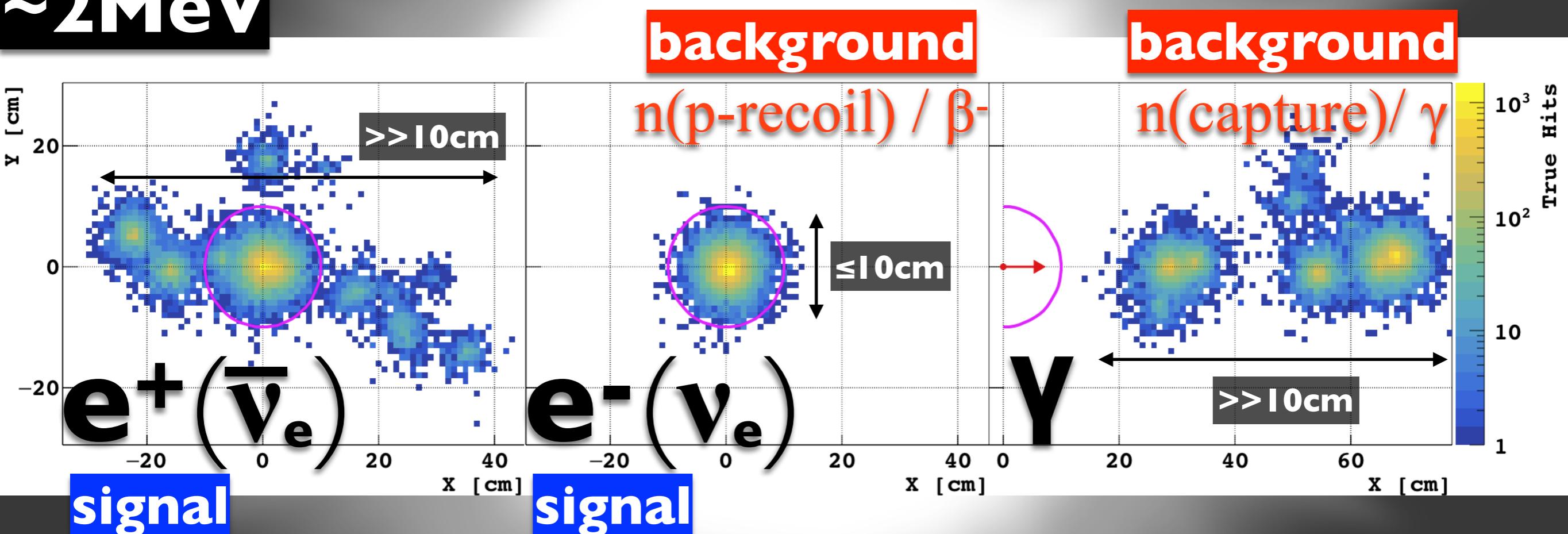


<https://zenodo.org/records/7922021>

unprecedented MeV imaging. . .

reduce overburden/shielding

~2MeV

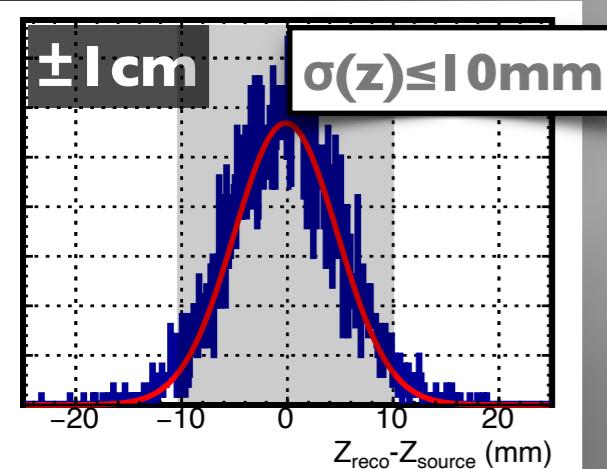
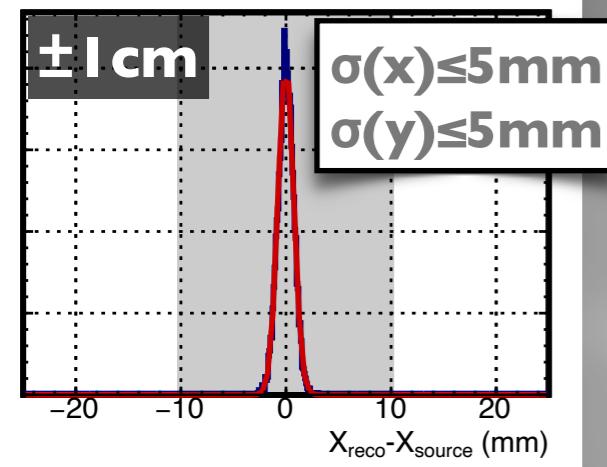


LiquidO: stochastic confinement (NO segmentation)

LiquidO \leftrightarrow stochastic light confinement

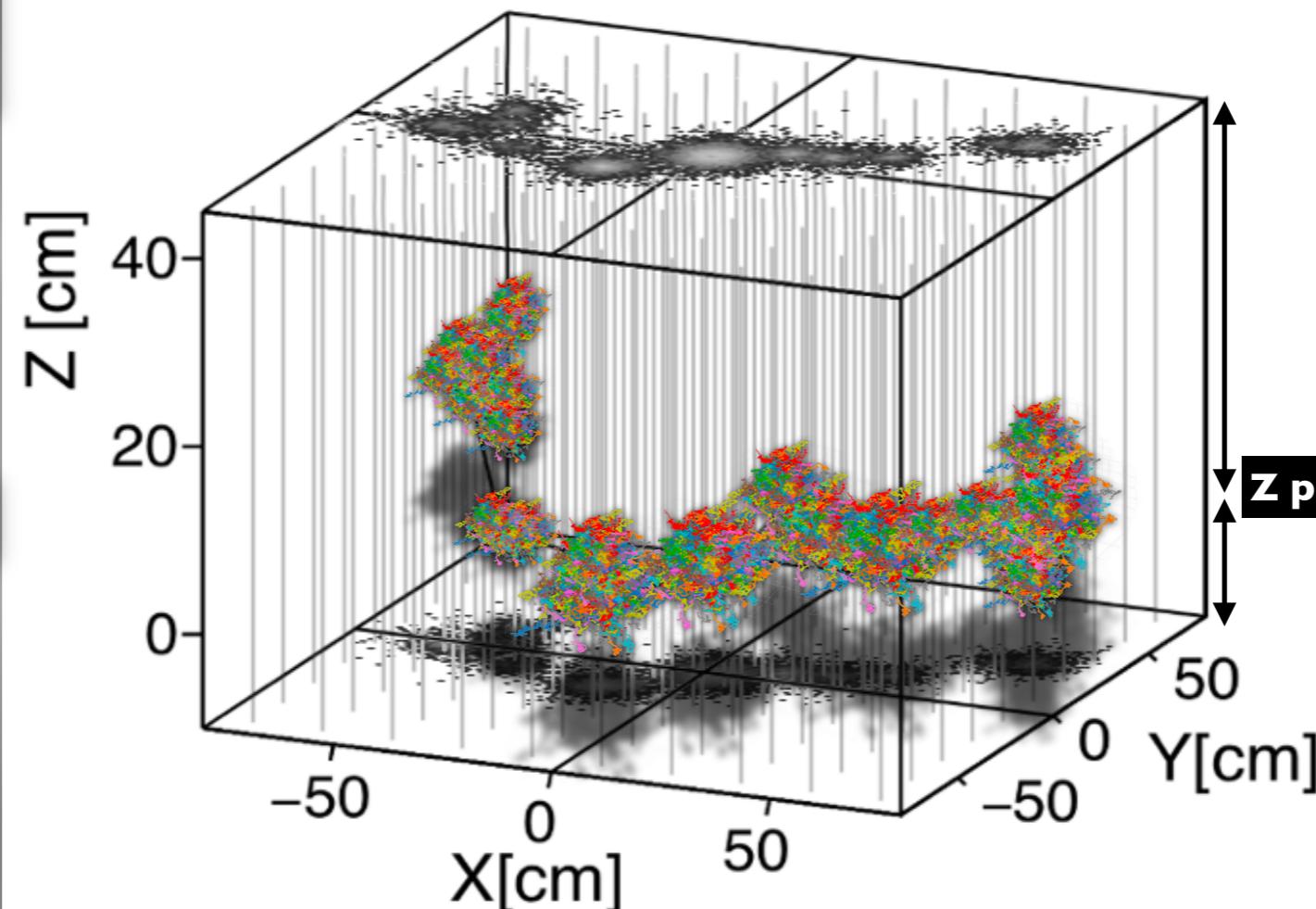
Topology (X,Y) direct & native (PID) \rightarrow possible **sub-mm vertex precision**

$\sim 1.0\text{MeV}$

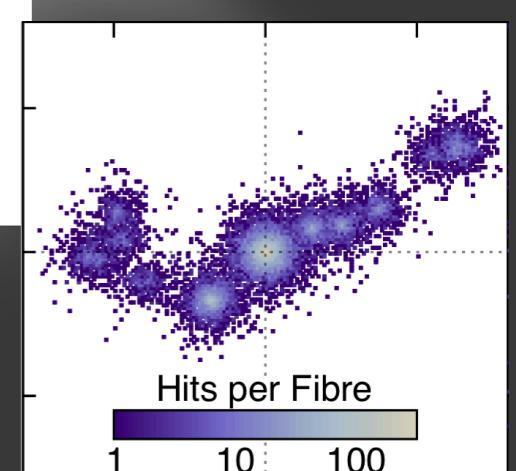
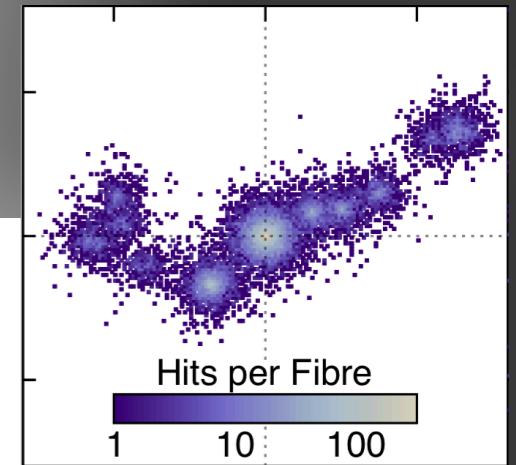


Vanilla LiquidO: 1D lattice (fibres along Z-axis only)

TOP VIEW: (X,Y) Projection \rightarrow direct readout



BOTTOM VIEW: (X,Y) Projection \rightarrow direct readout



LiquidO can have up 3 orthogonal fibre lattice orientations (3D)

latest results — March 2025

The Stochastic Light Confinement of LiquidO

LiquidO Collaboration^{(a-z)*}

Publication “LiquidO-I”

J. Apilluelo^{l,z}, L. Asquith^b, E. F. Bannister^b, N. P. Barradas^{k α} , J. L. Beney^p, M. Berberan e Santos^{k β} , X. de la Bernardie^p, T. J. C. Bezerra  ^b, M. Bongrand^p, C. Bourgeois^q, D. Breton^q, C. Buck  ¹, J. Bustosⁿ, K. Burns^q, A. Cabrera  ^{q,c,2}, A. Cadiou^p, E. Calvo^l, E. Chauveau^f, B. J. Cattermole^b, M. Chen^h, P. Chimentiⁱ, D. F. Cowen^{x α ,x β} , S. Dusini  ^{r α} , A. Earle^b, M. Felizardo^{k α} , C. Frigerio Martinsⁱ, J. Galán^z, J. A. García^z, R. Gazzini^q, A. Gibson-Foster^b, C. Girard-Carillo^{m α} , B. Gramlich¹, M. Grassi^{2,r β} , W. C. Griffith^b, J. J. Gómez-Cadenas^u, M. Guitière^p, F. Haddad^p, J. Hartnell^b, A. Holin^d, I. G. Irastorza^z, I. Jovanovic  ^a, A. Kling^{k α} , L. Koch  ^{m α} , P. Lasorak^b, J. F. Le Du^{q,c}, C. Lefebvre^h, F. Lefevre^p, P. Loaiza^q, J. A. Lock^b, G. Luzón^z, J. Maalmi^q, J. P. Malhado^j, F. Mantovani^{e α ,e β} , J. G. Marques^{k α} , C. Marquet^f, M. Martínez^z, D. Navas-Nicolás  ^{q,l}, H. Nunokawa^t, M. Obolensky², J. P. Ochoa-Ricoux  ^g, T. Palmeira^{k β} , C. Palomares^l, B. Pedras^{k β} , D. Petyt^d, P. Pillot^p, A. Pin^f, J. C. C. Porter^b, M. S. Pravikoff  ^f, N. Rodrigues^{k β} , M. Roche^f, R. Rosero^y, B. Roskovec^s, N. Roy^q, M. L. Sarsa^z, S. Schoppmann  ^{m $\beta,1$} , A. Serafini^{r $\alpha,r\beta$} , C. Shepherd-Themistocleous^d, W. Shorrock  ^b, M. Silva^{k γ} , L. Simard^q, S. R. Soleti^u, H. Th. J. Steiger^{m $\alpha,m\beta$} , D. Stocco^p, V. Strati^{e $\alpha,e\beta$} , J. S. Stutzmann^p, F. Suekane^v, A. Tunc^{m α} , N. Tuccori  ^b, A. Verdugo^l, B. Viaud^p, S. M. Wakely  ^{m α} , A. Weber  ^{m α} , G. Wendel^{x β} , A. S. Wilhelm  ^a, M. Yeh^y, and F. Yermia^p

[arXiv:2503.02541](https://arxiv.org/abs/2503.02541) — under publication

latest results — July 2025

Muon tracking in a LiquidO opaque scintillator detector

Publication “LiquidO-II”

LiquidO Collaboration

J. Apilluelo,^z L. Asquith,^b E. F. Bannister,^b N. P. Barradas,^{k α} C. L. Baylis,^b J. L. Beney,^p
M. Berberan e Santos,^{k β} X. de la Bernardie,^p T. J. C. Bezerra ,^b M. Bongrand,^p
C. Bourgeois,^q D. Breton,^q J. Bustos,ⁿ A. Cabrera ,^{q,c} A. Cadiou,^p E. Calvo,^l
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C. Frigerio Martins,ⁱ J. Galán,^z J. A. García,^z A. Gibson-Foster,^b C. Girard-Carillo,^{m α}
W. C. Griffith,^b J. J. Gómez-Cadenas,^u M. Guitière,^p F. Haddad,^p J. Hartnell ,^b A. Holin,^d
I. G. Irastorza,^z I. Jovanovic ,^a A. Kling,^{k α} L. Koch ,^{m α} P. Lasorak,^b J. F. Le Du,^{q,c}
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A. Verdugo,^l B. Viaud,^p S. M. Wakely ,^{m α} A. Weber ,^{m α} G. Wendel ,^{x β} A. S. Wilhelm
,^a A. W. R. Wong,^{b,d} M. Yeh,^y F. Yermia^p

[arXiv:2507.13864](https://arxiv.org/abs/2507.13864) — under publication

IFAE just joined...

LiquidO Collaboration^{(a–z)*}

J. Apilluelo^z, L. Asquith^b, E. F. Bannister^b, N. P. Barradas^{kα}, C. L. Baylis^b, J. L. Beney^p, M. Berberan e Santos^{kβ}, X. de la Bernardie^p, T. J. C. Bezerra^b, M. Bongrand^p, C. Bourgeois^q, D. Breton^q, J. Bustosⁿ, A. Cabrera^{q,c}, A. Cadiou^p, E. Calvo^l, M. de Carlos Generowicz^b, E. Chauveau^f, B. J. Cattermole^b, M. Chen^h, P. Chimentiⁱ, D. F. Cowen^{xα,xβ}, S. Kr. Das^b, S. Dusini^{rα}, A. Earle^b, M. Felizardo^{kα}, C. Frigerio Martinsⁱ, J. Galán^z, J. A. García^z, A. Gibson-Foster^b, C. Girard-Carillo^{mα}, W. C. Griffith^b, J. J. Gómez-Cadenas^u, M. Guitière^p, F. Haddad^p, J. Hartnell^b, A. Holin^d, I. G. Irastorza^z, I. Jovanovic^a, A. Kling^{kα}, L. Koch^{mα}, P. Lasorak^b, J. F. Le Du^{q,c}, F. Lefevre^p, P. Loaiza^q, J. A. Lock^b, G. Luzón^z, J. Maalmi^q, J. P. Malhado^j, F. Mantovani^{eα,eβ}, J. G. Marques^{kα}, C. Marquet^f, M. Martínez^z, J. T. Moffat^{xβ}, D. Navas-Nicolás^l, H. Nunokawa^t, J. P. Ochoa-Ricoux^g, T. Palmeira^{kβ}, C. Palomares^l, D. Petyt^d, P. Pillot^p, A. Pin^f, J. C. C. Porter^b, M. S. Pravikoff^f, S. Richards^d, N. Rodrigues^{kβ}, M. Roche^f, R. Rosero^y, B. Roskovec^s, M. L. Sarsa^z, A. Serafini^{rα,rβ}, C. Shepherd-Themistocleous^d, W. Shorrock^b, M. Silva^{kγ}, L. Simard^q, M. Sistio^o, S. R. Soleti^u, D. Stocco^p, V. Strati^{eα,eβ}, J. S. Stutzmann^p, F. Suekane^v, N. Tuccori^b, A. Verdugo^l, B. Viaud^p, S. M. Wakely^{mα}, A. Weber^{mα}, G. Wendel^{xβ}, A. S. Wilhelm^a, A. W. R. Wong^b, M. Yeh^y, and F. Yermia^p

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^yBrookhaven National Laboratory, Upton, NY, USA

^zCentro de Astropartículas y Física de Altas Energías (CAPA), Universidad de Zaragoza, Zaragoza, Spain

Spokespersons:

- A. Cabrera — IJCLab / Université Paris-Saclay (France)
- F. Suekane — Tohoku University / RCNS (Japan)

Web: <https://liquid.ijclab.in2p3.fr/>

Chooz (most powerful reactor) \oplus **UND** ($\geq 25\text{m}$ baseline) \oplus **LiquidO** (BG rejection)
[**EDF** within the team — unprecedented]

C L U D

I - II - III

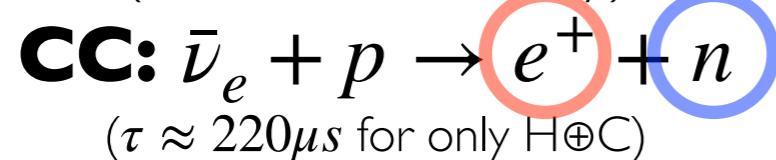
CLOUD's sequence...

the power of coincidences

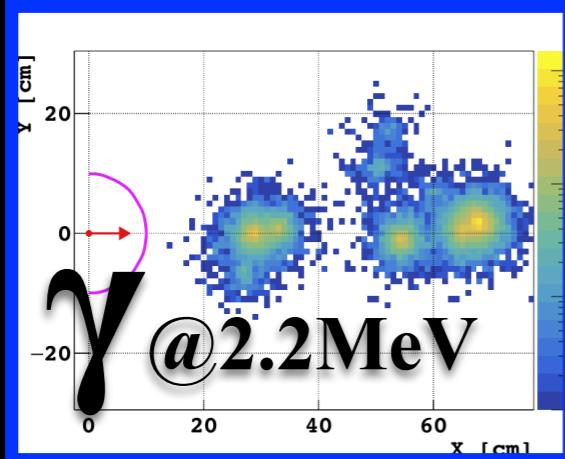
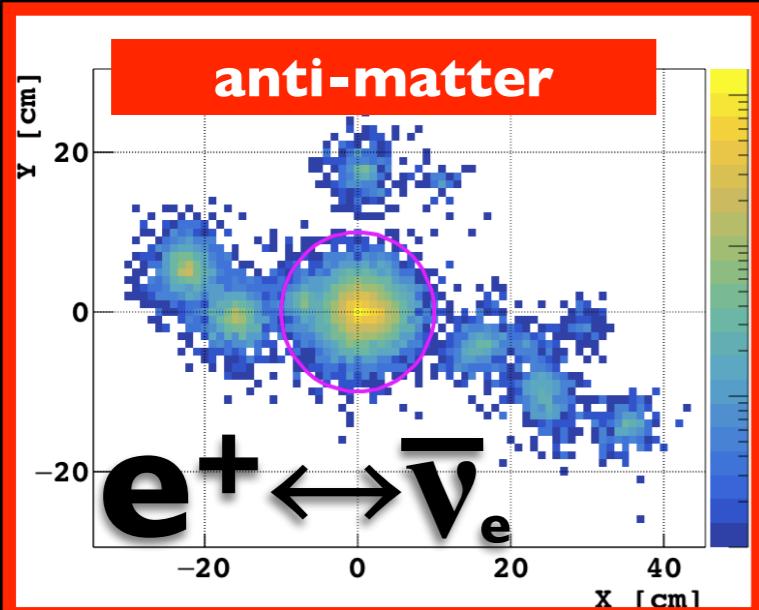
being at **the right “place & time & energy & PID” — huge rejection(s)**

Reines et al ‘50s

(neutrino discovery)



(anti)neutrino **discovery** [τ_n & $\Delta m_{p \sim n}$]

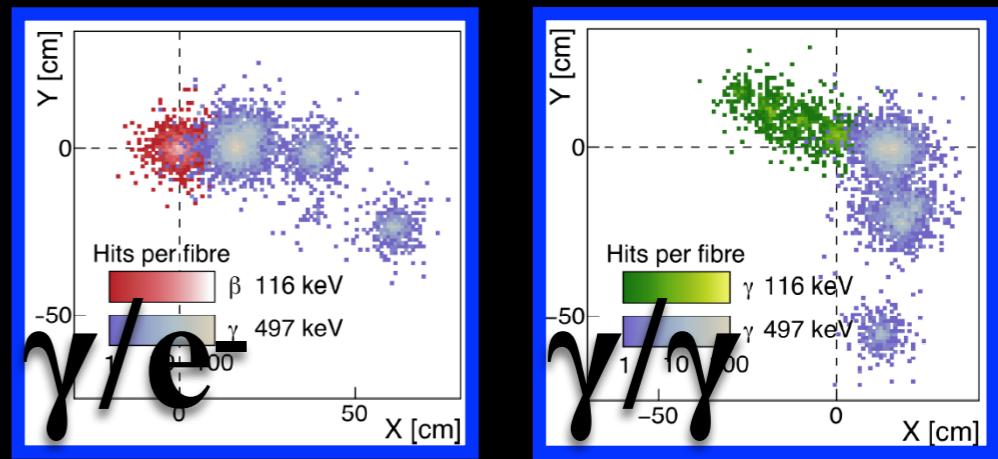
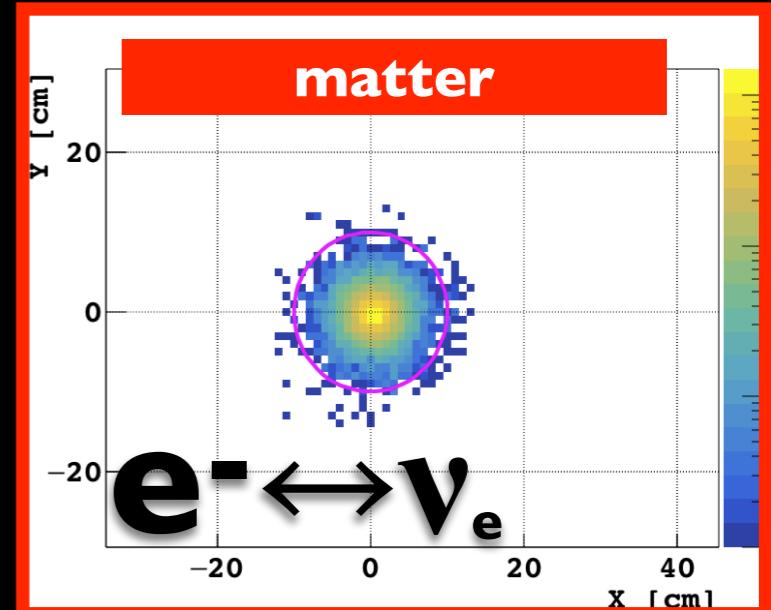


Raghavan et al ‘70s

(pp solar neutrino — unobserved)



major **R&D** [~ 2 decades] by **LENS** et al.



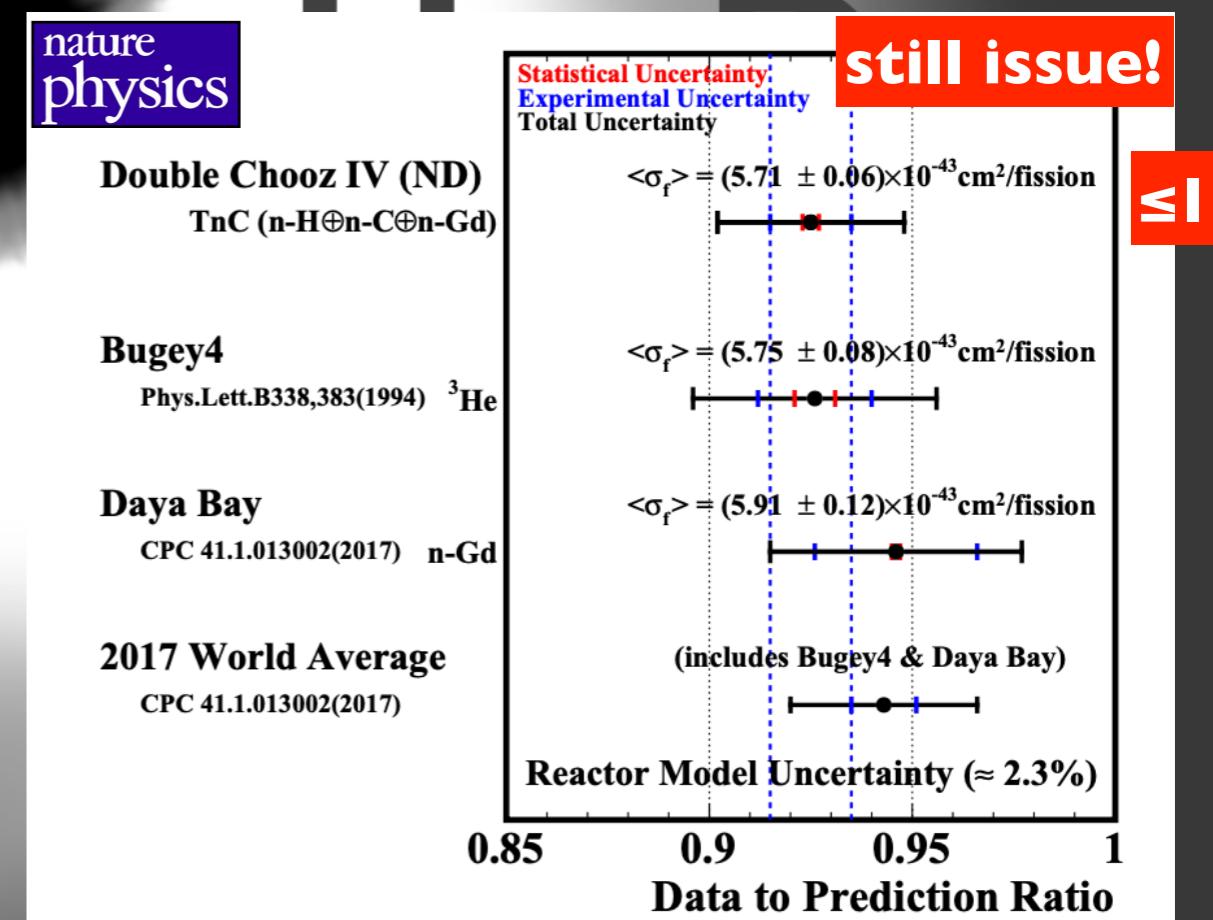
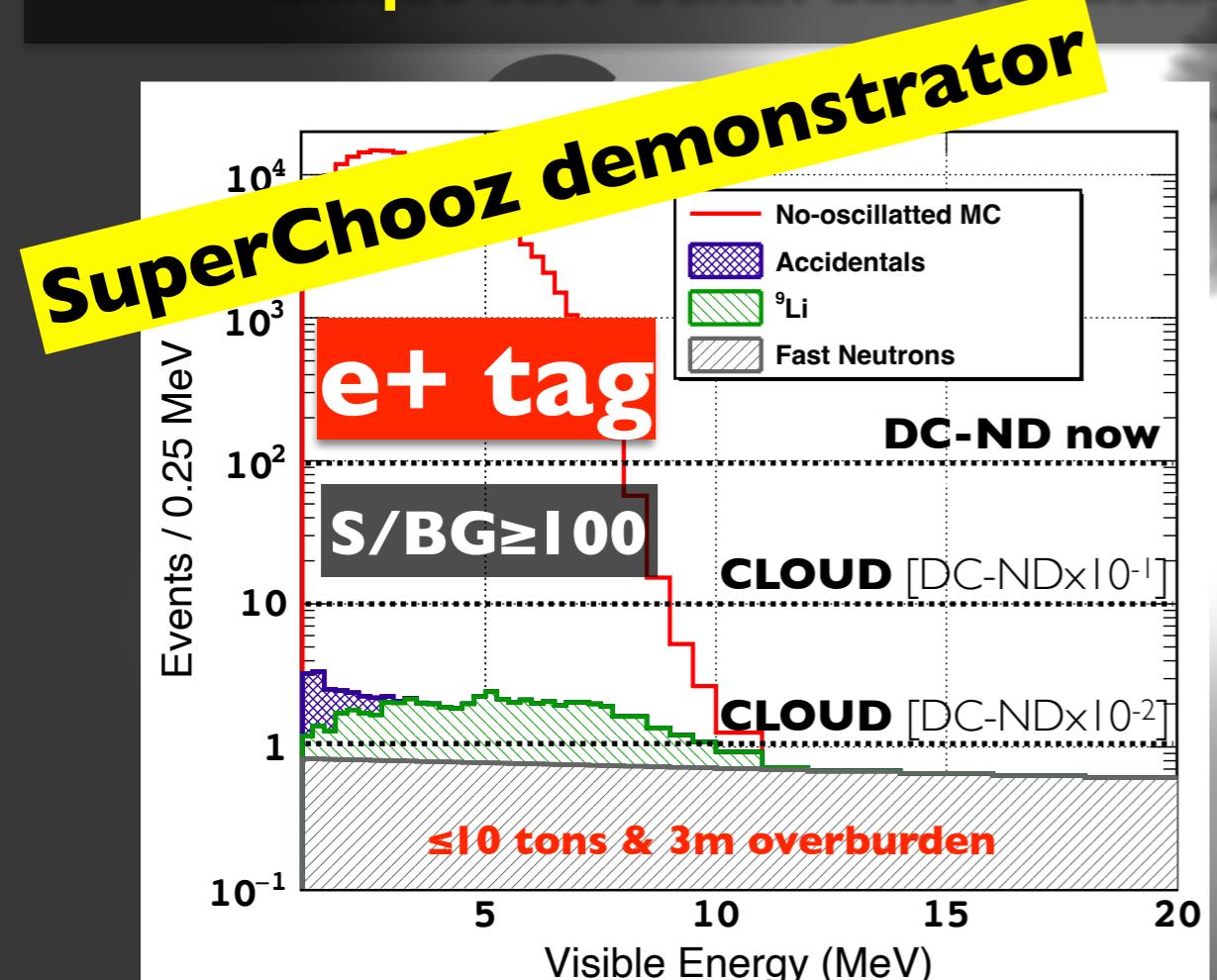
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AntiMatter-OTech — synergy
SuperChooz's antineutrino golden channel **demonstration** — byproduct

antineutrino CC & NC? (doping) . . .

CLOUD-I physics programme: IBD@p...

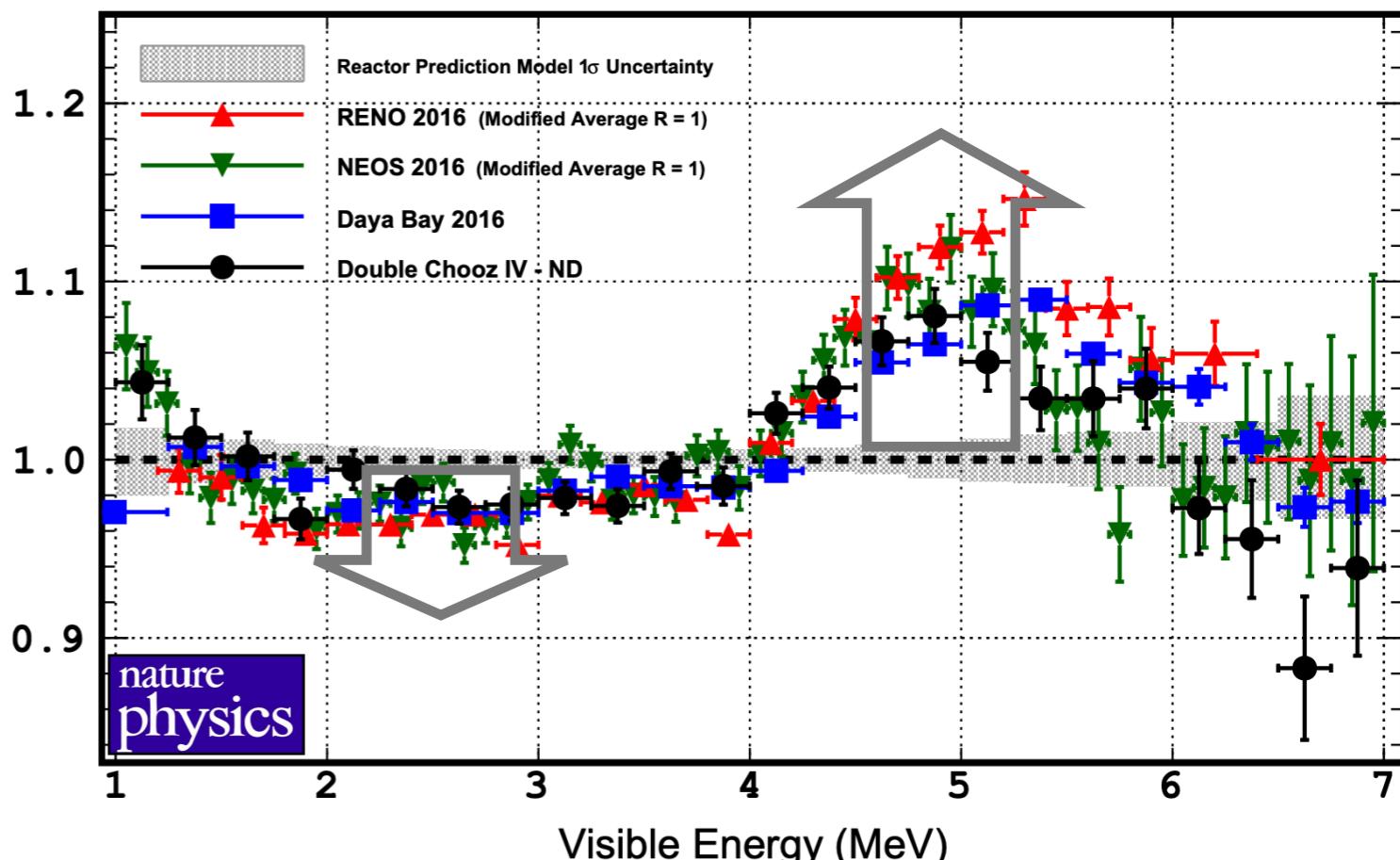
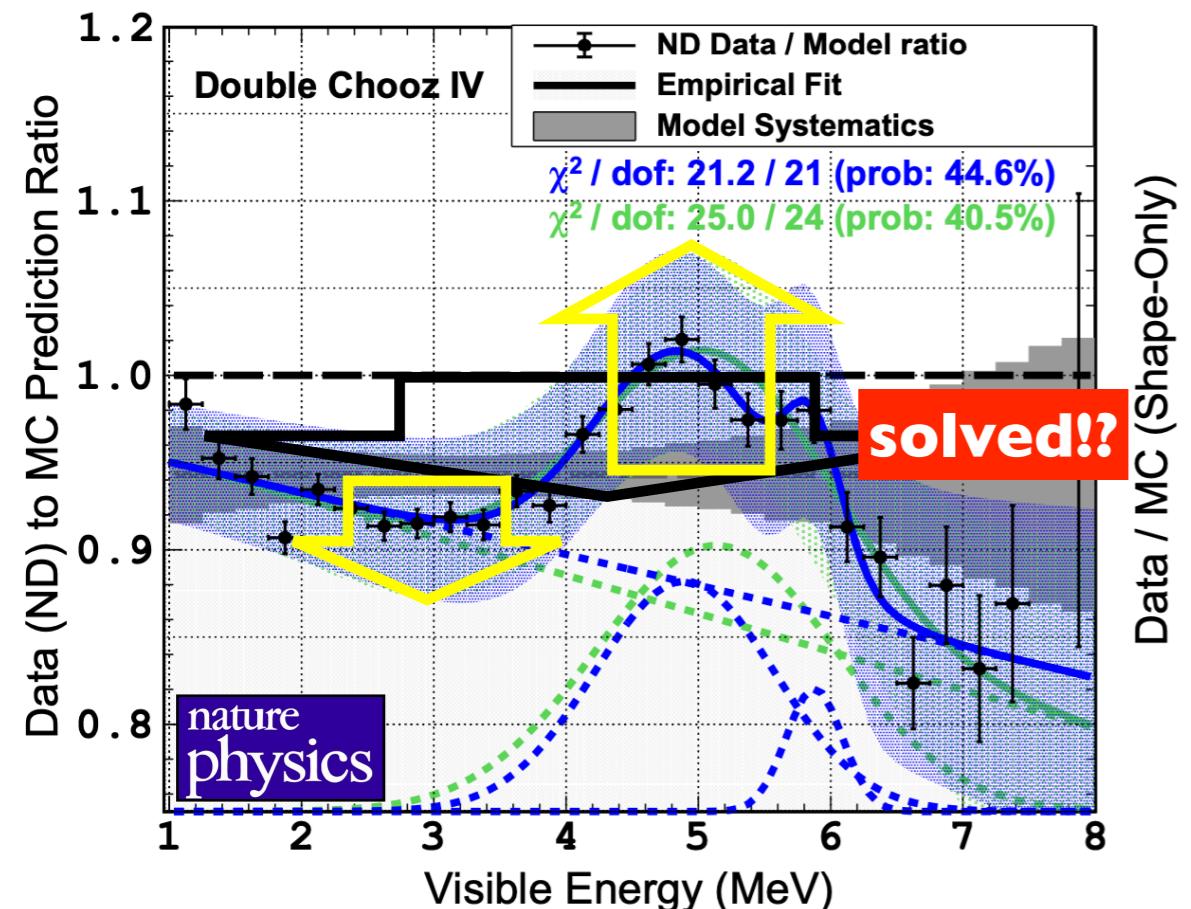
- IBD@p (anti-v CC): **≥10,000** interactions per day in **10tons** [**≥3M** interactions per year]
 - **LiquidO** reach a **background-less** regime — improve **≥3x** today's BG control (ex. DC-ND)
 - **Signal(ON)-to-BG ≥100** — unprecedented high precision reactor characterisation
 - dominant **~0.5% (thermal power) uncertainty** & accurate **U/Pu composition**
 - **Signal(OFF)-to-BG ≥1** — unprecedented **reactor-fuel monitoring**
 - accurate monitoring of **transitions OFF-ON-OFF** — interesting physics too
 - **unique test-bench data for accurate prediction** — validate uncertainties, too?



CLOUD precision $\geq 0.6\%$ ⇒ Unitarity Violation? (if predictions are improved)!

still issues to be resolved — beyond the sterile neutrino hypothesis (disfavoured)

all experiments consistent — except **Bugey3?!**



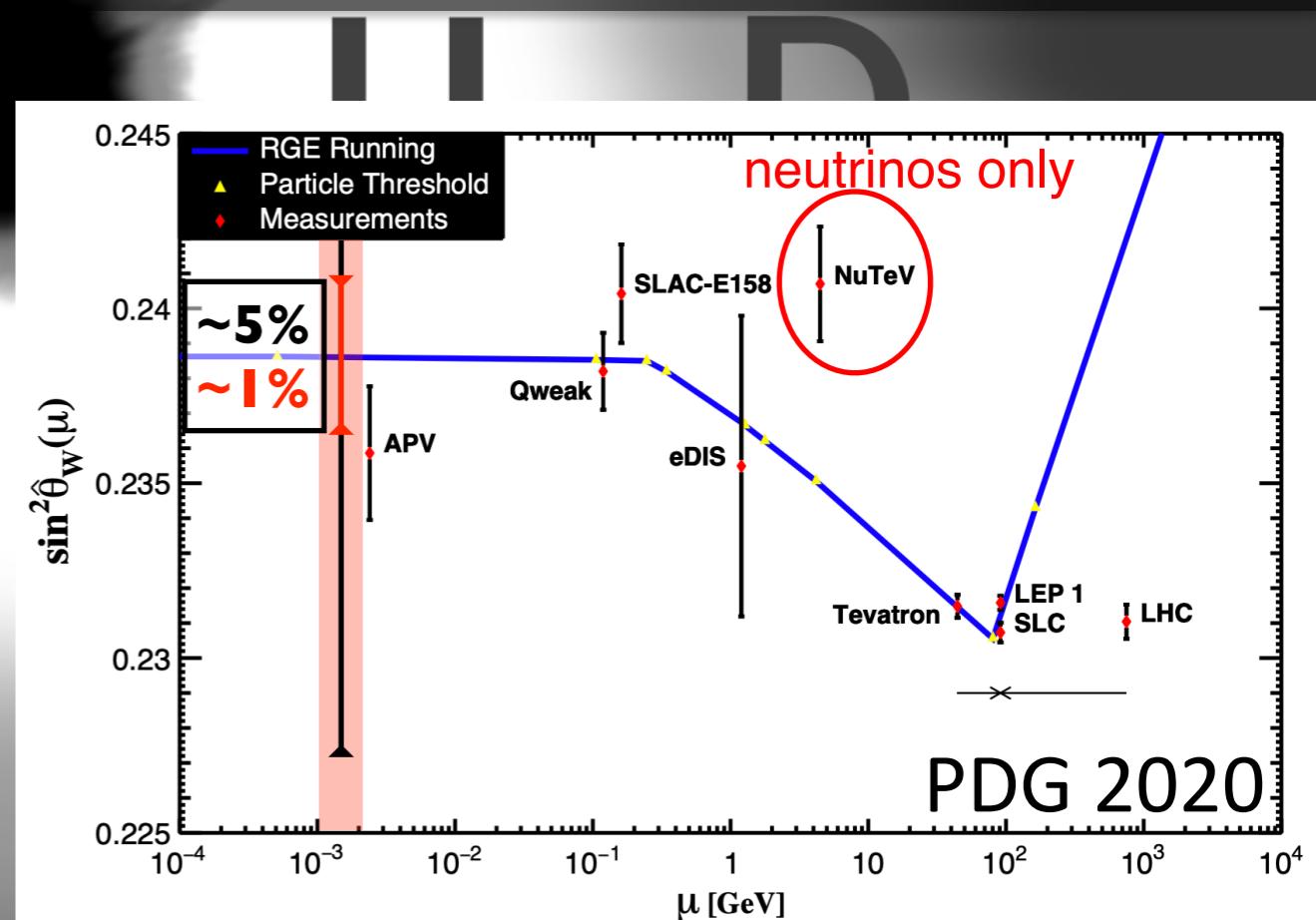
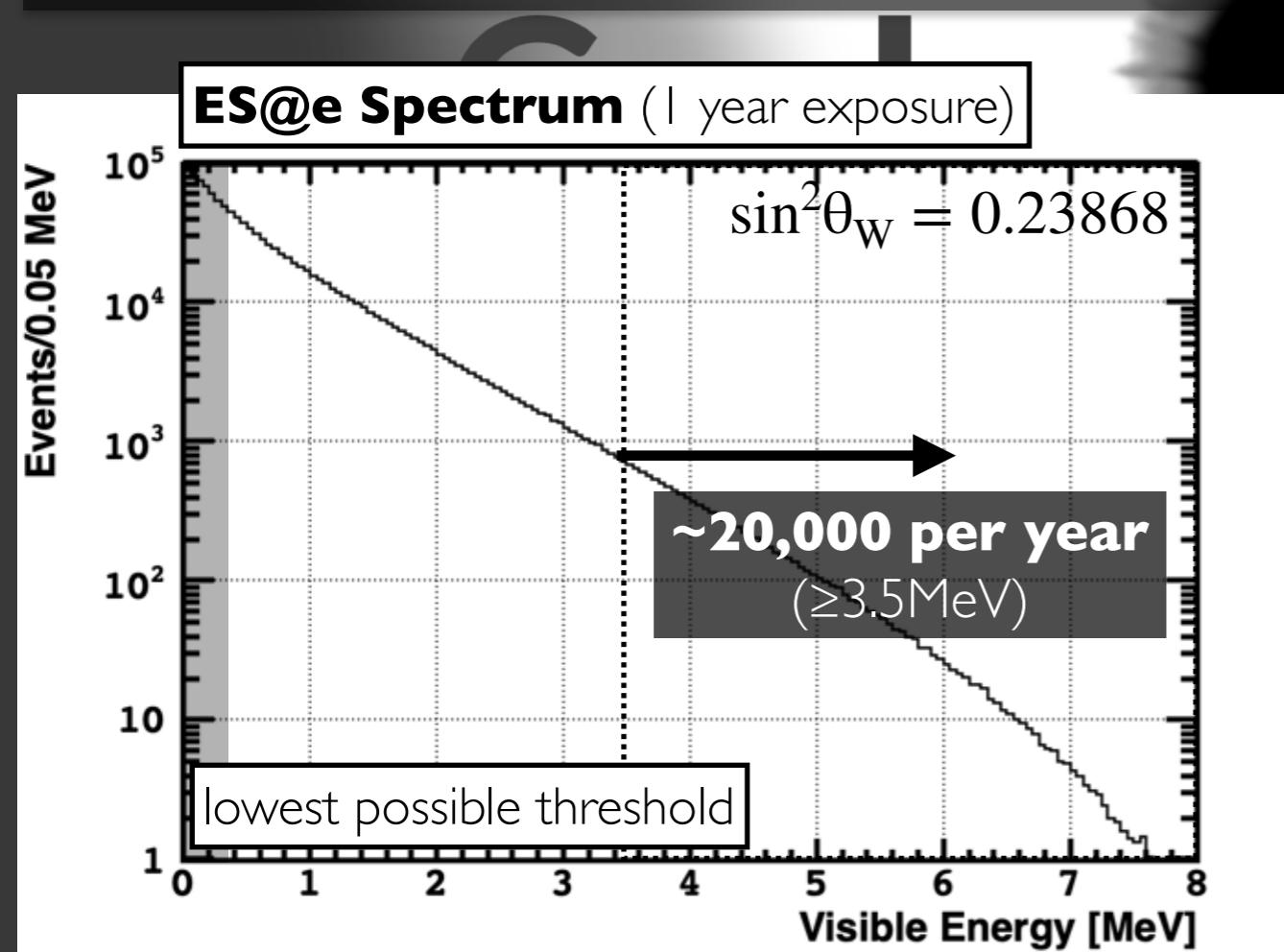
ABSOLUTE FLUX: the future of reactor-antineutrino physics

we must solve this “mess” \Rightarrow the reward **possible new physics!**
(if so, prediction should not use neutrino input \Rightarrow **no** new physics)

must understand flux $\leq 1\%$...

CLOUD-I physics programme: ES@e...

- **eES** (anti-v CC+NC): **≤5,000** interactions per day for **10tons ID** [$\leq 2\text{M}$ interactions per year]
 - interference CC & NC — different for neutrino (easier) and antineutrino (harder)
 - measure θ_w or use to **decompose the NC flux** component
 - PDG-2022's $\sin^2\theta_w \approx [0.231, 0.239]$ — running due to SM's renormalisation
- **major challenge: LiquidO** isolate “e-like” PID and exploit **high-rate reactor modulation**
 - likely strong **fiducial volume & higher energies** — reduce detected rate drastically
 - **≤10% precision** ($\geq 5\sigma$ observation) tolerates much BG but $\leq 1\% \Rightarrow S/BG \geq 2(!!)$ **impossible?**



R.L. Workman et al. (Particle Data Group), Prog. Theor. Exp. Phys. 2022, 083C01 (2022)

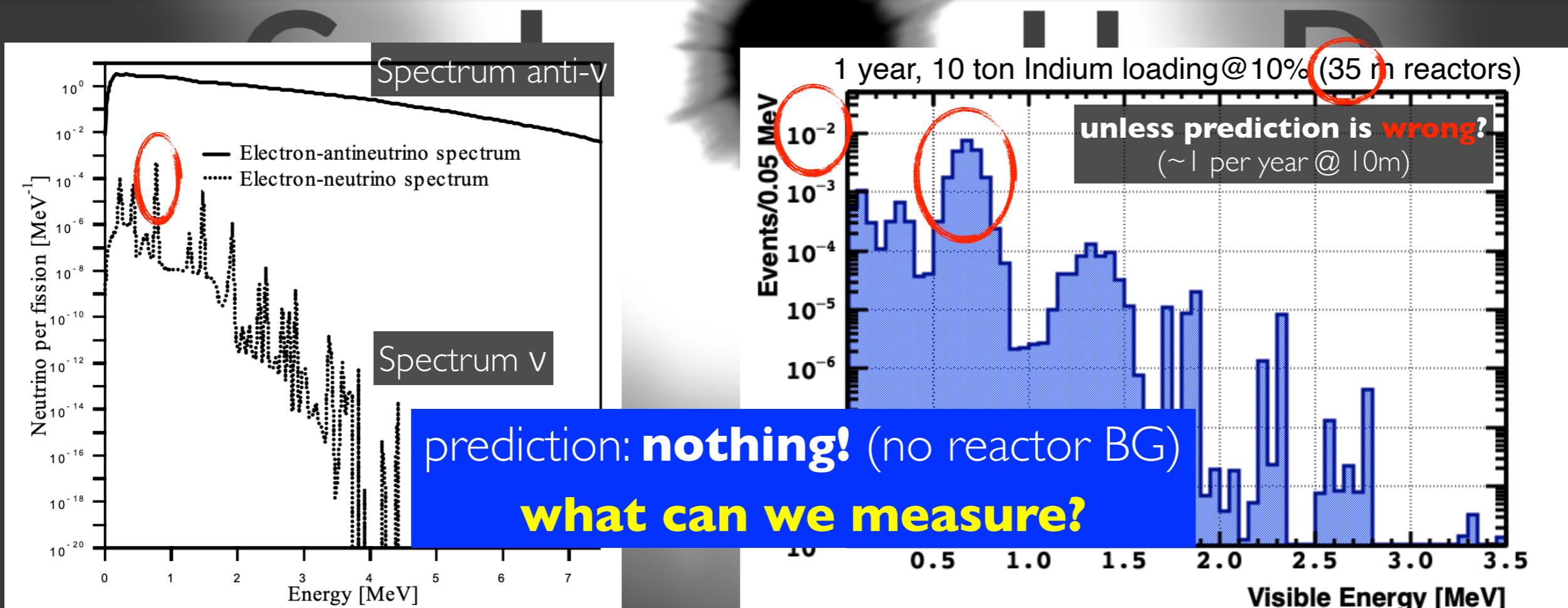
C L II U D

SuperChooz's neutrino golden channel **demonstration — byproduct**

neutrino CC (doping)...

CLOUD-II physics programme: neutrino...

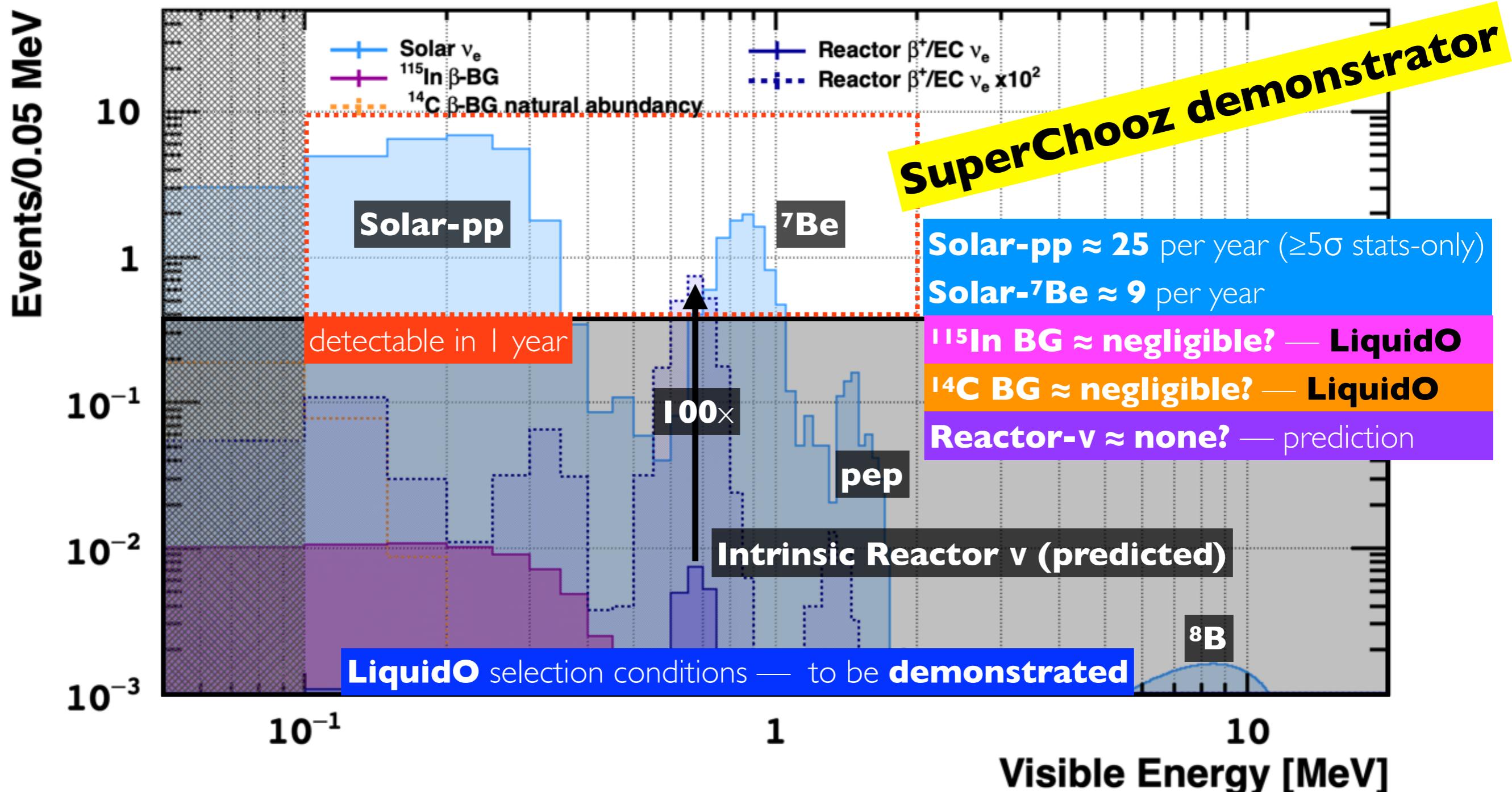
- loading **indium** on the detector — **unique strong coincidence $\geq 114\text{keV}$** (“solar-pp” in mind)
 - low threshold \oplus high natural-abundance \oplus high-ish cross-section \oplus BG-killer (coincidence)
 - CC interaction: $\nu_e + ^{115}\text{In} \rightarrow e^- + ^{115}\text{Sn}^*$ [$\tau:4.8\mu\text{s}$ decay: $\gamma/e(116\text{keV}) + \gamma(496\text{keV})$]
 - reactor neutrino **modulate with the reactor power** — no ambiguity whatsoever
- why to detect neutrinos close to a reactor?
 - **reactor neutrinos** (from β^+/EC): **rate(ν) $\approx 10^{-5}$ rate(anti- ν)** — prediction (both correlated)
 - could **reactors be the missing MeV neutrino source?** [otherwise impractical]



the big picture of neutrinos @CLOUD...

assuming the **LENS “BG model”** — valid at **overburden $\sim 3\text{m}$?** (to be demonstrated)

1 year, 10 ton Indium loading@10% (35 m reactors)



detection **solar-pp neutrinos** on a **10ton** detector **almost on surface** right **next a nuclear reactor?**

C L I P U D

R&D for low energy reactor-fuel monitoring & geoneutrino ^{40}K discovery — demonstration

new antineutrino CC (doping)...

Probing Earth's Missing Potassium using the Unique Antimatter Signature of Geoneutrinos

arXiv:2308.04154

A. Cabrera^{*12 α ,2,a}, M. Chen^{†6}, F. Mantovani^{‡3 α ,3 β} , A. Serafini^{§3 α ,3 β ,13 α ,13 β} , V. Strati^{¶3 α ,3 β} , J. Apilluelo¹⁸, L. Asquith¹, J.L. Beney¹¹, T.J.C. Bezerra¹, M. Bongrand¹¹, C. Bourgeois^{12 α} , D. Breton^{12 α} , M. Briere^{12 α} , J. Bustos¹⁰, A. Cadiou¹¹, E. Calvo⁸, V. Chaumat^{12 α} , E. Chauveau⁴, B.J. Cattermole¹, P. Chimenti⁷, C. Delafosse^{12 α} , H. de Kerret^{¶a}, S. Dusini^{13 α} , A. Earle¹, C. Frigerio-Martins⁷, J. Galán¹⁸, J. A. García¹⁸, R. Gazzini^{12 α} , A. Gibson-Foster¹, A. Gallas^{12 α} , C. Girard-Carillo^{9 α} , W.C. Griffith¹, F. Haddad¹¹, J. Hartnell¹, A. Hourlier¹⁷, G. Hull^{12 α} , I. G. Irastorza¹⁸, L. Koch^{9 α} , P. Laniéce^{12 α ,12 β} , J.F. Le Du^{12 α ,2}, C. Lefebvre⁶, F. Lefevre¹¹, F. Legrand^{12 α} , P. Loaiza^{12 α} , J. A. Lock¹, G. Luzón¹⁸, J. Maalmi^{12 α} , C. Marquet⁴, M. Martínez¹⁸, B. Mathon^{12 α} , L. Ménard^{12 α ,12 β} , D. Navas-Nicolás^{12 α} , H. Nunokawa¹⁵, J.P. Ochoa-Ricoux⁵, M. Obolensky^a, C. Palomares⁸, P. Pillot¹¹, J.C.C. Porter¹, M.S. Pravikoff⁴, H. Ramarijaona^{12 α} , M. Roche⁴, P. Rosier^{12 α} , B. Roskovec¹⁴, M.L. Sarsa¹⁸, S. Schoppmann^{9 β} , W. Shorrock¹, L. Simard^{12 α} , H.Th.J. Steiger^{9 α ,9 β} , D. Stocco¹¹, J.S. Stutzmann¹¹, F. Suekane^{16,a}, A. Tunc^{9 α} , M.-A. Verdier^{12 α ,12 β} , A. Verdugo⁸, B. Viaud¹¹, S. M. Wakely^{9 α} , A. Weber^{9 α} , and F. Yermia¹¹

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^a Université de Paris Cité, CNRS, APC, Paris, France

(LiquidO Consortium)

4⁰K geoneutrino new methodology → good enough for discovery?

C L U D

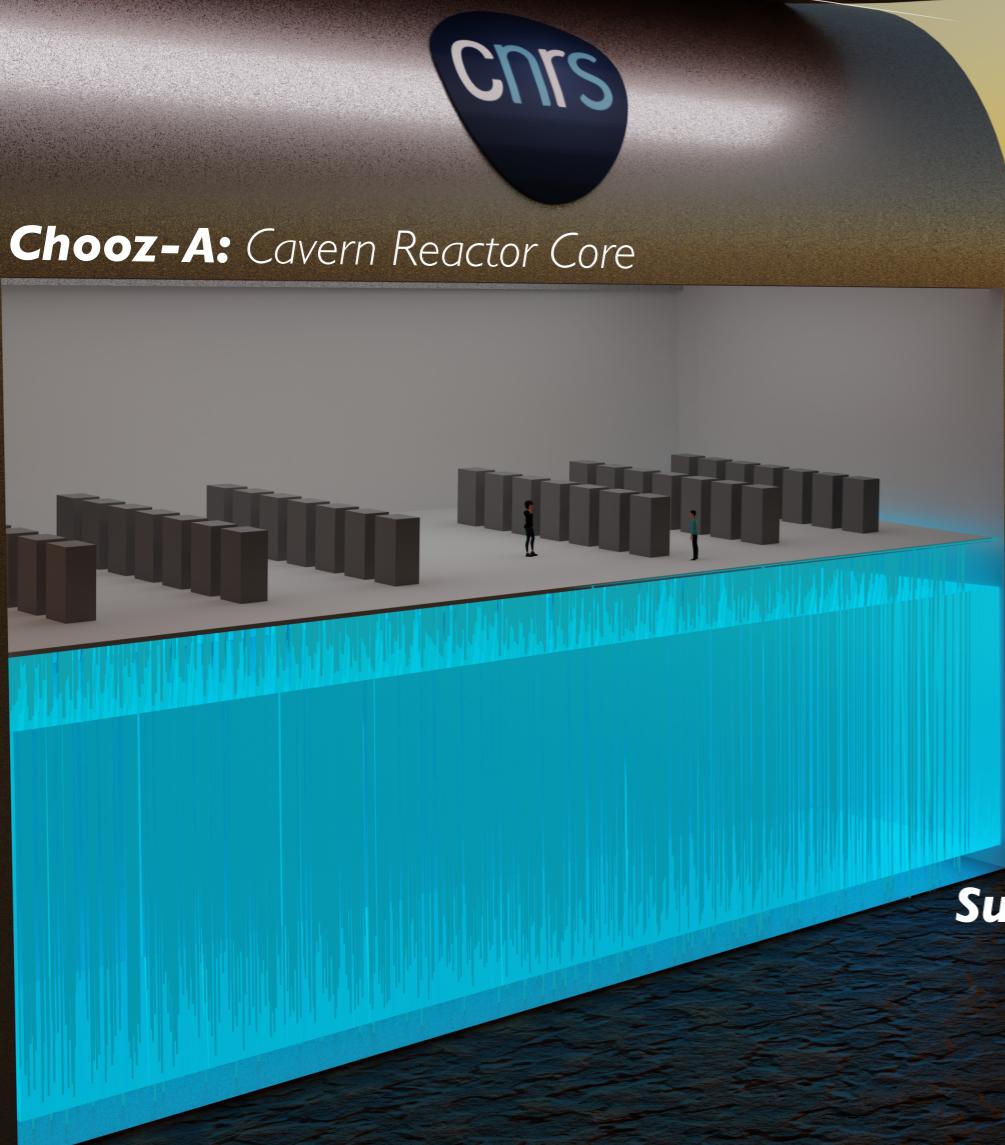


a long story short...

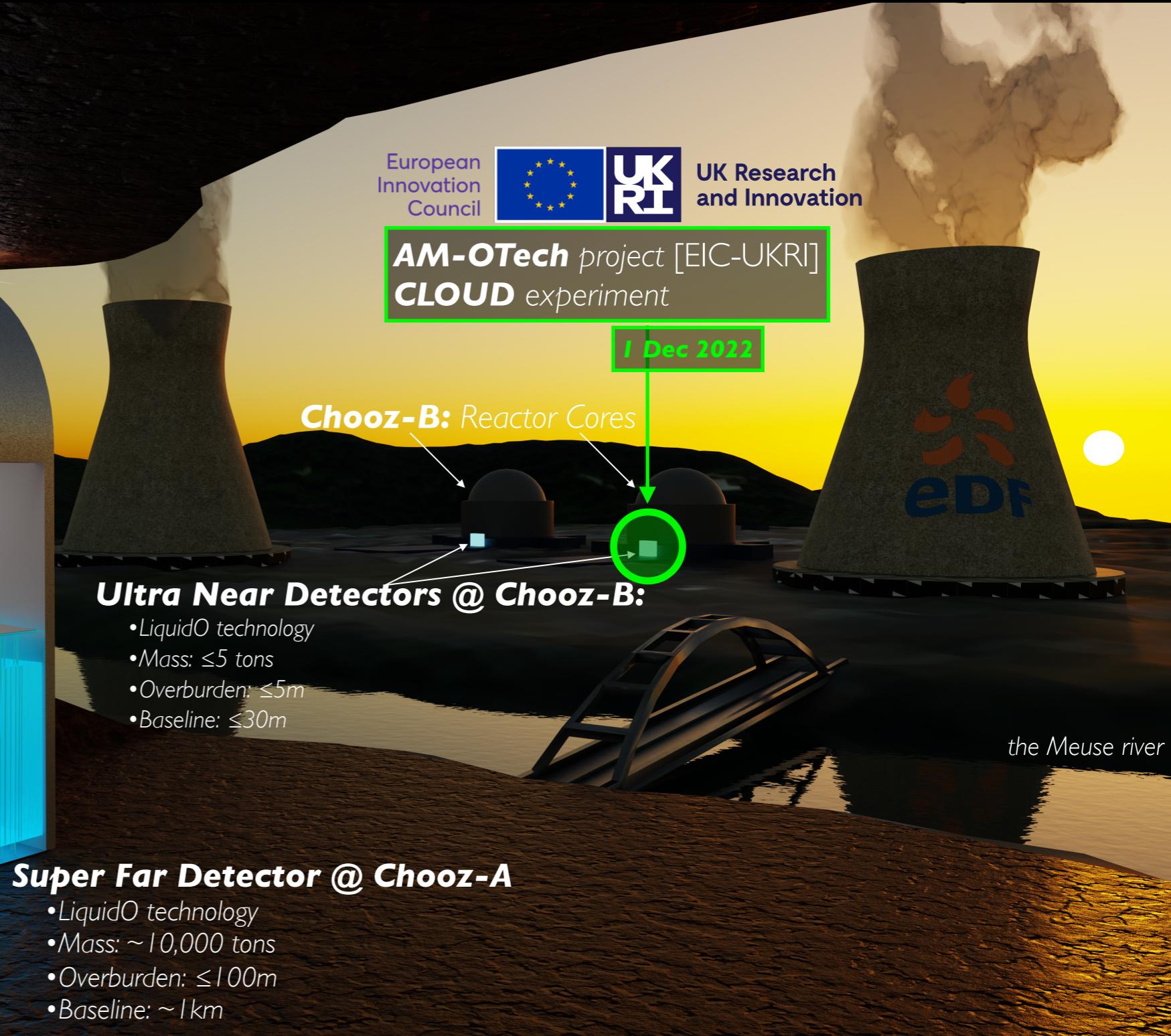
SuperChooz exploration...

flagship neutrino oscillation experiment in Europe?

the Ardennes mountains



Chooz-A: Cavern Reactor Core



- **CLOUD** demonstrator for **LiquidO's (anti)neutrino detection capabilities — a revolution?**
 - byproduct to **new reactor monitoring capability?** [a dream since '70s]
- **CLOUD-I: approved & funded [AM-OTech] plan: data by 2025**
 - **goal:** possible **background-less regime?** [→ LiquidO]
 - **most precise absolute CC-antineutrino flux — new physics?**
 - possible **first NC-(anti)neutrino flux — new physics?**
- **CLOUD-II: under feasibility study** (→ new **indium**-loaded opaque scintillator)
 - (first) **absolute CC-neutrino flux reactor — new physics?**
 - **measure solar-pp ($\geq 5\sigma$)** in a “tiny detector” almost on the surface? ⇒ **a breakthrough?**
- **CLOUD-III: under feasibility study** (→ new **copper**-loaded opaque scintillator)
 - probe **reactor flux at low energies?** — **surprises?** [first time ever below 1.8MeV]
 - demonstration for **^{40}K detection methodology** — **a discovery one day?**

an even vaster future of reactor (anti)neutrinos ahead?

conclusions...

SuperChooz demonstrator

our collaboration...

European
Innovation
Council



UK Research
and Innovation

C L U D

CLOUD International collaboration

- **EDF** (France) — **first time in neutrino science**
- **Brookhaven National Laboratory** (USA)
- **Charles University** (Czechia)
- **CIEMAT** (Spain)
- **IJCLab** / Université Paris-Saclay (France)
- **Imperial College London** (UK)
- **INFN-Padova** (Italy)
- **Instituto Superior Técnico** (Portugal)
- **Johannes Gutenberg Universität Mainz** (Germany)
- **Pennsylvania State University** (USA)
- **Pontifícia Universidade Católica do Rio de Janeiro** (Brazil)
- **Queen's University** (Canada)
- **Subatech / Nantes Université** (France)
- **Tohoku University / RCNS** (Japan)
- **Universidad de Zaragoza** (Spain)
- **Universidade Estadual de Londrina** (Brazil)
- **University of California Irvine** (USA)
- **University of Michigan** (USA)
- **University of Sussex** (UK)

⇒ 19 institutions in 11 countries

Spokespersons:

- A. Cabrera — IJCLab / Université Paris-Saclay (France)
- J. Hartnell — Sussex University (UK)

IB Chair:

- M. Chen — Queen's University (Canada)

Webs:

- <https://antimatter-otech.ijclab.in2p3.fr/> [AMOTech]
- <https://liquido.ijclab.in2p3.fr/nucloud> [via LiquidO]

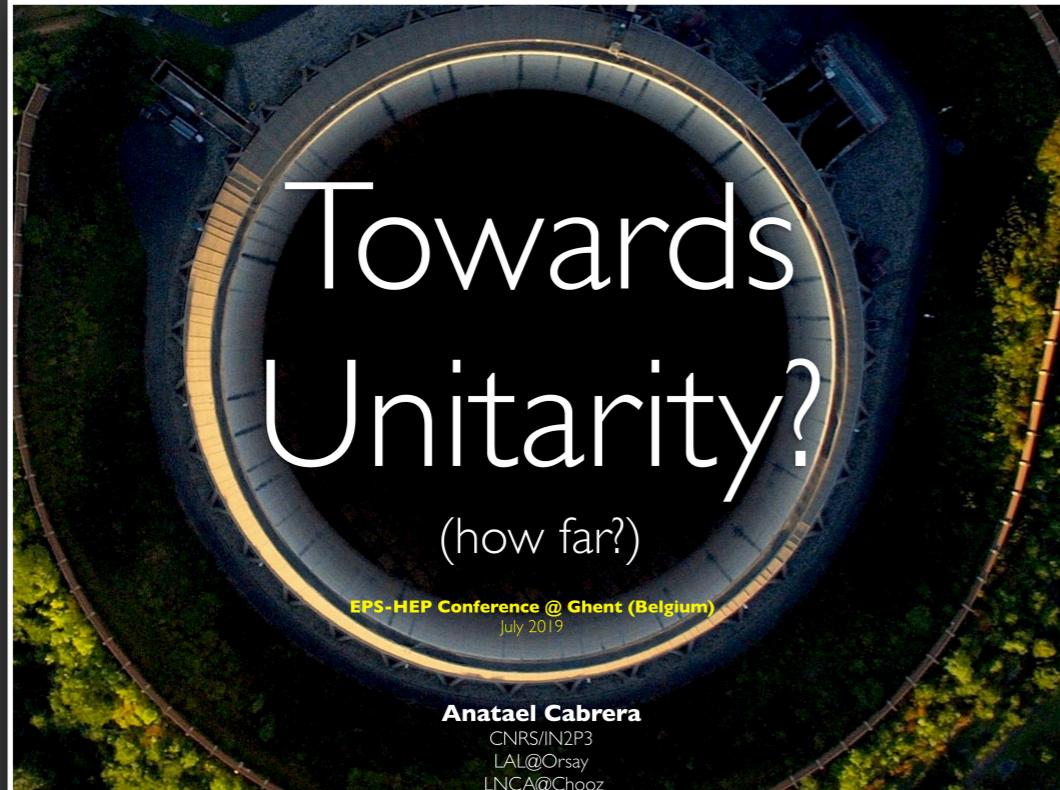
C L U D



backup slides...



HEP-European Physics Society
(July 2019 @ Ghent Belgium)



EP Seminar

The SuperChooz Experiment: Unveiling the Opportunity

by Dr Anatael CABRERA (IJCLab - IN2P3/CNRS)

Tuesday 29 Nov 2022, 11:00 → 12:00 Europe/Zurich

222/R-001 (CERN)



tightly linked to **LiquidO**, **AM-OTech/CLOUD**, and **SuperChooz** collaborations/consortia & specially **EDF**



<https://indico.cern.ch/event/577856/contributions/3421609/>

<https://indico.cern.ch/event/1215214/>

<https://zenodo.org/record/7504162>

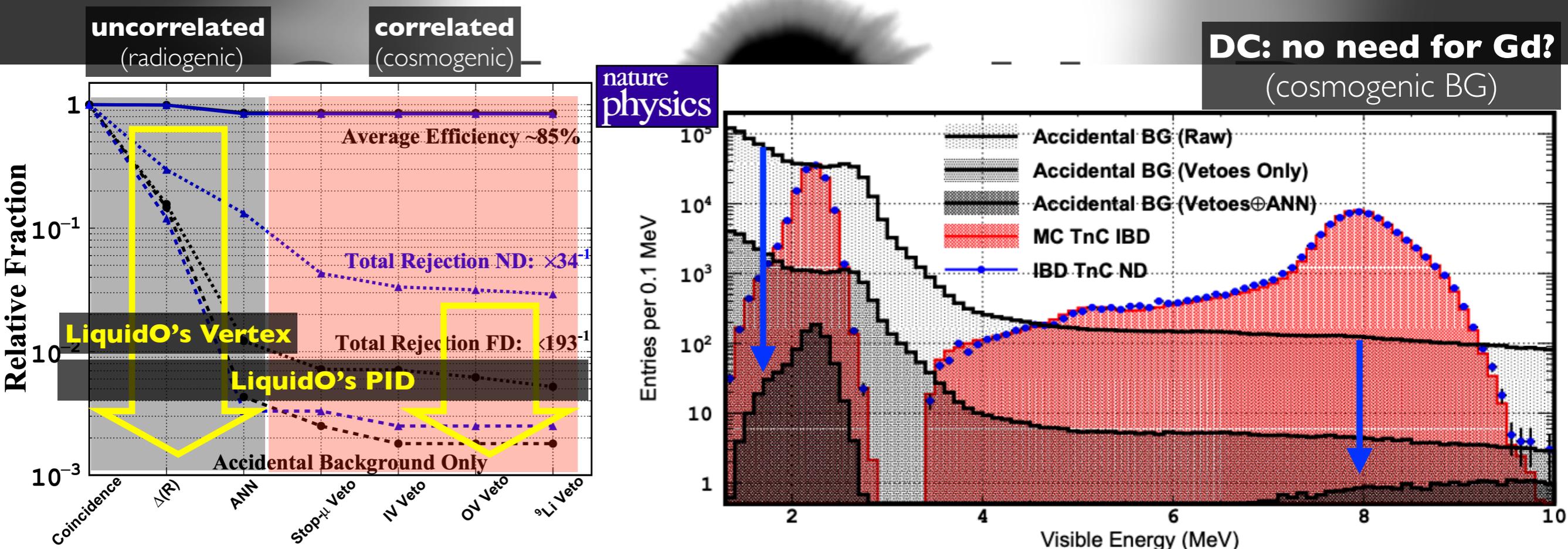
<https://liquido.ijclab.in2p3.fr/>

exploring since 2018...

active BG rejection and control...

- detection using **coincidence-signal** (ex. **IBD@p**) \Rightarrow prompt-delayed correlations
- **combinatory (uncorrelated) BG(s): 5D-coincidence** ($\Delta t \oplus \Delta r \oplus \Delta E$) — **LiquidO's** mm-vertex
- **cosmogenic (correlated) BG(s): particle-ID** — **LiquidO's** imaging [**impossible so far**]
- **active rejection** \rightarrow rejected-BG as **data-driven BG input** (high accuracy physics extraction)
- **radiogenic control**: in-situ radiogenic BG model tuning (radiopurity control order $\leq 10^{-14} g/g$)

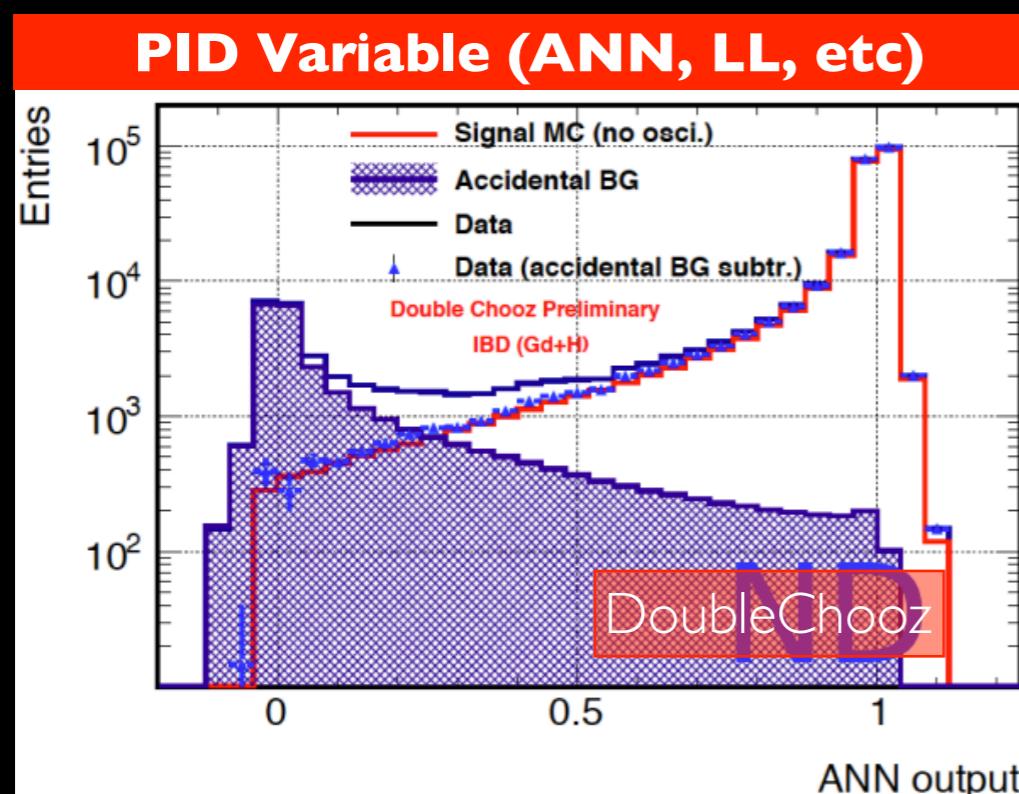
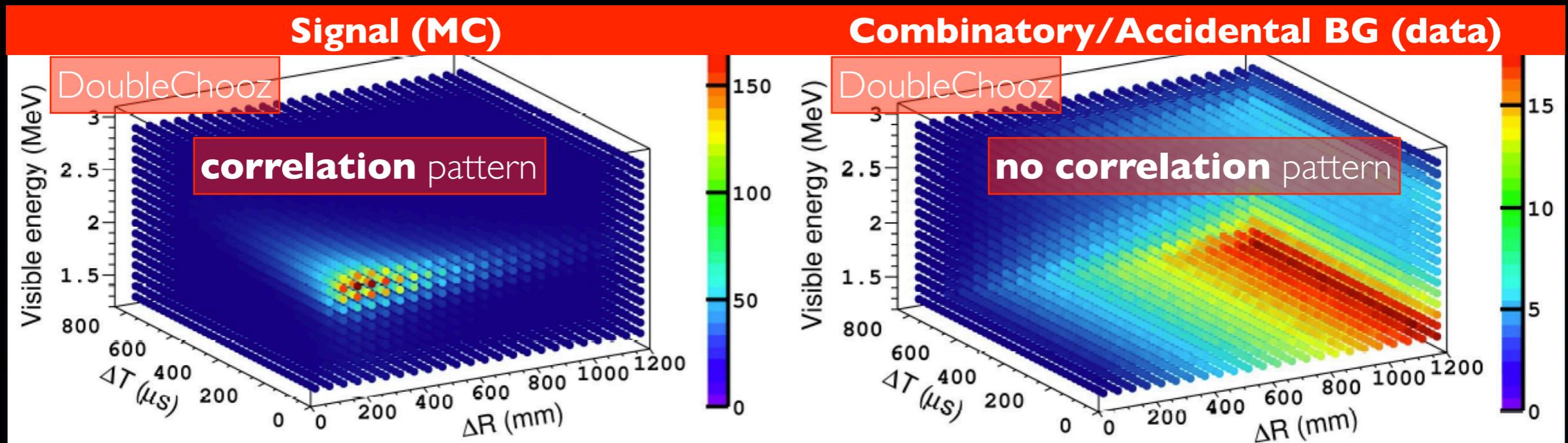
at right **place \oplus time \oplus energy \oplus PID — many orders of magnitude**



easier to lower **combinatory-BG** (~3 orders of magnitude) than **cosmogenic-BG** (~1 order of magnitude)

“combinatory” background...

at $\sim 1 \text{ MeV}$ **backgrounds** due to **radiogenic** (radio-purity only $\sim 10^{-15} \text{ g/g}$) & **cosmogenic**



LiquidO's vertex precision sub-cm — major

CC antineutrino (reactor)

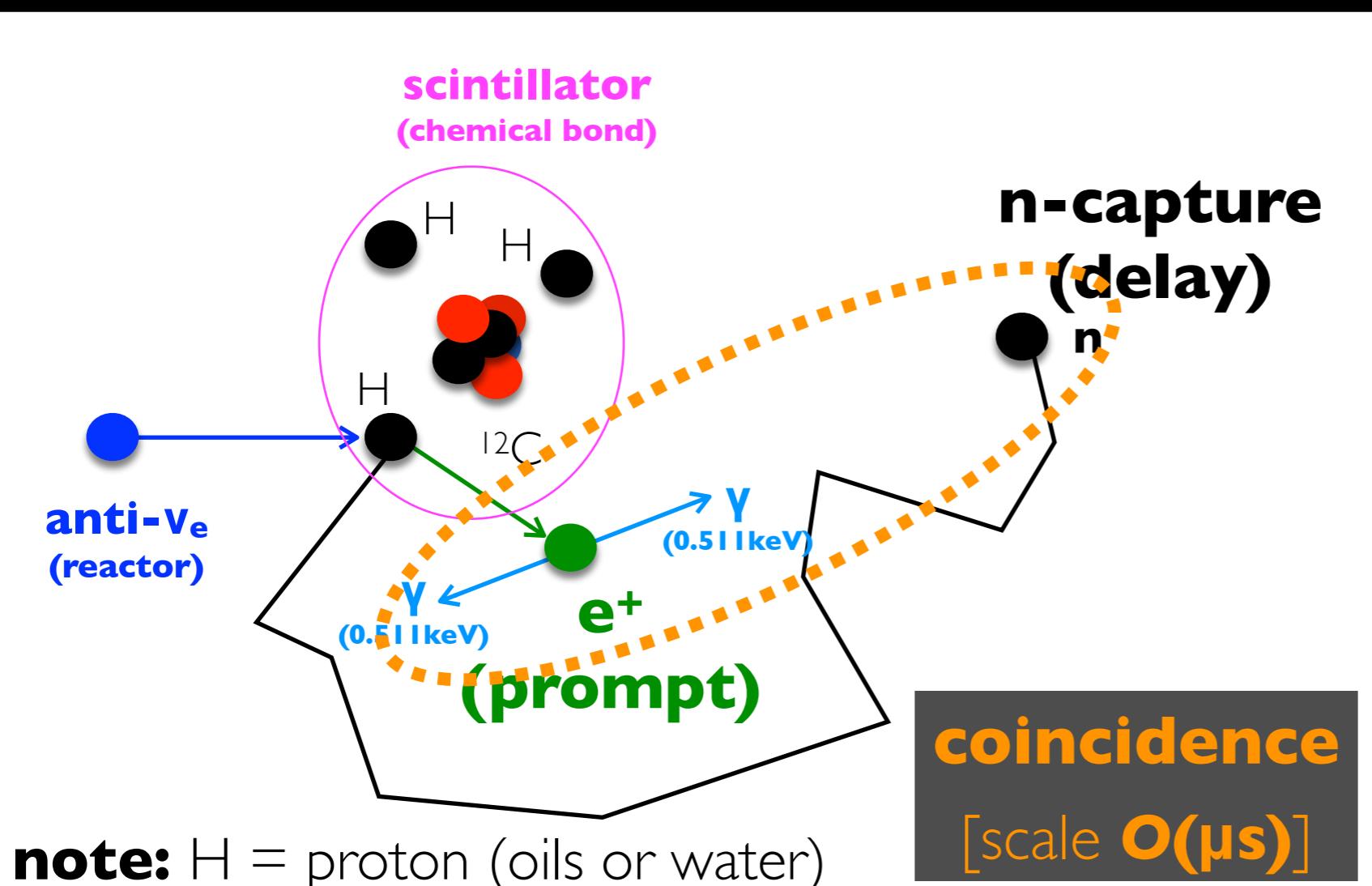
- **sub-dominant BG** — enough shielding (easy)
- **dominant: correlated cosmogenic**

CC neutrino (solar)

- **dominant BG: β^- decay of ^{115}In**
[^{14}C is lower: oil from underground petrol]

inverse- β decay (IBD) interaction...

IBD: anti- ν_e + p → e^+ + n



IBD detection art...

- n-H (native)**
- n-C (native oil)**
- n-O? (native water)**
- n-Cd** (non-native)
- n-Li** (non-native)
- n-Gd** (non-native)
- 3He** (non-native)

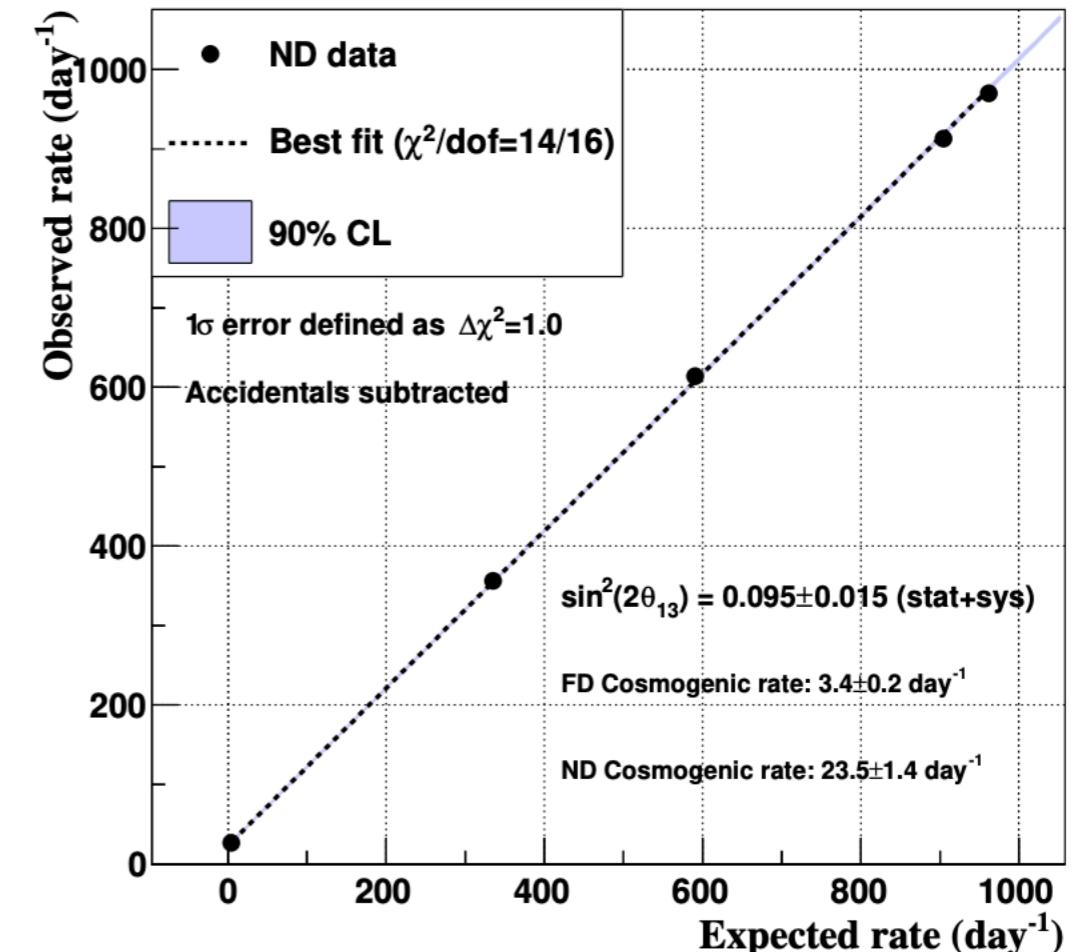
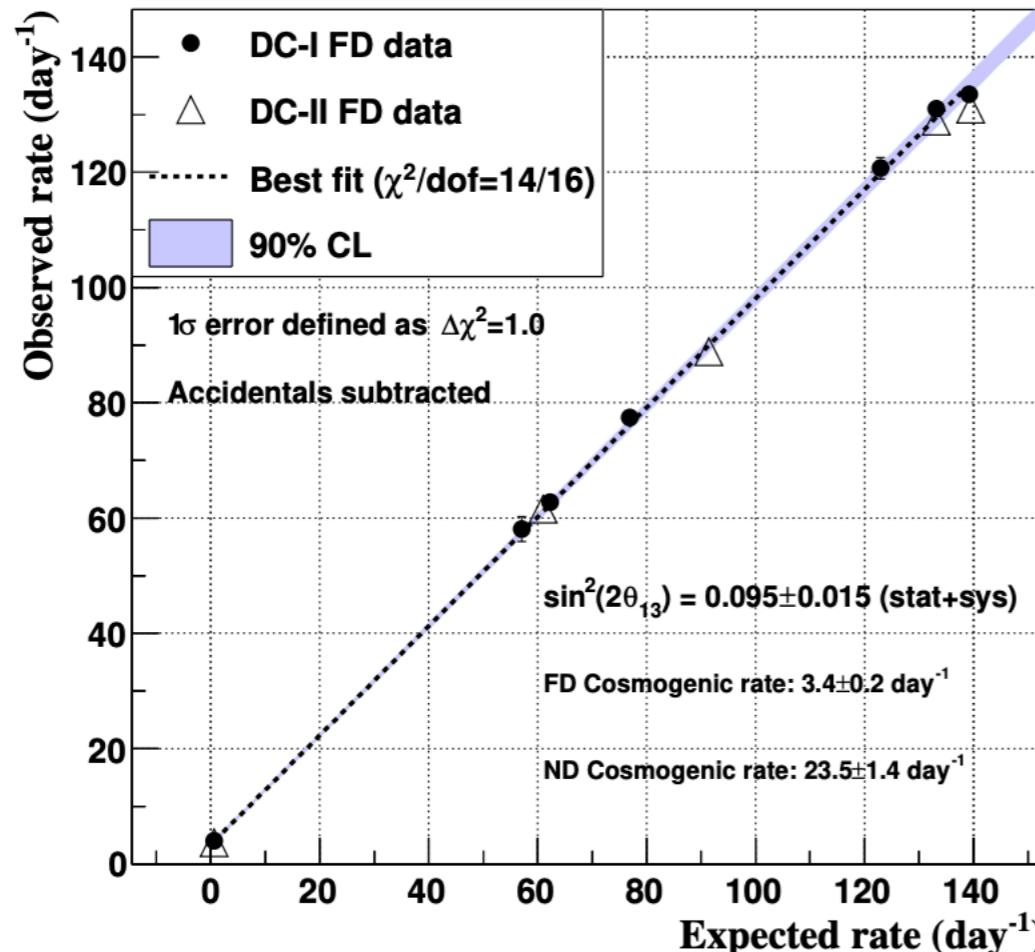
how to catch the n?

no e^+ PID implies

$\gamma \approx e^- \approx e^+ \approx a \approx p\text{-recoil (fast-n)}$

Rate⊕Shape⊕Rate-Modulation Analysis: differentiates . . .

- something **scaling with reactor power**: **reactor-signal**
- something **not-scaling with reactor power**: **background or non-reactor-signal**

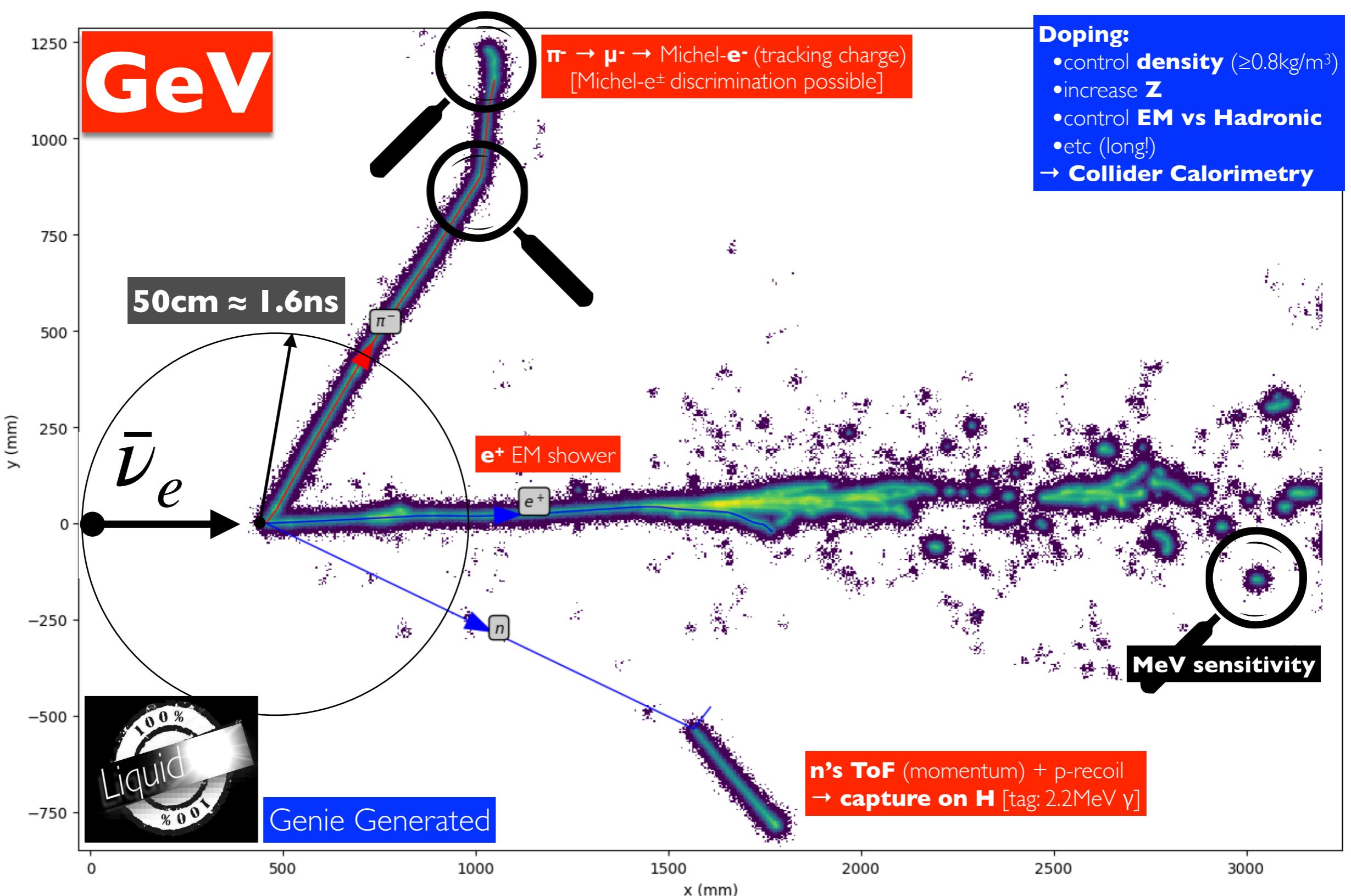


Rate⊕Shape⊕Rate-Modulation Analysis: works . . .

- for any “something” (agnostic) — **explorations beyond IBD?**
- Signal & BG estimates are **inclusive** — i.e. **Signal-model or BG-model indentent**
⇒ **no need to understand your BG!!** (except one little caveat)

Powerful way to target unknown discovery channels!

complex GeV with LiquidO ...



Stochastic calorimetry order 0.1% [$\sim 10^5 \text{ PE/GeV}$] — excellent control of non-stochastic