



TWENTY-SECOND LOMONOSOV CONFERENCE August, 21-27, 2025 ON ELEMENTARY PARTICLE PHYSICS MOSCOW STATE UNIVERSITY

*Neutrino oscillation inside supernovae:
exact description*

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23/8/2025



Outlines

- Neutrino Oscillations: 2 ν vs 3 ν Scheme
- Neutrino Oscillations in 3 ν 's Scheme
- Non-Adiabatic Resonances
- Numerical Results
- Conclusion



Neutrino Oscillations: 2 nu vs 3 nu Scheme

The flavor basis $|\nu_\alpha\rangle$ ($\alpha = e, \mu$) is related to mass basis $|\nu_i\rangle$ ($i = 1, 2$) as

$$|\nu_\alpha\rangle = \sum_{i=1}^2 X_{\alpha i} |\nu_i\rangle, \text{ with } X = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \text{ Then,}$$

$$E_i \approx p + \frac{m_i^2}{2p} \approx E + \frac{m_i^2}{2E}. \quad H = \frac{1}{2E} \begin{pmatrix} m_1^2 & 0 \\ 0 & m_2^2 \end{pmatrix}. \quad H_f = X H_m X^\dagger,$$

$$H_{vac} = \frac{\Delta m^2}{4E} \begin{pmatrix} -\cos 2\theta & \sin 2\theta \\ \sin 2\theta & \cos 2\theta \end{pmatrix}$$

Inside matter, this Hamiltonian is generalized into

$$H_{mat} = \frac{\Delta m^2}{4E} \begin{pmatrix} -\cos 2\theta + A & \sin 2\theta \\ \sin 2\theta & \cos 2\theta - A \end{pmatrix}, \quad A \equiv \frac{2\sqrt{2}G_F N_e E}{\Delta m^2}.$$

that gives the eigenvalues $\lambda_\pm = \pm \frac{\Delta m_m^2}{4E}$, via the mixing angle $\tan 2\theta_m = \frac{\sin 2\theta}{\cos 2\theta - A}$,

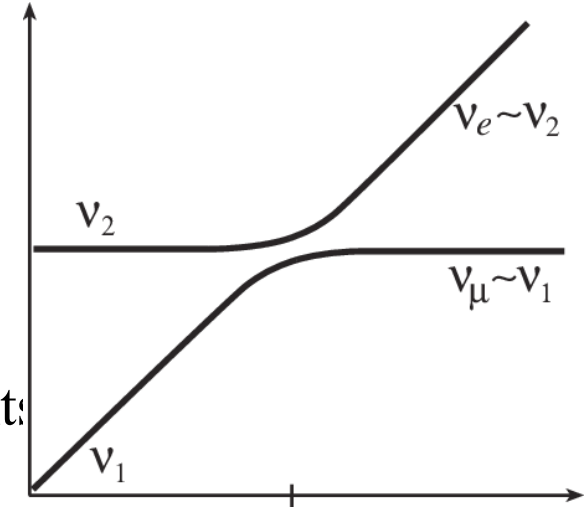
$$\Delta m_m^2 = \Delta m^2 \sqrt{(\cos 2\theta - A)^2 + \sin^2 2\theta}.$$

If the matter density is changing, the Schrodinger equation is written as

$$i \frac{d}{dr} \nu_i(r) = H_{ik}^m \nu_k(t) = U_M^\dagger H_{ik}^f U_M \nu_k(t) + i \left(\frac{d}{dr} U_M^\dagger(r) \right) U_M(r).$$

and the mass eigenstates behave as

$$\tilde{H}_{mat} = \begin{pmatrix} \lambda_- & -i \frac{\partial \theta_m}{\partial r} \\ i \frac{\partial \theta_m}{\partial r} & \lambda_+ \end{pmatrix}$$



The adiabaticity is ensured if the off-diagonal elements are very small w.r.t. the diagonal ones, i.e., $\gamma_{ij} \gg 1$

$$\gamma(x) = \frac{\Delta m_m^2(x)}{2E} \left[\frac{d}{dx} \theta_m(x) \right]^{-1}$$

If not, a possible jump could happen between λ_{\pm} at (around) the resonance region that defined by $(H_{mat})_{11} = (H_{mat})_{22}$ or by minimizing $|\lambda_+ - \lambda_-|$

At the resonance, we got $\gamma_{ij}(x) = \frac{\Delta m_{ij}^2 \sin^2 2\theta_{ij}}{2E \cos 2\theta_{ij}} \left| \frac{d \ln N_e(x)}{dx} \right|_{x_{res}}^{-1}$.

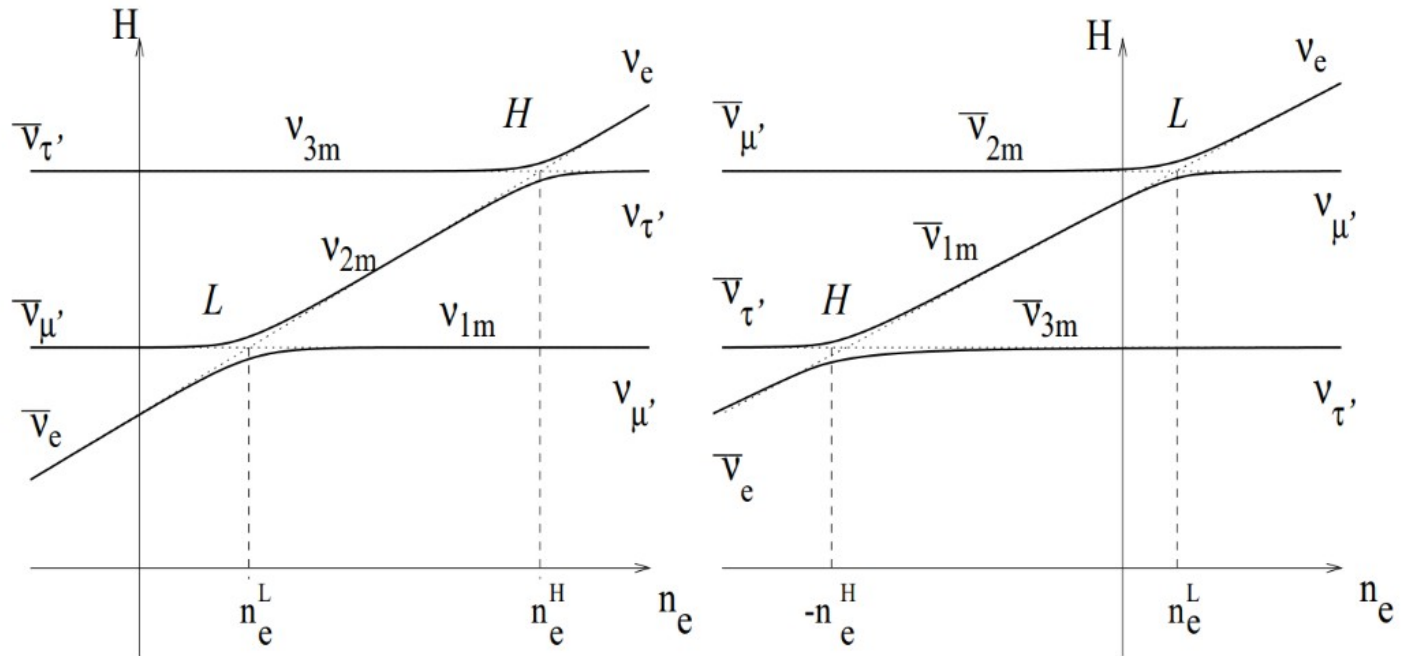
The transition probability is given by the LZ formula

$$P_{LZ} = \exp\left(-\frac{\pi}{2}\gamma\right) = \exp\left(-\frac{\pi}{4} \frac{(\Delta m_{ij}^2) \sin^2 2\theta_{ij}}{E \cos 2\theta_{ij} \left|\frac{d \ln N_e}{dr}\right|_{\text{res}}}\right).$$

In the 3nu scheme, the matter Hamiltonian can be written in flavor modified basis as

$$i \frac{d}{dx} \begin{pmatrix} \nu_e \\ \nu'_\mu \\ \nu'_\tau \end{pmatrix} = \frac{1}{2E} \begin{pmatrix} \#_{11} + A & \#_{12} & \#_{13} \\ \#_{12} & \#_{22} & 0 \\ \#_{13} + 0 & 0 & \#_{33} \end{pmatrix} \begin{pmatrix} \nu_e \\ \nu'_\mu \\ \nu'_\tau \end{pmatrix}$$

which allows to use the 2nu scheme results. Thus in a dense object like SN we have



Are these results exact? Are these approximations reliable?

Neutrino Oscillations in 3 nu's Scheme

Starting by the Schrodinger equation in the flavor basis $|\nu_\alpha\rangle^T = (\nu_e \ \nu_\mu \ \nu_\tau)$.

$$i\frac{d}{dr}|\nu_\alpha\rangle = H_f|\nu_\alpha\rangle, \quad \text{which can be written in the mass basis as}$$

$$i\frac{d}{dr}|\nu_i\rangle = U_M^\dagger H_f U_M |\nu_i\rangle + i\frac{d}{dr} \left(U_M^\dagger \right) U_M |\nu_i\rangle$$

with $|\nu_\alpha\rangle = U_M |\nu_i\rangle$. then using these dimensionless quantities, one writes

$$i\frac{d}{d\tilde{\zeta}} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix} = \left[U_M^\dagger U \begin{pmatrix} 0 & 0 & 0 \\ 0 & \alpha & 0 \\ 0 & 0 & \beta \end{pmatrix} U^\dagger U_M + \kappa U_M^\dagger \begin{pmatrix} f & 0 & 0 \\ 0 & 0 & 0 \\ 0 & 0 & 0 \end{pmatrix} U_M + \omega Q \right] \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix},$$

$$\tilde{\zeta} = r/r_0 \quad \rho = \rho_0 f(\xi) \quad \omega = \frac{df}{d\tilde{\zeta}} \quad Q = i \left(\frac{d}{df} U_M^\dagger \right) U_M \quad \kappa = 4.5416 \times 10^{12}$$

$$\alpha = 3.0205 \times 10^{-2} \left(\frac{E}{10 \text{ MeV}} \right)^{-1} \quad \beta_{NO} = 1.0219 \left(\frac{E}{10 \text{ MeV}} \right)^{-1} \quad \beta_{IH} = -1.0122 \left(\frac{E}{10 \text{ MeV}} \right)^{-1}$$

Clearly, the Hamiltonian is not real but complex Hermitian; and for the case of vanishing Dirac CP phase in the PMNS matrix, it has the form

$$\begin{pmatrix} \epsilon_1 & i\eta_{12} & i\eta_{13} \\ -i\eta_{12} & \epsilon_2 & i\eta_{23} \\ -i\eta_{13} & -i\eta_{23} & \epsilon_3 \end{pmatrix} \quad \text{Here, the epsilon's are the dimensionless energy eigenvalues, and the eta's are analogous to the effective angles derivatives.}$$

Howere, we know that $\delta_D \neq 0$, then the real part of the Hamiltonian is not diagonal!!

$$\mathcal{H}_{\Re}^{(0)} = diag\{\epsilon_1^{(0)}, \epsilon_2^{(0)}, \epsilon_3^{(0)}\} + \Re(Q^{(0)})$$

Then, it needs to be diagonalized again $diag\{\epsilon_1^{(1)}, \epsilon_2^{(1)}, \epsilon_3^{(1)}\} = U_M^{(1)\dagger} \mathcal{H}_{\Re}^{(0)} U_M^{(1)}$ and again the resulting Hamiltonian real part is not diagonal, so one diagonalizes it again! Then, the final form is

$$\mathcal{H}^{(\infty)} = \begin{pmatrix} \epsilon_1 & i\eta_{12} & i\eta_{13} \\ -i\eta_{12} & \epsilon_2 & i\eta_{23} \\ -i\eta_{13} & -i\eta_{23} & \epsilon_3 \end{pmatrix} \quad \tilde{\nu}_i = (\tilde{U}_M)_{i\alpha} \nu_\alpha$$

$$\tilde{U}_M = U_M^{(\infty)} \dots U_M^{(1)} U_M^{(0)}$$

Practically, all the unitary matrices $U_M^{(n>0)}$ are almost identical to the unit matrix for the majority of the SN density profiles, at least the power law one.

This form of $\mathcal{H}^{(\infty)}$ is now suitable to study possible transitions.

Non-Adiabatic Resonances

Assuming that all possible resonances are separated, then 3nu system with the Hamiltonian

$$\mathcal{H}^{(\infty)} = \begin{pmatrix} \epsilon_1 & i\eta_{12} & i\eta_{13} \\ -i\eta_{12} & \epsilon_2 & i\eta_{23} \\ -i\eta_{13} & -i\eta_{23} & \epsilon_3 \end{pmatrix}$$

can be reduced into 3 two nu subsystems $\{\nu_i, \nu_j\}$.

So, is it possible to have 3 (non)-adiabatic transitions with the adiabaticity parameter?

$$\gamma_{ij} = \left| \frac{\epsilon_i - \epsilon_j}{2\eta_{ij}} \right|_{res}.$$

If yes, how is the resonance region can be defined?

- by making the diagonal elements in the flavor Hamiltonian equals? (as in 2nu scheme).
- by minimizing the eigenvalues difference $|\epsilon_i - \epsilon_j|$?
- by minimizing the adiabaticity parameter γ_{ij} ? ==> This should be the most precise and meaningful?

If there exists a resonance between $\tilde{\nu}_1 \longleftrightarrow \tilde{\nu}_3$ in the SN, its corresponding density should be between the L and H resonance densities, so we call it M (middle) resonance.

The neutrino final spectrum can be written as

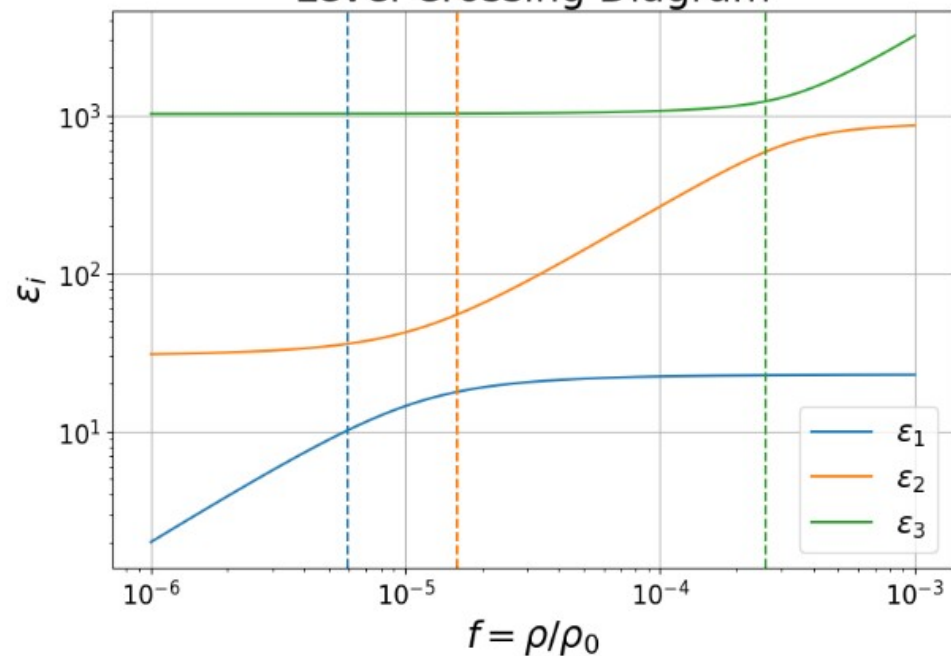
$$\begin{pmatrix} F_e \\ F_\mu \\ F_\tau \end{pmatrix} = \begin{pmatrix} |U_{e1}|^2 & |U_{e2}|^2 & |U_{e3}|^2 \\ |U_{\mu1}|^2 & |U_{\mu2}|^2 & |U_{\mu3}|^2 \\ |U_{\tau1}|^2 & |U_{\tau2}|^2 & |U_{\tau3}|^2 \end{pmatrix} \begin{pmatrix} 1-P_L & P_L & 0 \\ P_L & 1-P_L & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} 1-P_M & 0 & P_M \\ 0 & 1 & 0 \\ P_M & 0 & 1-P_M \end{pmatrix} \begin{pmatrix} 1 & 0 & 0 \\ 0 & 1-P_H & P_H \\ 0 & P_H & 1-P_H \end{pmatrix} \begin{pmatrix} F_\mu^0 \\ F_\tau^0 \\ F_e^0 \end{pmatrix}$$

$$P_L = \exp\left(-\frac{\pi}{2}\gamma_{12}\right), P_M = \exp\left(-\frac{\pi}{2}\gamma_{13}\right), P_H = \exp\left(-\frac{\pi}{2}\gamma_{23}\right)$$

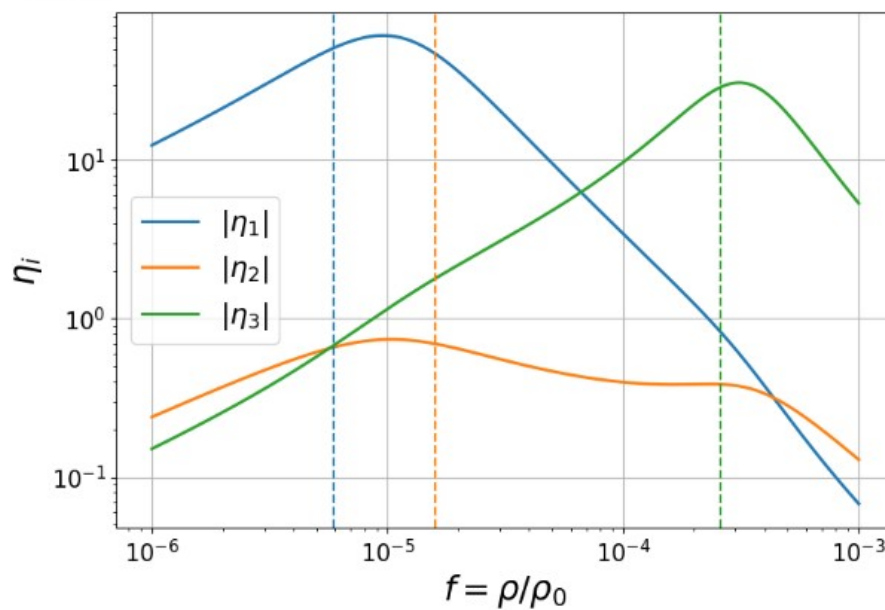
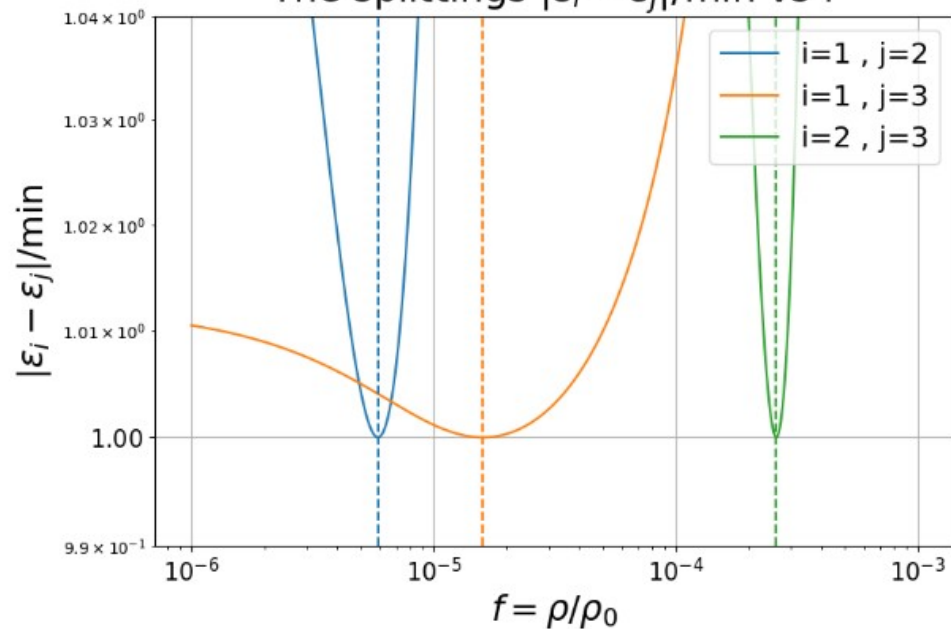
In the IH case, this definition is different and it depends on where the M resonance occurs? In the neutrino or antineutrino sector?

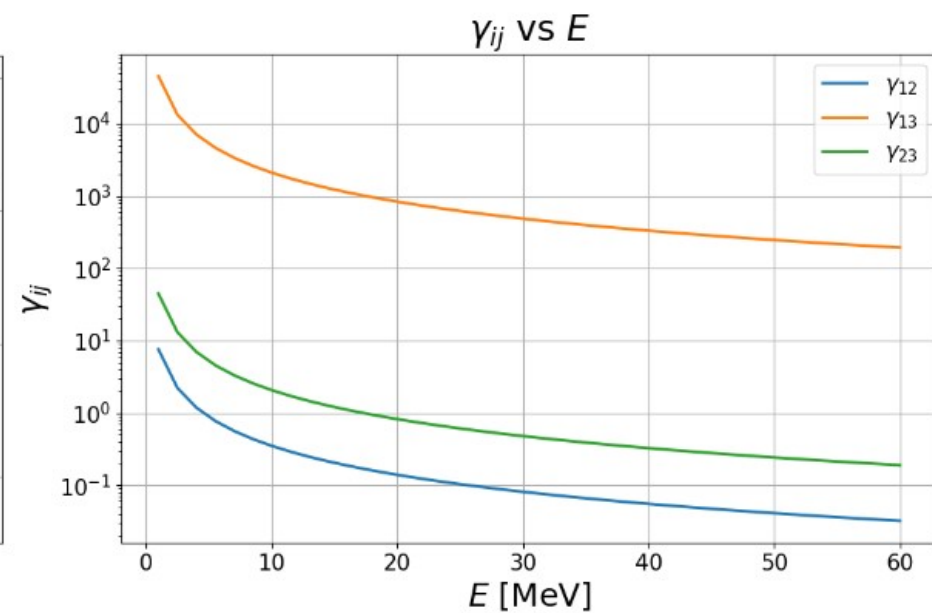
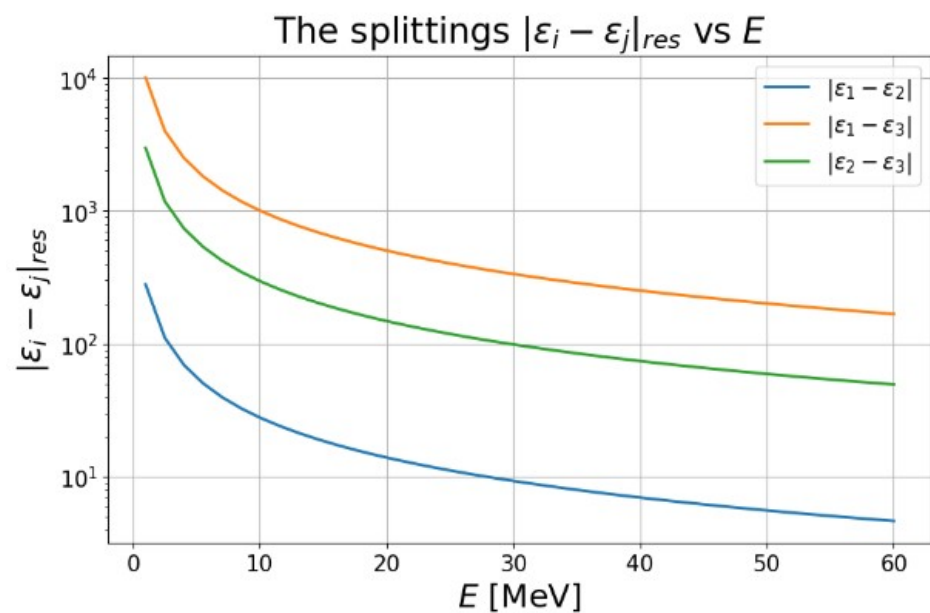
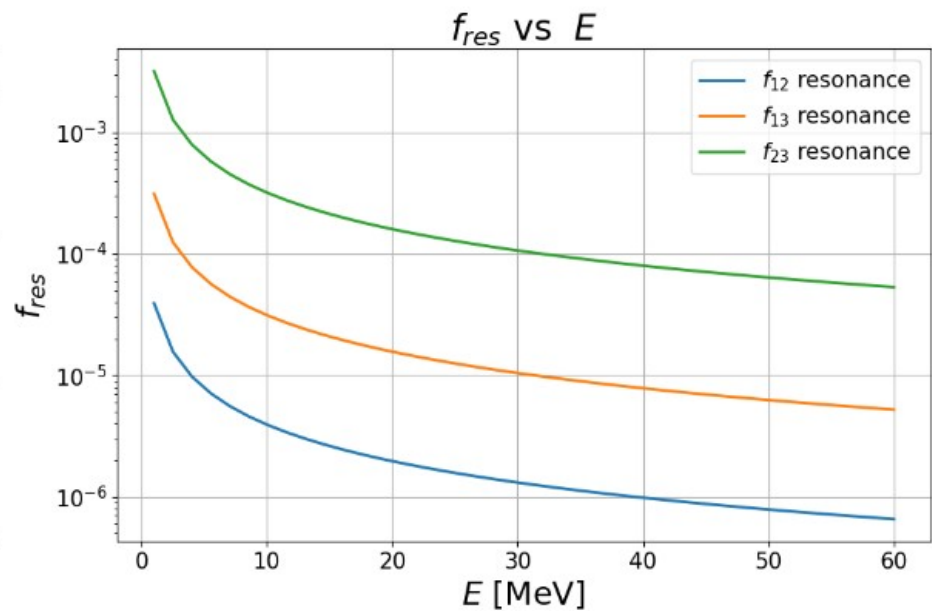
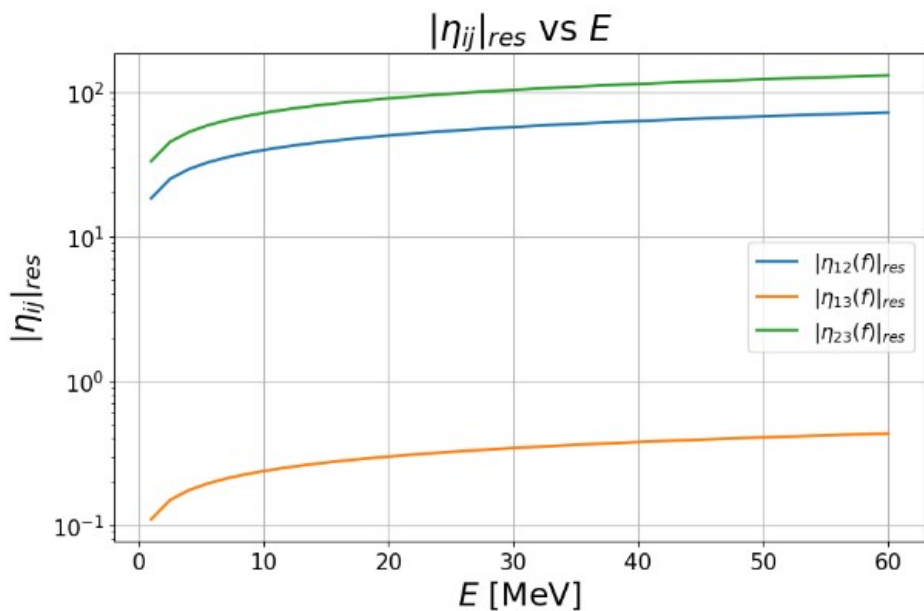
Numerical Results

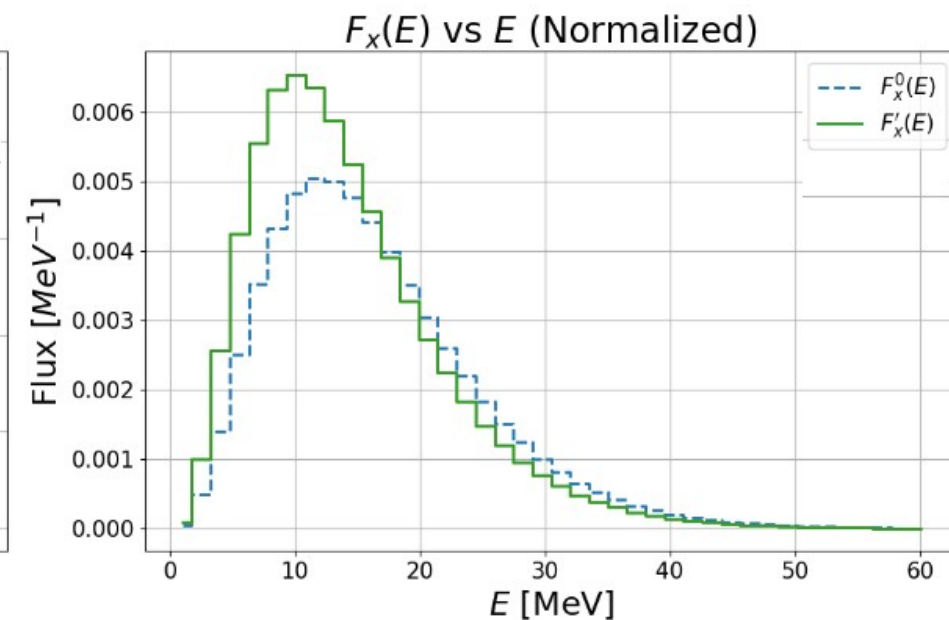
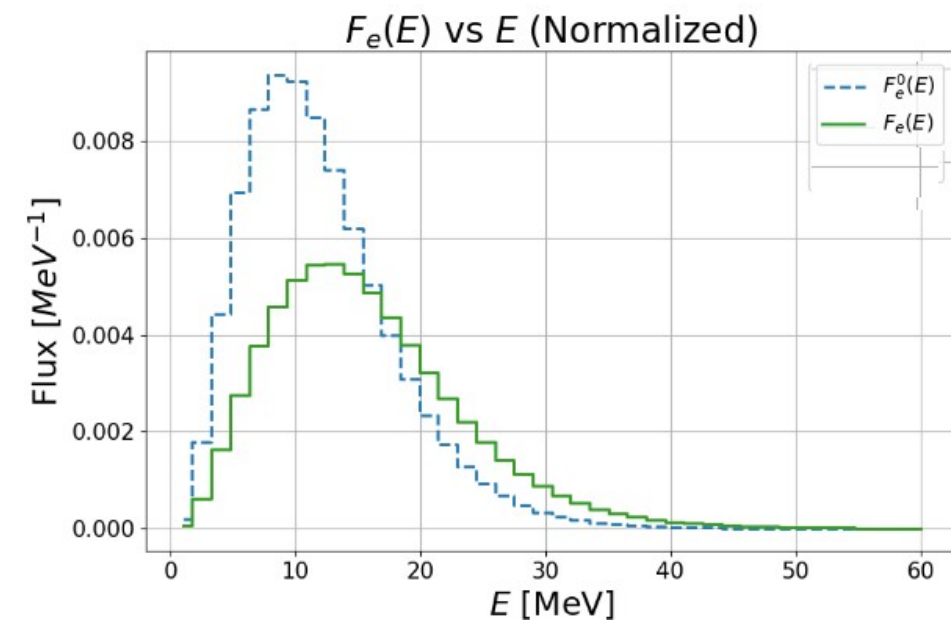
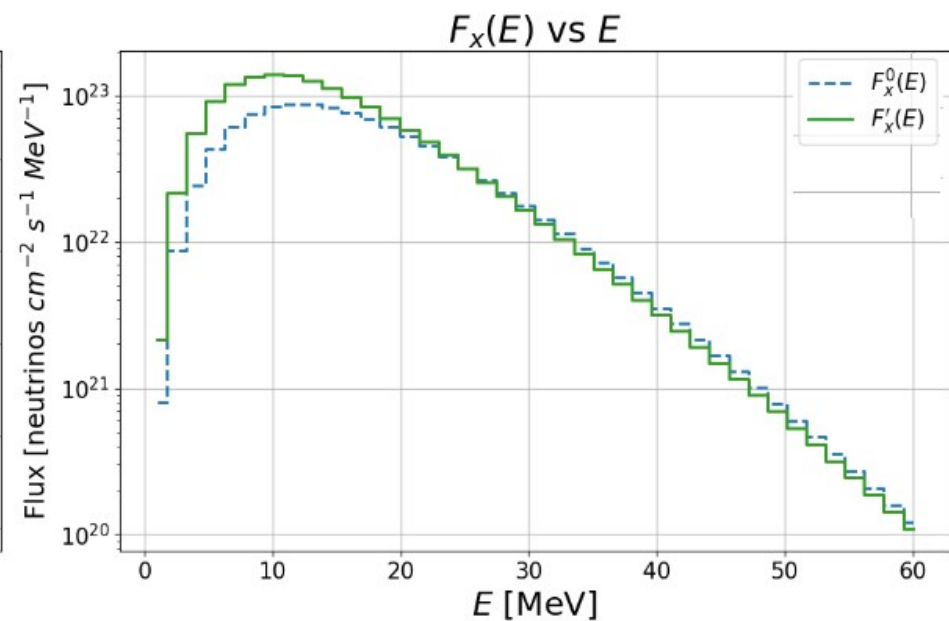
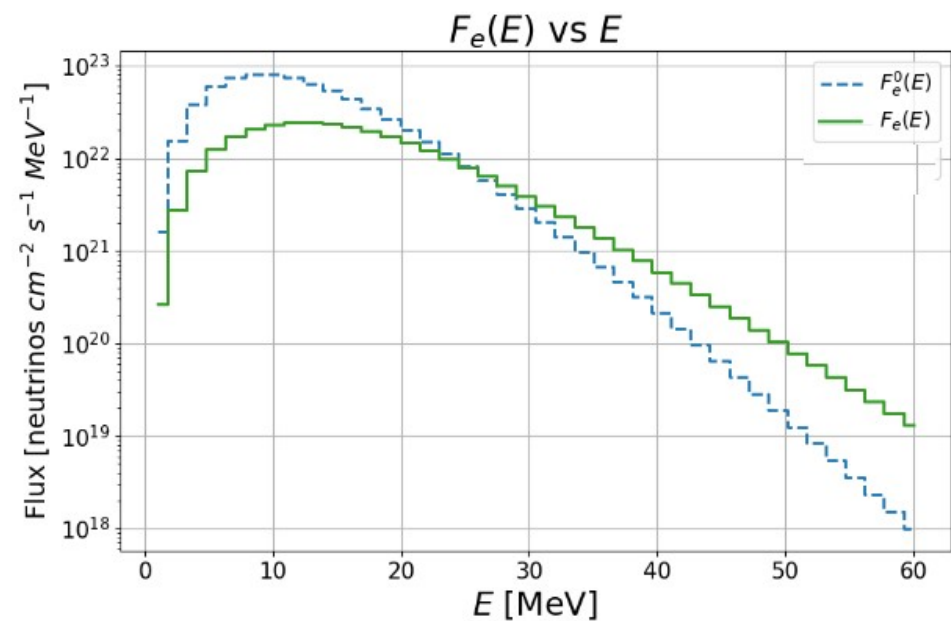
Level Crossing Diagram



The splittings $|\varepsilon_i - \varepsilon_j|/\min$ vs f







Conclusion

We numerically describe exactly the neutrino propagation inside a varying density medium (like SN), where we found:

- Possible new resonance (M, middle), purely adiabatic for a power law density profile, maybe not for realistic models.
- The resonance region should correspond a minimal adiabaticity parameter.
- For realistic SN density profiles, the neutrino final spectrum can help in probing the ν hierarchy, the SN density profile, and probably non-standard effect.

Thank You