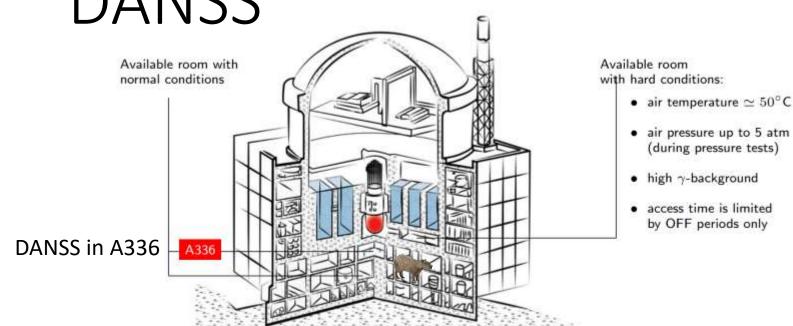


TWENTY-SECOND LOMONOSOV

CONFERENCE August, 21-27, 2025 ON ELEMENTARY PARTICLE PHYSICS MOSCOW STATE UNIVERSITY

Searches for physics beyond the SM at DANSS



Petr Gorovtsov, LPI RAS

On behalf of the DANSS collaboration

Motivation - there are several anomalies in neutrino physics

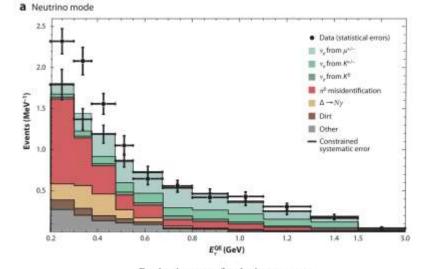
LSND, MiniBooNE, 29.05.2013, Annual Review of Nuclear and Particle Science vol 63

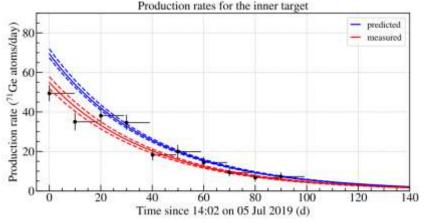
- LSND & MiniBooNE result: v_e in the v_μ beam (6 σ): $\Delta m^2 \sim 0.2 1 \, \mathrm{eV}^2$?
 - MicroBooNE does not confirm this anomaly
 P. Arbatenko et al., (MicroBooNE), Phys. Rev. Lett. 130, 011801, 2023
- Gallium anomaly (GA): deficit of ν events in the calibration measurements in gallium experiments SAGE, GALLEX and BEST (5 σ)

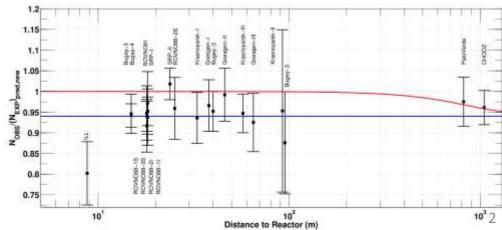
V.V. Barinov et al, (BEST), Phys. Rev. Lett. 128, 232501, 2022

• Reactor Antineutrino Anomaly (RAA): deficit of $\bar{\nu}_e$ from the reactors in comparison with theoretical predictions (2.8 σ)

G. Mention et al, Phys. Rev. D 83, 073006, 2011







Possible explanations

 Sterile neutrinos: at short distances we can consider only oscillations to sterile state

$$P(\bar{\nu}_e \to \bar{\nu}_e) \approx 1 - \sin^2(2\theta_{ee}) \sin\left(1.267 \frac{\Delta m_{41}^2 [eV^2] L[m]}{E[MeV]}\right)$$

- Large Extra Dimensions (LED) (N. Arkani-Hamed, S. Dimopoulos, G. Dvali, Phys. Let. B, vol 429, issues 3-4, pages 263-272, 1998)
 - Extra dimensions renormalize Planck's mass: $M_{Pl}^2 = (M_{Pl}^*)^{2+n}a^n$
 - $n \ge 2$ needed for Hierarchy problem but one dimension can be bigger than others. In this approach survival probability for (anti)neutrino: $P(\bar{\nu}_e \to \bar{\nu}_e) = \left|\sum_{flavours} |U_{ei}|^2 A_i\right|^2$

$$A_{i} = \left(1 - \frac{\pi^{2}}{6} m_{i}^{2} a^{2}\right)^{2} \exp\left(i \frac{m_{i}^{2} L}{2E}\right) + 2m_{i}^{2} a^{2} \exp\left(i \frac{m_{i}^{2} L}{E}\right) \sum \frac{\exp\left(i \frac{n^{2} L}{2E a^{2}}\right)}{n^{2}}, ma \ll 1$$

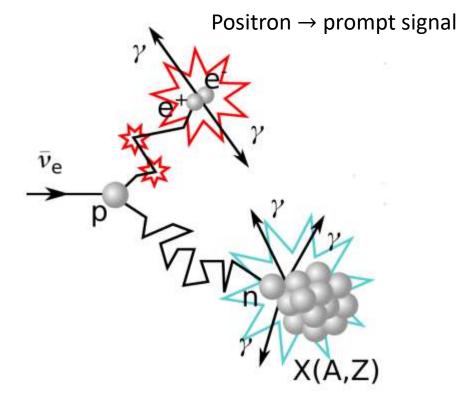
Absolute mass scale

Size of the **dominant** dimension

Kaluza-Klein tower

Detector of AntiNeutrino based on Solid Scintillator

- Detector is placed under the reactor core at Kalinin NPP and consists of safe, nonflammable materials
- The goal is to scrutinize the sterile neutrino hypothesis (and other BSM models)
- A movable platform → spectral ratio analysis not sensitive to detector efficiency and antineutrino spectrum
- Antineutrinos are detected in the IBD reaction:
- The energy of antineutrino is determined by the energy of positron



Neutron \rightarrow delayed signal

$$\overline{\nu_e} + p \to n + e^+$$

$$E_{\overline{\nu_e}} \approx E_{e^+} + 1.8 MeV$$

 $\sim 5000 events/day$

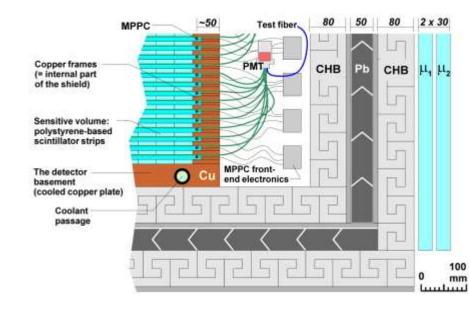
Largest in the world antineutrino statistics, IBD is used

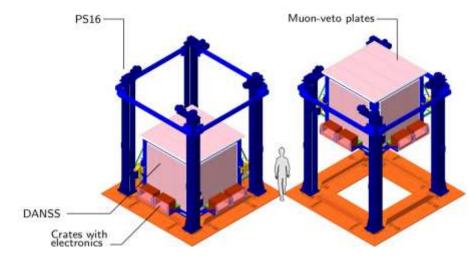
Design of the detector

• Shielding: muon veto, 8 cm of boron-doped polyethylene, 5 cm of lead, another 8 cm of boron-doped polyethylene, and a 5 cm layer of copper

 2500 scintillator strips coated with a gadoliniumcontaining layer for neutron capture. Light is collected with wavelength shifting fibers connected to SiPM and PMT

 Currently the detector's upgrade is coming: new scintillation counters, more of sensitive volume, better light collection



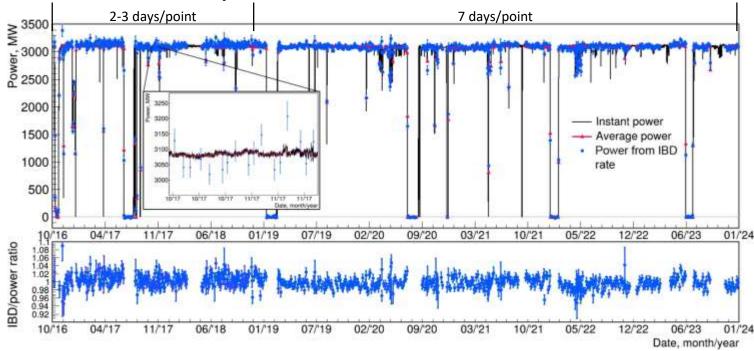


First precise measurement of reactor power with antineutrino

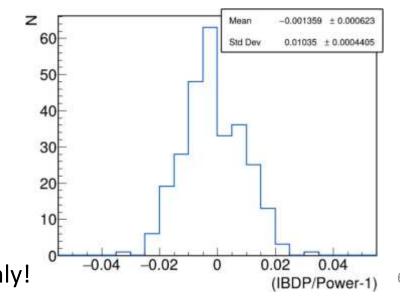
<u>I. Alekseev et al, (DANSS), Phys. Let. B, vol 866, 139575, 2025</u>

Antineutrino flux:

$$\frac{d^2N}{dE_{\nu}dt} = N_p \sigma(E_{\nu}) \varepsilon(E_{\nu}) \frac{d^2 \phi}{dE_{\nu}dt} \frac{1}{4\pi L^2} P(E_{\nu}, L)$$



- Uncertainty in flux models → normalization to flux during 1 month in 2016
- Corrections applied:
 - Change of the efficiency due to dead channels amount (< 3%)
 - Variation of the dead time due to VETO rate (1%)
 - The fuel evolution (provided by KNPP)
- Excellent agreement between DANSS and KNPP: 1.0% during 7 years
- Statistical accuracy of each point is around 0.7%
- Additional systematic between KNPP and DANSS measurements is 0.79% only!

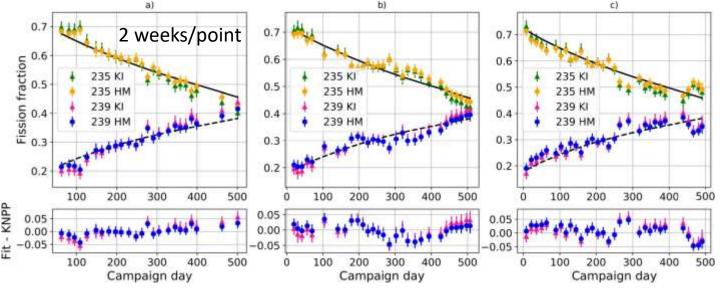


First measurement of fission fractions with antineutrino

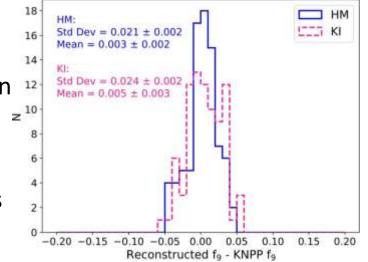
spectrum

I. Alekseev et al, (DANSS), Phys. Let. B, vol 866, 139575, 2025

 The antineutrino spectrum is different for different isotopes



- Correction for efficiency, dead time and neighbouring reactors (0.6%)
- Fit doesn't include $E_{e^+} \in [3; 5,5]$ MeV due to theoretical uncertainties
- 5th campaign is used to fix energy calibration and ^{238}U and ^{241}Pu fractions (taken 12 from KNPP)
- Statistical error of IBD rate is 0.6%
- $\bar{\nu}_e$ spectra for isotopes are taken from HM or KI model. DANSS shows that results depend weakly on the choice of the model. KNPP data is based on the solving diffusion equation for neutron flux
- Agreement with KNPP is within 2.1% although two methods based on the completely different physical processes



Oscillation analysis

- For different points in parameter space θ spectra ratios R were calculated
 - $\theta = (\sin^2 2\theta_{ee}; \Delta m_{41}^2)$ for sterile neutrino
 - $\theta = (a, m_0)$ for LED searches
- χ^2 statistics is minimized over systematic parameters η
- Gaussian CL_s method or confidence intervals (Wilks theorem assumed) were used

$$\chi^{2}(\theta,\eta) = \sum_{bins} \frac{\left(R_{bt}^{obs} - R_{bt}^{pre}(\theta,\eta)\right)^{2}}{\sigma^{2}}$$

• Data taking in 2 positions
$$R_{bt} = bottom/top$$

$$+ \sum_{bins} (R_{bt}^{obs} - R_{bt}^{pre}; R_{mbt}^{obs} - R_{mbt}^{pre}) \cdot W^{-1} \cdot \begin{pmatrix} R_{bt}^{obs} - R_{bt}^{pre} \\ R_{mbt}^{obs} - R_{mbt}^{pre} \end{pmatrix}$$

$$+ \sum_{bins} (R_{bt}^{obs} - R_{bt}^{pre}; R_{mbt}^{obs} - R_{mbt}^{pre}) \cdot W^{-1} \cdot \begin{pmatrix} R_{bt}^{obs} - R_{bt}^{pre} \\ R_{mbt}^{obs} - R_{mbt}^{pre} \end{pmatrix}$$

$$+ \sum_{bins} (R_{bt}^{obs} - R_{bt}^{pre}; R_{mbt}^{obs} - R_{mbt}^{pre}) \cdot W^{-1} \cdot \begin{pmatrix} R_{bt}^{obs} - R_{bt}^{pre} \\ R_{mbt}^{obs} - R_{mbt}^{pre} \end{pmatrix}$$

$$+ \sum_{bins} (R_{bt}^{obs} - R_{bt}^{pre}; R_{mbt}^{obs} - R_{mbt}^{pre}) \cdot W^{-1} \cdot \begin{pmatrix} R_{bt}^{obs} - R_{bt}^{pre} \\ R_{mbt}^{obs} - R_{mbt}^{pre} \end{pmatrix}$$

$$+ \sum_{bins} (R_{bt}^{obs} - R_{bt}^{pre}; R_{mbt}^{obs} - R_{mbt}^{pre}) \cdot W^{-1} \cdot \begin{pmatrix} R_{bt}^{obs} - R_{bt}^{pre} \\ R_{mbt}^{obs} - R_{mbt}^{pre} \end{pmatrix}$$

• Data taking in 3 positions
$$R_{mbt} = middle/\sqrt{bottom \cdot top}$$

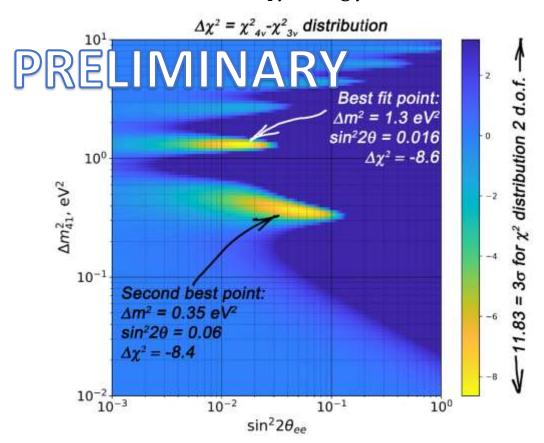
$$W - \text{covariance matrix}$$

$$+\sum_{syst}\frac{(\eta-\eta^0)^2}{\sigma_\eta^2}$$

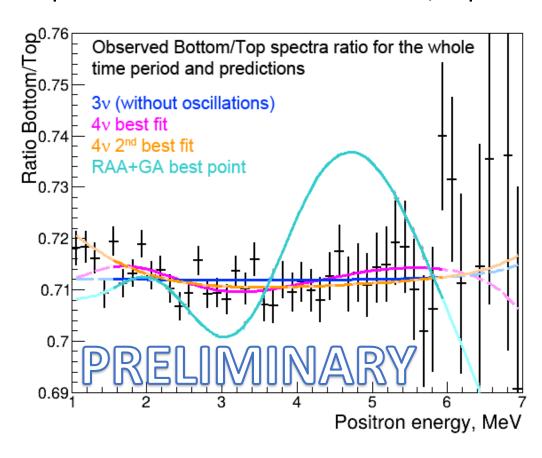
 Penalty term for systematics: relative efficiency, background, energy scale, distance to reactor core

Sterile neutrino (only relative counts)

• Color indicates $\chi^2_{4\nu} - \chi^2_{3\nu}$



predicted and observed Bottom/Top ratio

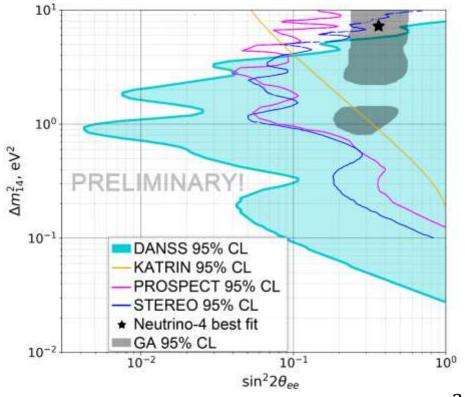


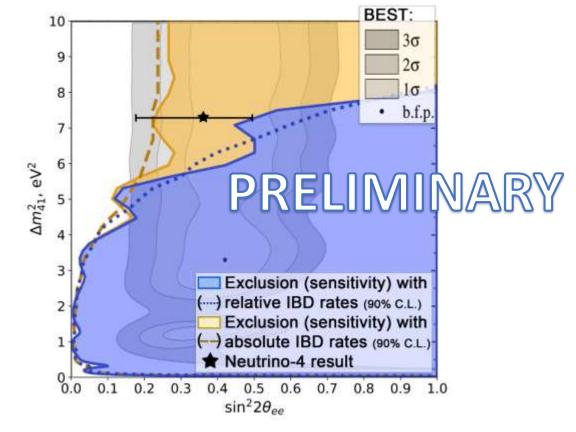
- October 2016 May 2025, 7.2 mln events in $E_{e^+} \in [1.5; 7]$ MeV
- Best point is $\Delta m_{41}^2=1.3~eV^2$; $\sin^2 2\theta_{ee}=0.016$ with significance of 2.5σ only
- No statistically significant evidence for sterile neutrino

Sterile neutrino

$$\chi^{2} += \frac{\left(\left(N_{top} + N_{mid} + N_{bot} \right)^{obs} - \left(N_{top} + k_{2} \cdot \sqrt{k_{1}} \cdot N_{mid} + k_{1} \cdot N_{bot} \right)^{pre} \right)^{2}}{\sigma_{abs}^{2}}$$

- With relative counts independent on model of antineutrino flux
 and on the uncertainty in the detector efficiency
 - With absolute counts Huber-Mueller model and conservative estimation of systematic uncertainty in the absolute rate

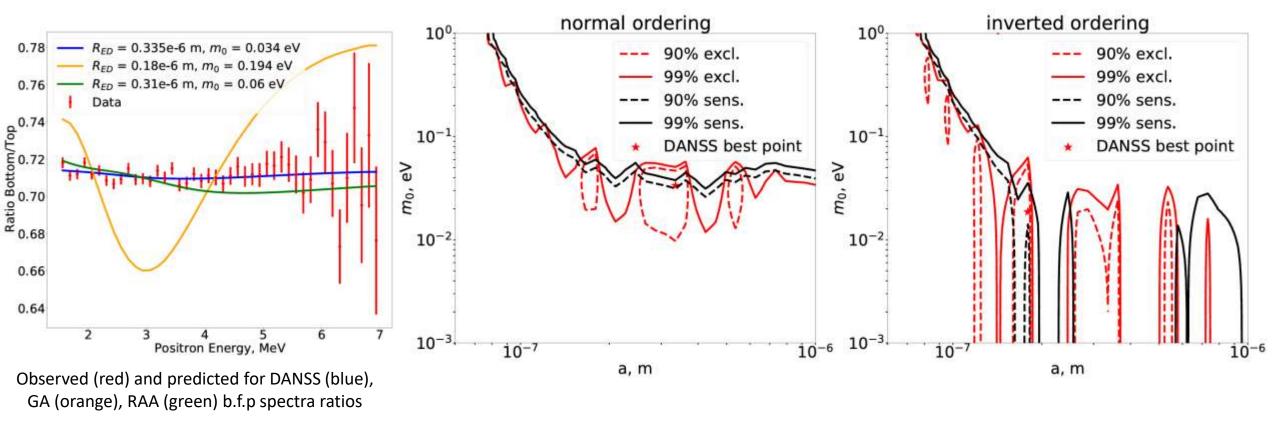




- Absolute counts provide sensitivity to larger Δm^2_{41}
- DANSS exclusions are the best in the world for $\Delta m^2_{41} \in [0.7; 4]~eV^2$
 - Area preferred by RAA and GA is excluded almost completely including b.f.p (much more than 5σ) in the model-independent analysis
 - Neutrino-4 best point is excluded in the analysis with absolute counts

Search for Large Extra Dimensions

I. Alekseev et al., DANSS, Letters to JETP, vol 122, issue 1, page 3, 2025



- Bottom/Top ratio is used
- October 2016 January 2024, 5.8 mln events
- Sensitivity is larger for inverted mass ordering

Search for Large Extra Dimensions

I. Alekseev et al., DANSS, Letters to JETP, vol 122, issue 1, page 3, 2025

• CI with Wilks theorem:

GA data

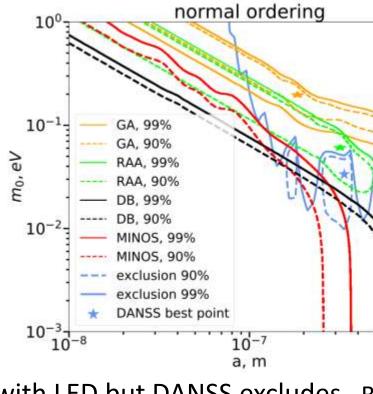
RAA data

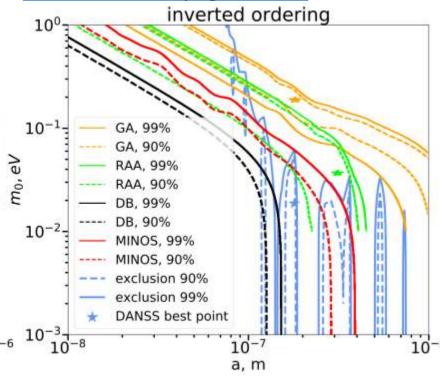
• DB data

• MINOS results

P. Adamson et al, (MINOS), Phys. Rev. D 94, 111101(R), 2016







- GA and RAA can be explained with LED but DANSS excludes corresponding b.f.p. ($> 5\sigma; 4\sigma$) with the most relevant part of parameter space
- Results obtained by DANSS in model-independent analysis contours are the most stringent in the world for some regions with $m_0 \sim 0.01~eV$ and $a \sim 0.1~\mu m$

Best points:

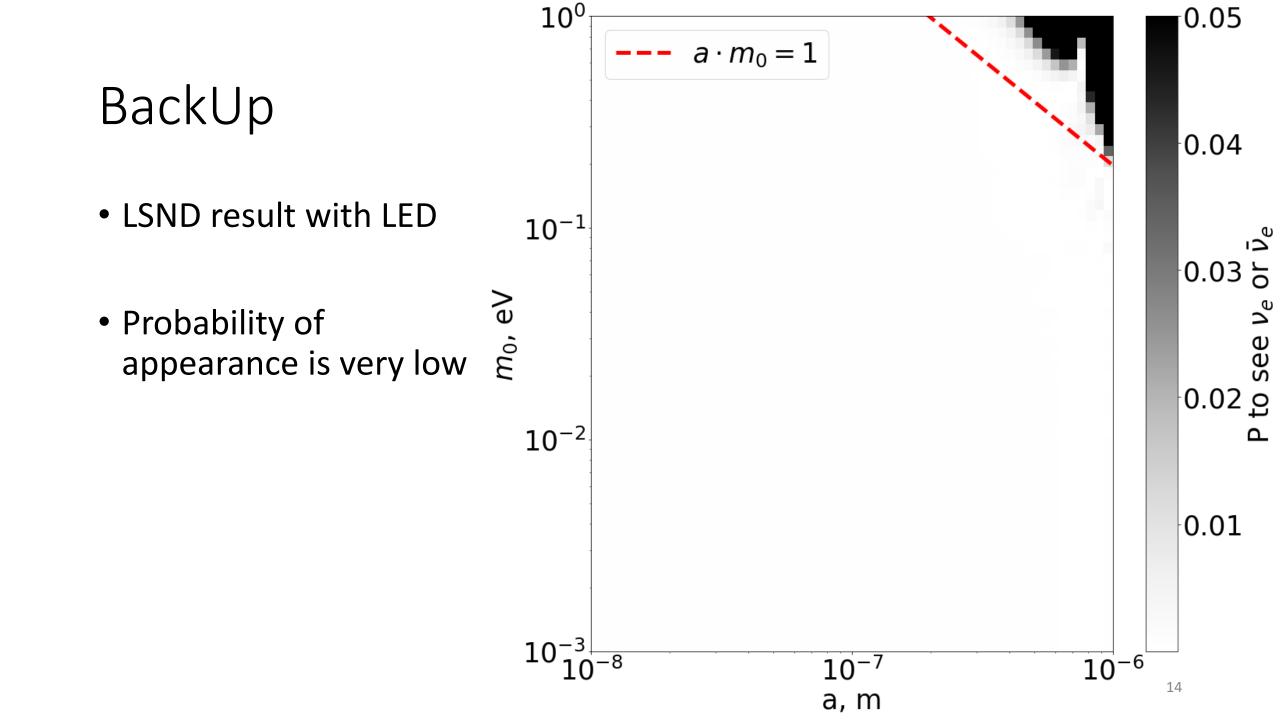
$$a = 3.4 \times 10^{-7} m; m_0 = 0.034 \text{ eV}, 2.0\sigma \text{ (NO)}$$

$$a = 1.8 \times 10^{-7} m; m_0 = 0.017 eV, 1.8\sigma$$
 (IO)

No significant evidence for oscillations to LED

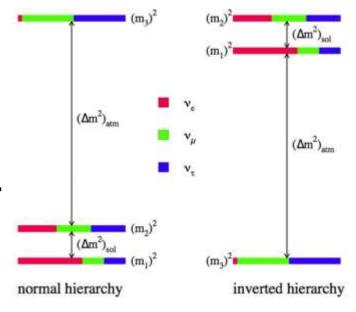
Conclusions

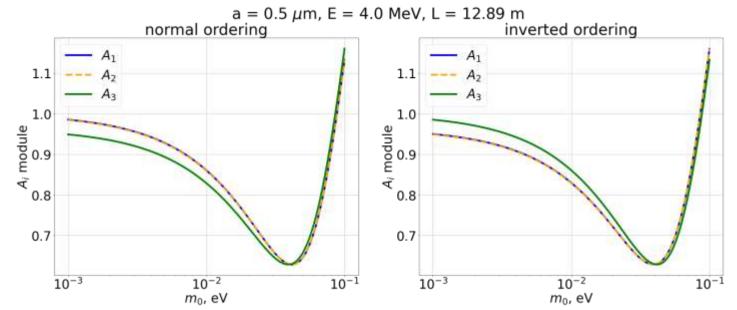
- DANSS acquires around 5000 IBD events per day with S/B ratio > 50
- Measured IBD rate allowed to measure relative reactor power with 1% accuracy in a weak measurements in comparison with traditional methods
- Fission fractions of ^{235}U and ^{239}Pu were determined from the shape of IBD spectrum for the first time with 2.1% precision
- In model-independent analysis DANSS put the most stringent limits on $\sin 2\theta_{ee}$ for $\Delta m_{41}^2 \in [0.7; 4] \ eV^2$
- In the analysis with HM model and conservative estimation of systematic DANSS excludes the best fit point of Neutrino-4
- DANSS excludes the large part of parameter space of LED theory preferred by GA and RAA data and puts the strictest limits for particular regions in (a, m_0) space
- See the report by I. Zhitnikov about upgrade of the DANSS



Mass ordering

• For the same m_0 in different ordering values of m_1, m_2, m_3 differ, hence, amplitudes are also different. They have different weights in the total probability





Survival probability:

$$P_{ee} = \left| \sum_{i=1}^{n} |U_{ei}|^2 A_i \right|^2$$

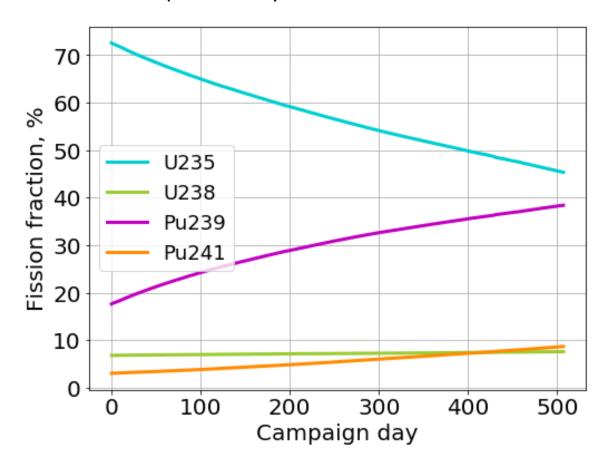
$$|U_{e1}| \sim 0.8$$

$$|U_{e2}| \sim 0.5$$

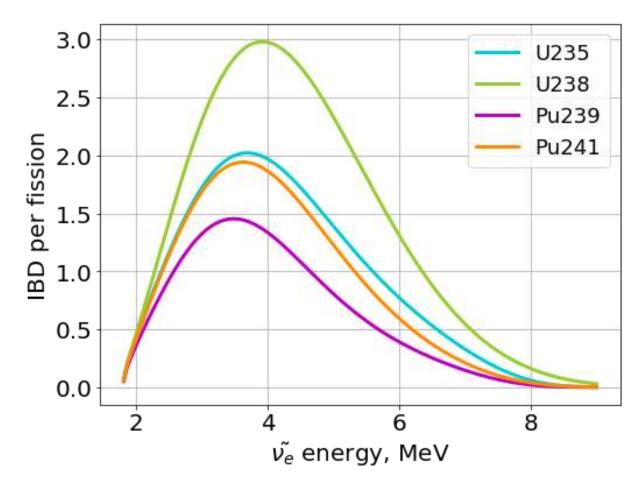
$$|U_{e3}| \sim 0.1$$

Fission fractions

Fission fraction provided by KNPP



Antineutrino spectra for different isotopes



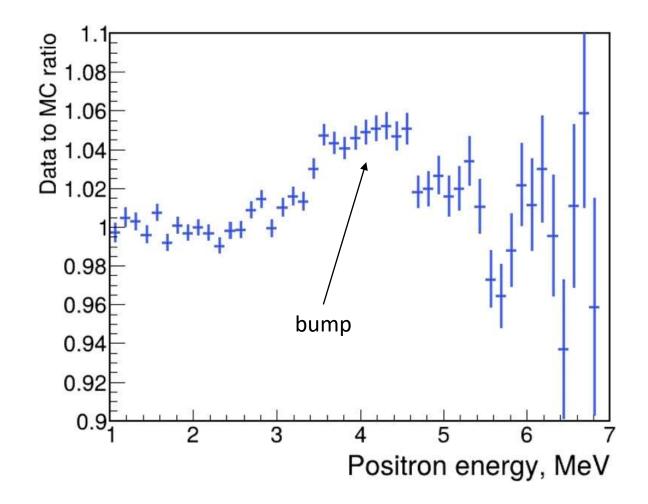
List of systematics

- η_0 nominal values
- σ_η standard deviations

η	k	b	K_E	smear	shift, кэВ	move, cm
η_0	1	0	1	0	0	0
σ_{η}	$4 \cdot 10^{-3}$	$7 \cdot 10^{-3}$	0.02	$\frac{6\%}{\sqrt{E}} + 2\%$	50	5

Bump in positron spectra

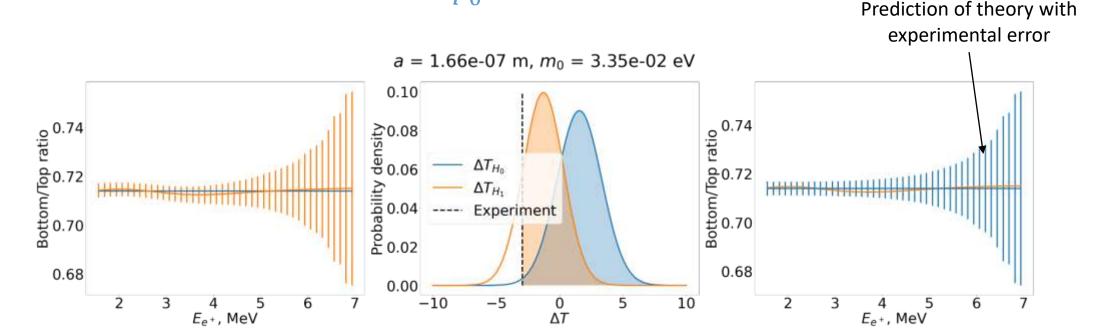
- Ratio of measured and predicted by HM spectra
- 3-5.5 MeV range is not included



Gaussian CL_s method

Significance of exclusion H_1 in comparison with H_0 without any suggestions on the structure of parameter space. H_0 is assumed to be true.

- $\chi_1^2 \chi_0^2$ (ΔT) is distributed according to normal distribution with parameters determined from Asimov datasets.
- Confidence level: $CL_s = \frac{1-p_1}{1-p_0}$



Unfolding of the antineutrino spectrum

Unfolding with SVD method

