TWENTY-SECOND LOMONOSOV CONFERENCE ON ELEMENTARY PARTICLE PHYSICS

Moscow, August 21-27, 2025

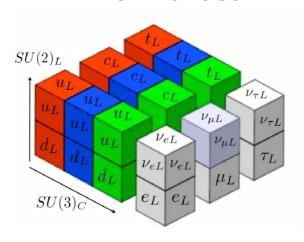


Quark-lepton mixing and the leptonic CP violation

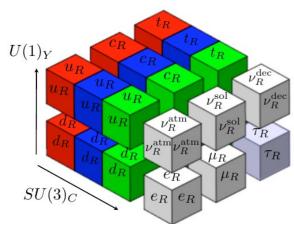
Davide Meloni Dipartimento di Matematica e Fisica Universita' Roma Tre

The Standard Model of Particle Physics

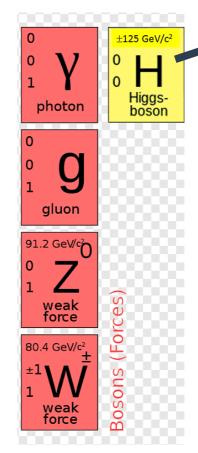
Left-handed



Right-handed



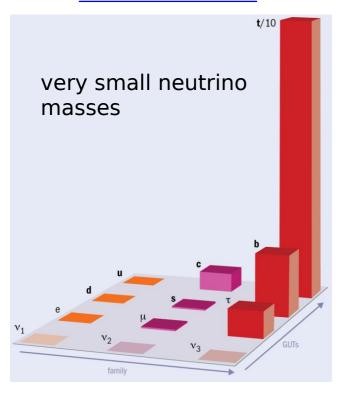




S.King, talk at Bethe Forum on Modular Flavor Symmetries Scalar sector

The Flavor Problem

Mass hierarchies



$$m_d \ll m_s \ll m_b$$
, $\frac{m_d}{m_s} = 5.02 \times 10^{-2}$,

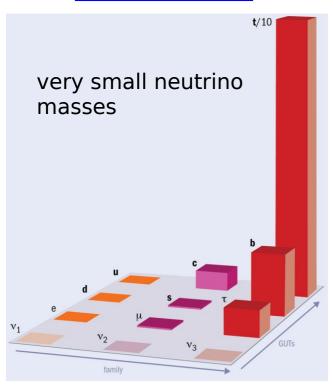
$$m_u \ll m_c \ll m_t \,, \,\,\, \frac{m_u}{m_c} = 1.7 \times 10^{-3} \,,$$

$$\frac{m_s}{m_b} = 2.22 \times 10^{-2} \,, \ m_b = 4.18 \ {
m GeV};$$

$$\frac{m_c}{m_t} = 7.3 \times 10^{-3}$$
, $m_t = 172.9$ GeV;

The Flavor Problem

Mass hierarchies



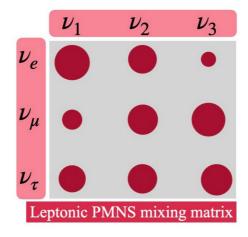
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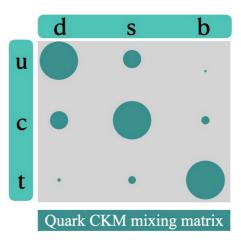
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Fermion mixing



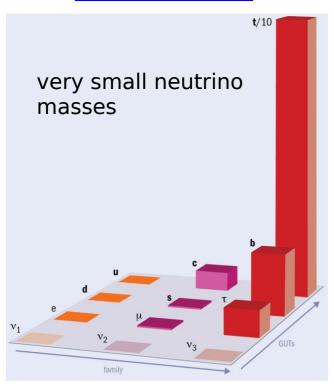




almost a diagonal matrix

The Flavor Problem

Mass hierarchies



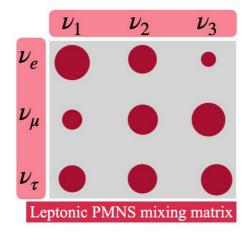
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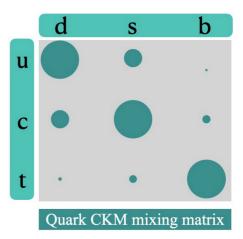
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Fermion mixing





all mixing are large but the 13 element

almost a diagonal matrix



Why are they so different?

Suggested solutions

* Hierarchical Pattern

Froggatt-Nielsen mechanism

$$L \sim \overline{\Psi}_L H \Psi_R \left(\frac{\theta}{\Lambda}\right)^n$$

Too many O(1) coefficients

Works better for small mixing: good for quarks, no for neutrinos

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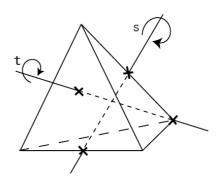
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* mixing angles

elegant explanation: non-Abelian discrete flavour symmetries



Complicated scalar sector. Good for neutrinos, not for quarks

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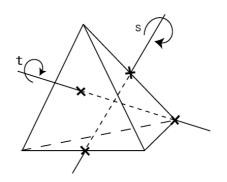
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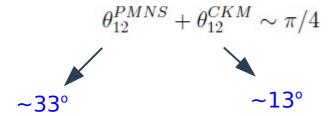
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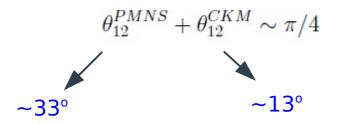
Complicated scalar sector. Good for neutrinos, not for quarks

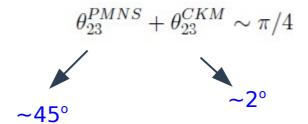
Mixings in the lepton and hadron sector are unrelated?

Quark-Lepton complementarity



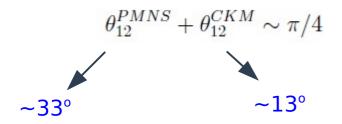
Quark-Lepton complementarity

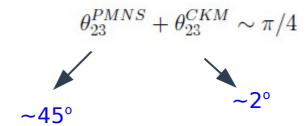




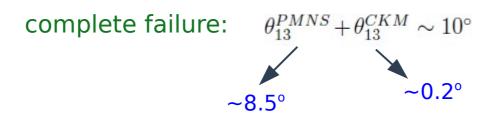
- appealing from a theoretical and phenomenological point of view
- no clue on which kind of symmetry could be responsible for them

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we need to replace the bad relation with a promising one:

$$\theta_{13}^{PMNS} = O(1) \cdot \theta_{12}^{CKM}$$

same order of magnitude

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Flavor symmetries

$$\begin{pmatrix} x & y & y \\ y & z & x-z \\ y & x-z & z \end{pmatrix}$$
 diagonalized b

neutrino mass
$$\begin{pmatrix} x & y & y \\ y & z & x-z \\ y & x-z & z \end{pmatrix} \xrightarrow{\text{diagonalized by}} \begin{pmatrix} \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 \\ \frac{1}{2} & \frac{1}{2} & -\frac{1}{\sqrt{2}} \\ \frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} = U_v$$

$$\theta_{13}^{PMNS} = 0$$
 $\theta_{12}^{PMNS} = 45^{\circ}$ $\theta_{23}^{PMNS} = 45^{\circ}$

good starting point

Flavor symmetries

Altarelli et al., 0903.1940

Corrections are needed from charged lepton diagonalization

$$U^{PMNS} = U_{cl}^{+} \cdot U_{v} \qquad U_{cl} \sim \begin{vmatrix} 1 & \lambda_{C} & \lambda_{C} \\ \lambda_{C} & 1 & 0 \\ \lambda_{C} & 0 & 1 \end{vmatrix}$$

introduced by hand (me/mu $\sim \lambda_c^2$)

good results:

$$\sin^2 \theta_{12} = \frac{1}{2} - O(\lambda_C)$$

$$\sin^2 \theta_{23} = \frac{1}{2}$$

$$\sin \theta_{13} = \frac{1}{\sqrt{2}} O(\lambda_C)$$

$$\left[\theta_{13}^{PMNS} = O(1) \cdot \theta_{12}^{CKM}\right]$$

<u>GUT: simple example from SU(5)</u>

Let us take the electron and down quark relation:

$$m_e = m_D^T$$

$$U^{PMNS} = U_{cl}^+ \cdot U_v$$

$$V^{CKM} = U_u^+ \cdot U_d$$

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Let us diagonalize the matrices:

$$U_{cl}\,m_e\,E_R^+=\!m_e^D$$
 $U_d\,m_d\,D_R^+=\!m_d^D$ this implies
$$U_{cl}\!=\!D_R^* \qquad \qquad U_d\!=\!E_R^*$$

relations involve <u>unobservable</u> right-handed rotations

Our approach

Our point of view: assume a dependence of neutrino mixing on the CKM

$$U^{PMNS}\!=\!V_{C\!K\!M}^*\!\cdot\!T^*$$

$$T = U_{23}(\widetilde{\theta}_{23})U_{13}(\widetilde{\theta}_{13})U_{12}(\widetilde{\theta}_{12})$$

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Strategy:

- take T as the well know leading order results
- correct them to match the experimental values of angles and phases
- check for neutrino mass predictions as well

Parameter	Best-fit value and 1σ range
$r \equiv \Delta m_{\rm sol}^2 / \Delta m_{\rm atm}^2 $	0.0295 ± 0.0008
$\tan(\theta_{12})$	0.666 ± 0.019
$\sin(\theta_{13})$	0.149 ± 0.002
$\tan(\theta_{23})$	0.912 ± 0.035
$J_{ m CP}$	-0.027 ± 0.010

$$J_{CP} = \text{Im} \left[(U_{PMNS})_{11} (U_{PMNS})_{12}^* (U_{PMNS})_{21}^* (U_{PMNS})_{22} \right]$$

Is the ansatz successfull?

No CKM corrections

$$U^{PMNS} = V_{CKM}^* \cdot T^*$$

$$T = U_{23}(\widetilde{\theta}_{23}) U_{13}(\widetilde{\theta}_{13}) U_{12}(\widetilde{\theta}_{12})$$

$$T_{BM} = \begin{pmatrix} \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 \\ \frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ -\frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad T_{TBM} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} \quad T_{GR} = \begin{pmatrix} c_{12} & s_{12} & 0 \\ \frac{s_{12}}{\sqrt{2}} & -\frac{c_{12}}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{s_{12}}{\sqrt{2}} & -\frac{c_{12}}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \end{pmatrix} \quad \text{Bi-maximal mixing}$$

No CKM corrections

$$U^{PMNS} = V_{CKM}^* \cdot T^*$$

$$T = U_{23}(\widetilde{\theta}_{23})U_{13}(\widetilde{\theta}_{13})U_{12}(\widetilde{\theta}_{12})$$

$$\tan \theta_{12} = 1/\phi$$
, with $\phi = (1 + \sqrt{5})/2$

$$T_{BM} = \left(\begin{array}{ccc} \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 \\ \frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ -\frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{array} \right)$$
 Bi-maximal mixing

$$T_{TBM} = \left(\begin{array}{ccc} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{array} \right)$$
 maximal

mixing

$$T_{BM} = \begin{pmatrix} \frac{1}{\sqrt{2}} & -\frac{1}{\sqrt{2}} & 0 \\ \frac{1}{2} & \frac{1}{2} & \frac{1}{\sqrt{2}} \\ -\frac{1}{2} & -\frac{1}{2} & \frac{1}{\sqrt{2}} \end{pmatrix} \qquad T_{TBM} = \begin{pmatrix} \sqrt{\frac{2}{3}} & \frac{1}{\sqrt{3}} & 0 \\ -\frac{1}{\sqrt{6}} & \frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{6}} & -\frac{1}{\sqrt{3}} & \frac{1}{\sqrt{2}} \end{pmatrix} \qquad T_{GR} = \begin{pmatrix} c_{12} & s_{12} & 0 \\ \frac{s_{12}}{\sqrt{2}} & -\frac{c_{12}}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{s_{12}}{\sqrt{2}} & -\frac{c_{12}}{\sqrt{2}} & -\frac{1}{\sqrt{2}} \end{pmatrix}$$

$$\text{Tri-Bi-maximal mixing} \qquad \text{Tri-Bi-maximal maximal mixing} \qquad \text{Tri-Bi-maximal maximal m$$

no CKM corrections

$$\sin(\theta_{13})=0$$

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$$\tan(\theta_{23}) = 1$$

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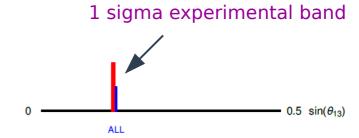
$$\tan(\theta_{12}) = 1$$

$$\tan(\theta_{12}) = \frac{1}{\sqrt{2}}$$

$$\tan(\theta_{12}) = \frac{2\sqrt{5}}{\sqrt{5+\sqrt{5}}}$$

$$U^{PMNS} = V_{CKM}^* \cdot T^*$$

after CKM corrections



\overline{T}	$\sin(\theta_{13})$	$\tan(\theta_{12})$	$\tan(\theta_{23})$	$J_{ m CP}$
U_{TBM}	$\frac{\lambda}{\sqrt{2}}$	$\frac{1}{\sqrt{2}} + \frac{3\lambda}{2\sqrt{2}}$	1	$-\frac{1}{6}A\eta\lambda^3$
U_{BM}	$\frac{\lambda}{\sqrt{2}}$	$1-\sqrt{2}\lambda$	1	$\frac{1}{4\sqrt{2}}A\eta\lambda^3$
U_{GR}	$\frac{\lambda}{\sqrt{2}}$	$\frac{2\sqrt{5}}{5+\sqrt{5}} + \frac{5\sqrt{2}}{5+\sqrt{5}}\lambda$	1	$-\frac{1}{2\sqrt{10}}A\eta\lambda^3$



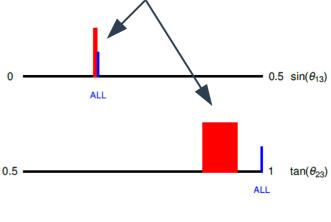
$$\theta_{13}^{PMNS} = O(1) \cdot \theta_{12}^{CKM}$$

$$U^{PMNS} = V_{CKM}^* \cdot T^*$$

after CKM corrections

T	$\sin(\theta_{13})$	$\tan(\theta_{12})$	$ an(heta_{23})$	$J_{ m CP}$ 0.5
U_{TBM}	$\frac{\lambda}{\sqrt{2}}$	$\frac{1}{\sqrt{2}} + \frac{3\lambda}{2\sqrt{2}}$	1	$-\frac{1}{6}A\eta\lambda^3$
U_{BM}	$\frac{\lambda}{\sqrt{2}}$	$1-\sqrt{2}\lambda$	1	$\frac{1}{4\sqrt{2}}A\eta\lambda^3$
U_{GR}	$\frac{\lambda}{\sqrt{2}}$	$\frac{2\sqrt{5}}{5+\sqrt{5}} + \frac{5\sqrt{2}}{5+\sqrt{5}}\lambda$	1	$-\frac{1}{2\sqrt{10}}A\eta\lambda^3$







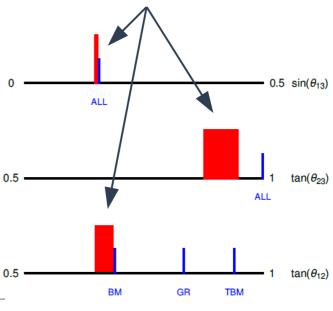
maximal atmospheric mixing

$$U^{PMNS} = V_{CKM}^* \cdot T^*$$

after CKM corrections

T	$\sin(\theta_{13})$	$\tan(\theta_{12})$	$\tan(\theta_{23})$	$J_{ m CP}$ 0.
U_{TBM}	$\frac{\lambda}{\sqrt{2}}$	$\frac{1}{\sqrt{2}} + \frac{3\lambda}{2\sqrt{2}}$	1	$-\frac{1}{6}A\eta$.
U_{BM}	$\frac{\lambda}{\sqrt{2}}$	$1-\sqrt{2}\lambda$	1	$\frac{1}{4\sqrt{2}}A\eta$.
U_{GR}	$\frac{\lambda}{\sqrt{2}}$	$\frac{2\sqrt{5}}{5+\sqrt{5}} + \frac{5\sqrt{2}}{5+\sqrt{5}}\lambda$	1	$-\frac{1}{2\sqrt{10}}A\eta\lambda^{-}$

1 sigma experimental band





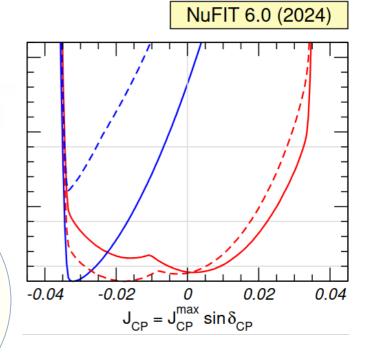
good for BM only

$$U^{PMNS} = V_{CKM}^* \cdot T^*$$

after CKM corrections



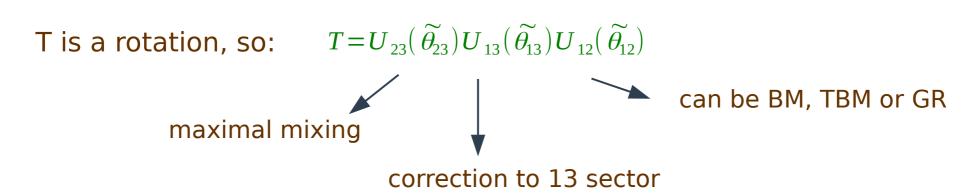
T	$\sin(\theta_{13})$	$\tan(\theta_{12})$	$\tan(\theta_2)$	$_{3})$ $J_{ m CP}$
U_{TBM}	$\frac{\lambda}{\sqrt{2}}$	$\frac{1}{\sqrt{2}} + \frac{3\lambda}{2\sqrt{2}}$	1	$-\frac{1}{6}A\eta\lambda^3$
U_{BM}	$\frac{\lambda}{\sqrt{2}}$	$1-\sqrt{2}\lambda$	1	$\frac{1}{4\sqrt{2}}A\eta\lambda^3$
U_{GR}	$\frac{\lambda}{\sqrt{2}}$	$\frac{2\sqrt{5}}{5+\sqrt{5}} + \frac{5\sqrt{2}}{5+\sqrt{5}}\lambda$	1	$-\frac{1}{2\sqrt{10}}A\eta\lambda^3$





too small CP violation

How to correct the wrong predictions?



How to correct the wrong predictions?

T is a rotation, so:
$$T = U_{23}(\widetilde{\theta}_{23})U_{13}(\widetilde{\theta}_{13})U_{12}(\widetilde{\theta}_{12})$$



correction to 12 sector

correction to 23 sector

correction to 13 sector

$$U_{13} = \begin{pmatrix} 1 - \frac{\lambda^2}{2} |u|^2 & 0 & u\lambda \\ 0 & 1 & 0 \\ -u^*\lambda & 0 & 1 - \frac{\lambda^2}{2} |u|^2 \end{pmatrix} \quad \text{u = complex parameter}$$

$$J_{CP} \sim \lambda * \Im(u)$$

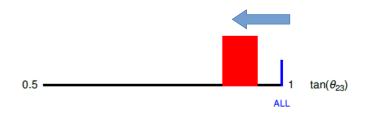


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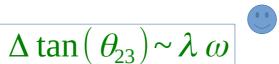


correction to 23 sector



2.
$$U_{23} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1}{\sqrt{2}} - \lambda\omega - \sqrt{2}\lambda^{2}\omega^{2} - 2\lambda^{3}\omega^{3} & \frac{1}{\sqrt{2}} + \omega\lambda \\ 0 & -\frac{1}{\sqrt{2}} - \omega\lambda & \frac{1}{\sqrt{2}} - \lambda\omega - \sqrt{2}\lambda^{2}\omega^{2} - 2\lambda^{3}\omega^{3} \end{pmatrix}$$

 ω = real parameter fixed by fit



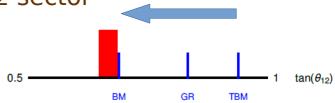


How to correct the wrong predictions?

T is a rotation, so: $T = U_{23}(\widetilde{\theta}_{23})U_{13}(\widetilde{\theta}_{13})U_{12}(\widetilde{\theta}_{12})$



correction to 12 sector



3.
$$U_{12} = \begin{pmatrix} K & \tilde{s}_{12} + z\lambda & 0 \\ -\tilde{s}_{12} - z\lambda & K & 0 \\ 0 & 0 & 1 \end{pmatrix}$$



z= real parameter
fixed by fit

 $\Delta \tan(\theta_{12}) \sim \lambda z$

Conclusions

 χ^2 4 parameter-fit of u, ω and z: all patterns agree with experiments

 \rightarrow our ansatz $U^{PMNS} = V_{CKM}^* \cdot T^*$ is phenomenological viable

atmospheric

angle

	Pattern	$\operatorname{Re}\left(u\right)$	$\operatorname{Im}\left(u\right)$	ω	z
$\chi^2 \sim 0$	TBM	-0.27	0.57	-0.27	-0.50
	BM	-0.27	0.57	-0.27	0.08
	GR	-0.27	0.57	-0.27	-0.73
					

CP violation

solar angle

Backup slides

Neutrino masses

