

# Status of Dark Matter Direct Detection Experiments

---

**Speaker : Qing Lin**

University of Science and Technology of China

Lomonosov Conference @ Moscow, 2025.08.22

# CONTENTS

---



**01**

**DM direct search**

**02**

**>10GeV Searches**

**03**

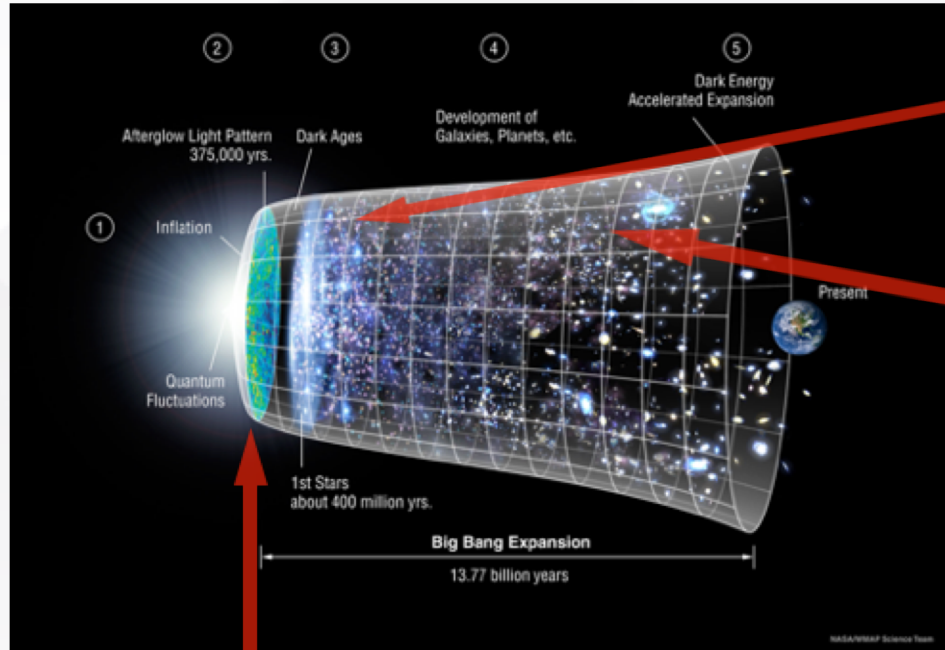
**~GeV Searches**

**\*\*Disclaimer: Apology for omittance!\*\***

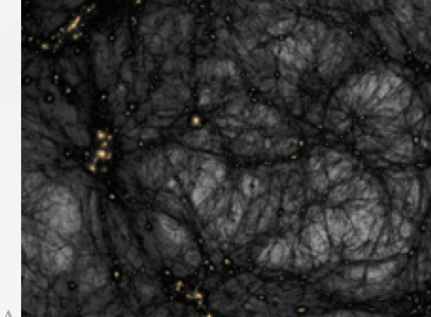
**\*\* Most materials are borrowed from public talks, especially from Prof J. Liu, Prof. Q. Yue's, and talks from IDM2024.\*\***



# Dark Matter and its Gravitational Evidence



Large Structure



Courtesy of NASA

Credit: <https://youtu.be/sI23cwbbNqs>

Small Structure

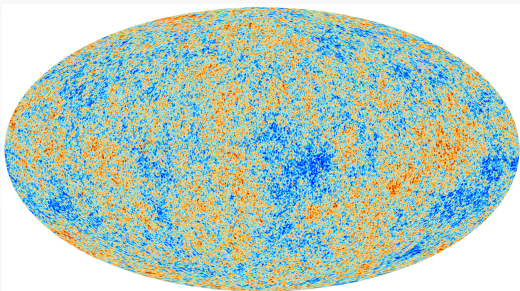


Bullet cluster collision

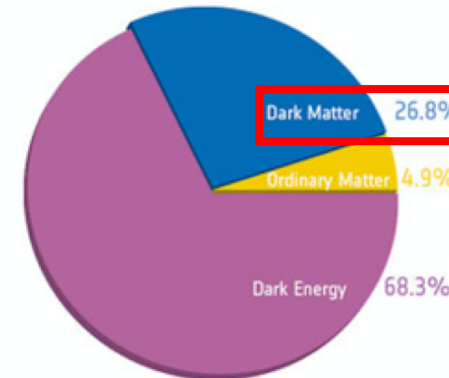
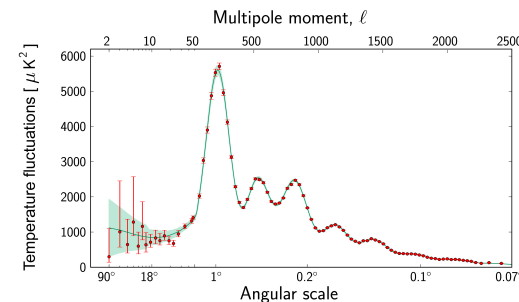


Galaxy rotation curve

Primordial Universe

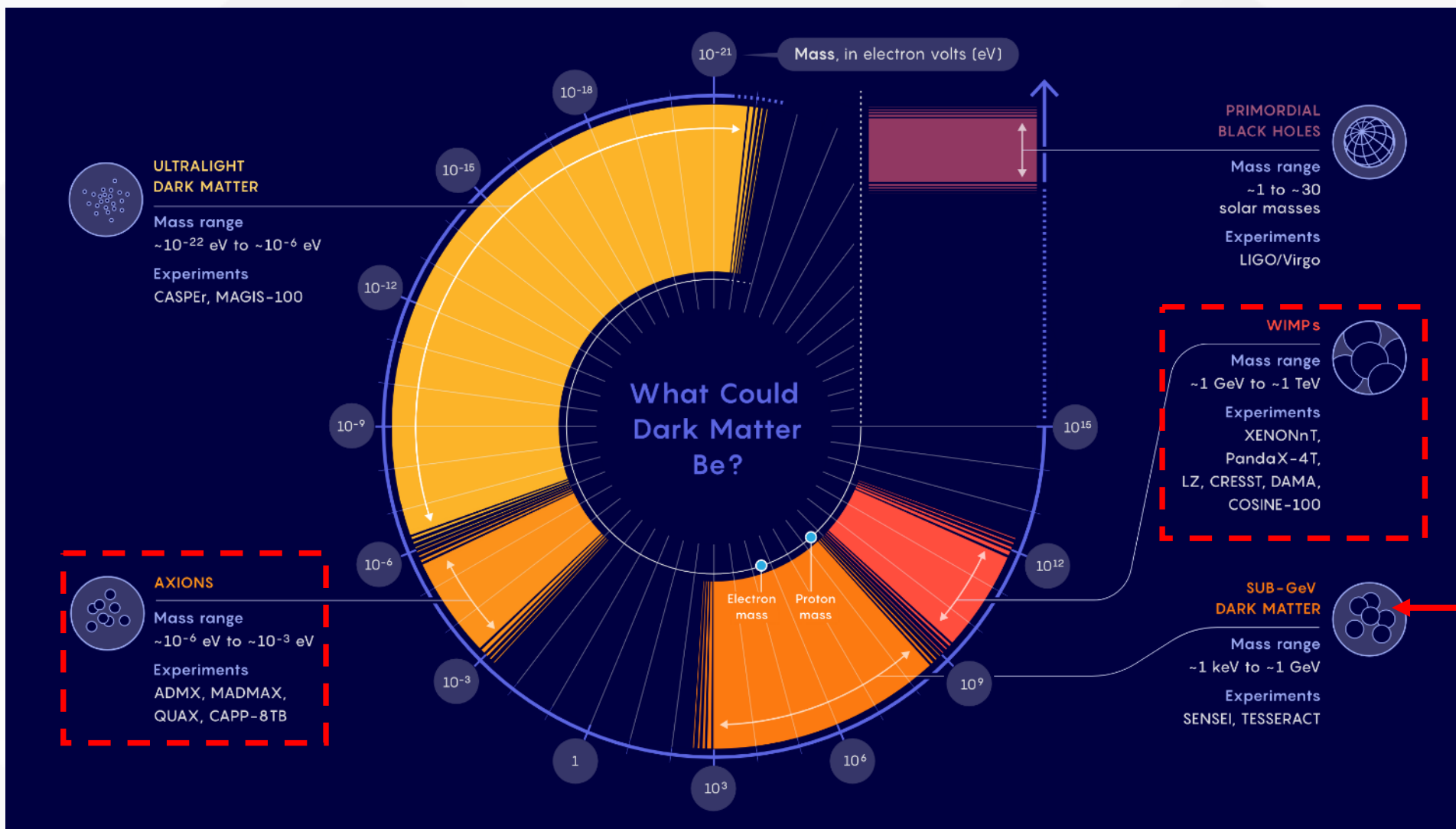


Courtesy of PLANCK

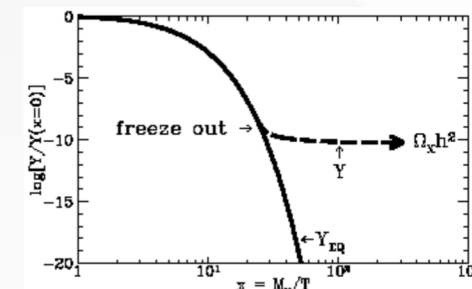


Gravitational evidences suggest **dark matter** is the **dominant form of matter** in Universe!

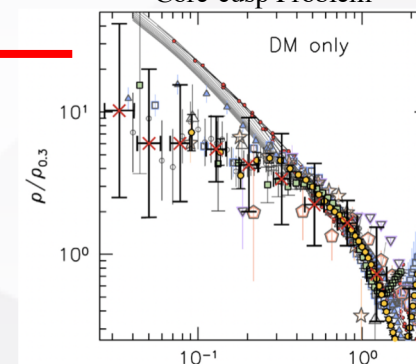
# Many Possibilities



**This talk!**



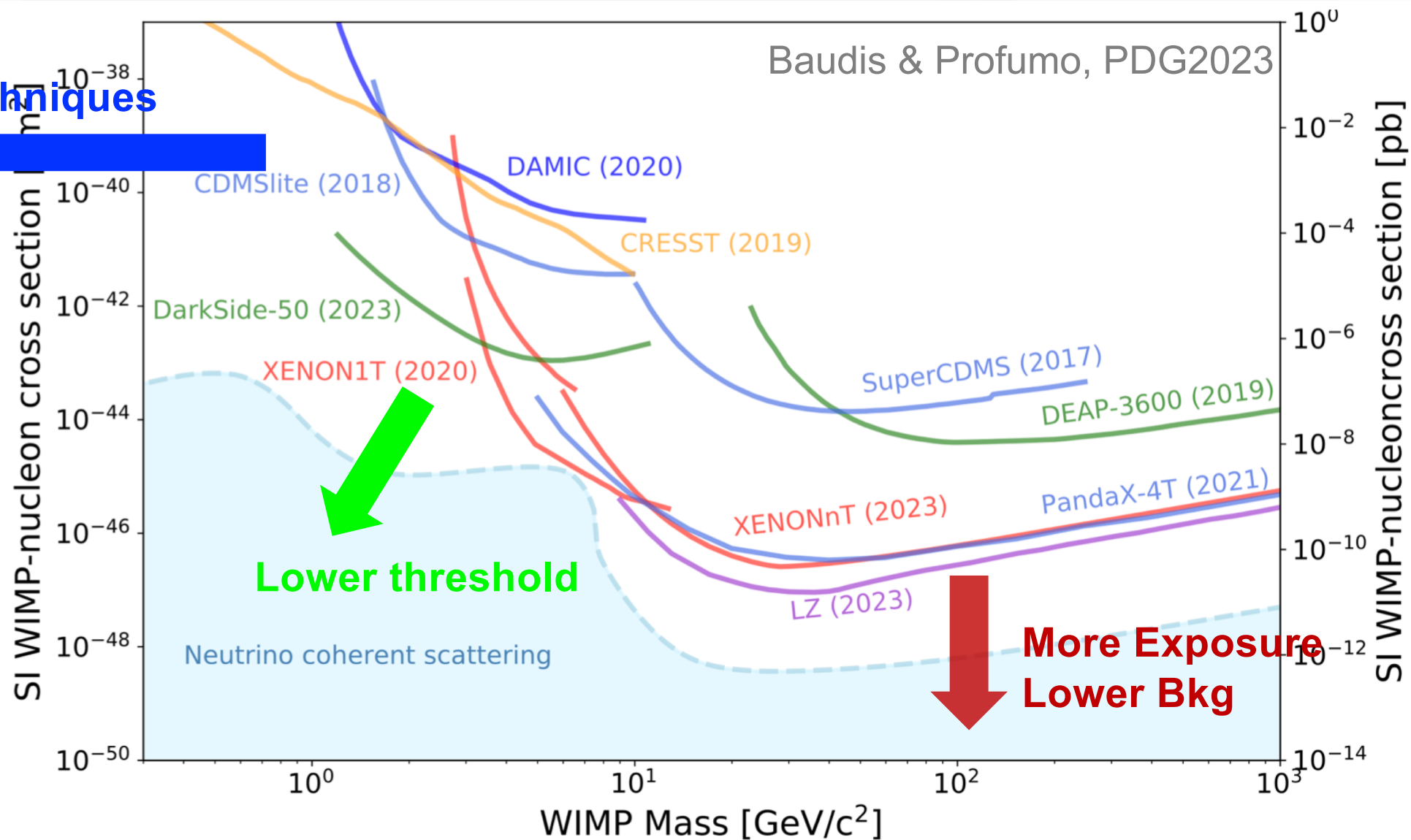
Core-cusp Problem



# Direct Search Overview

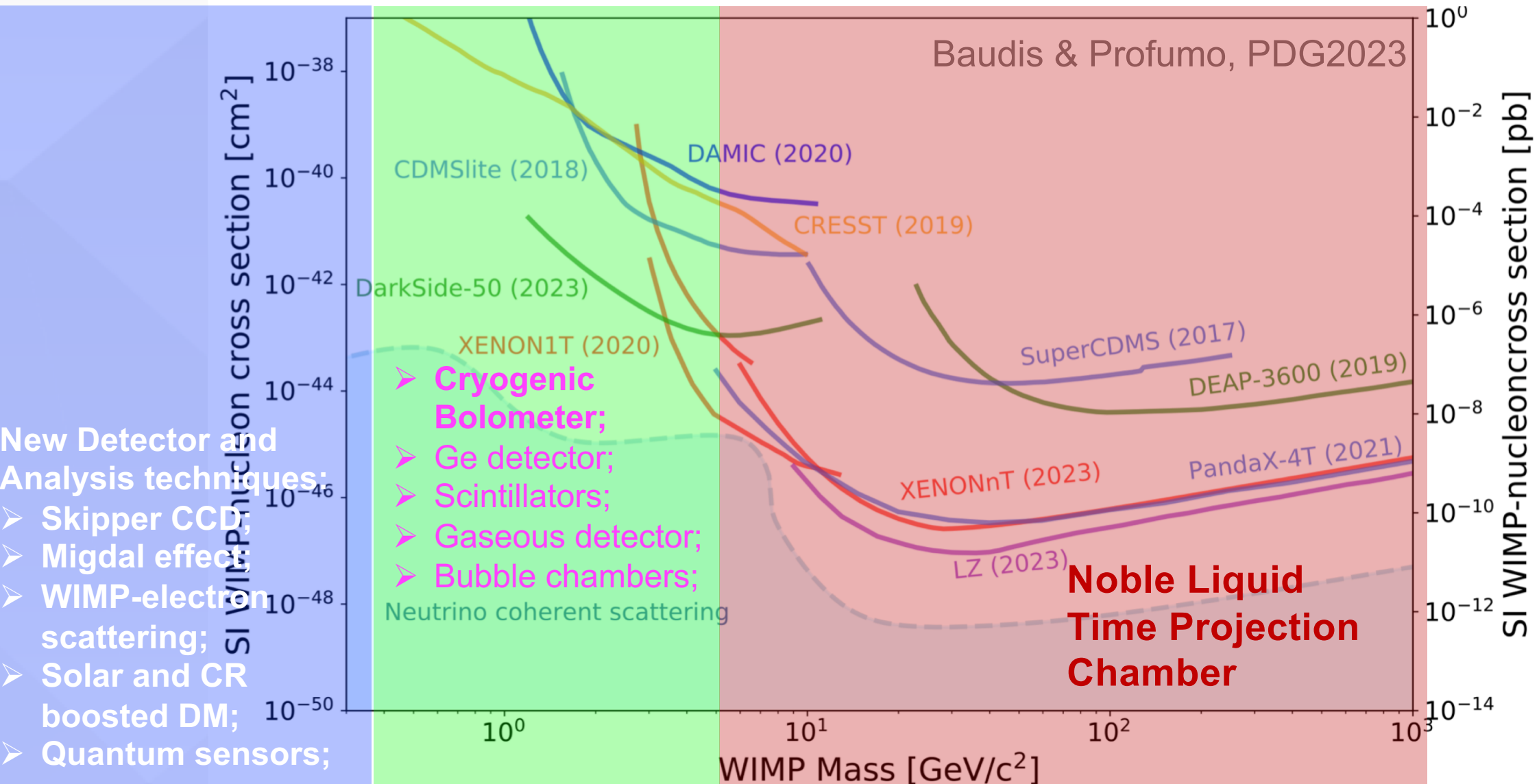


New Techniques





# Direct Search Overview



# CONTENTS

---



**01**

**DM direct search**

**02**

**>10GeV Searches**

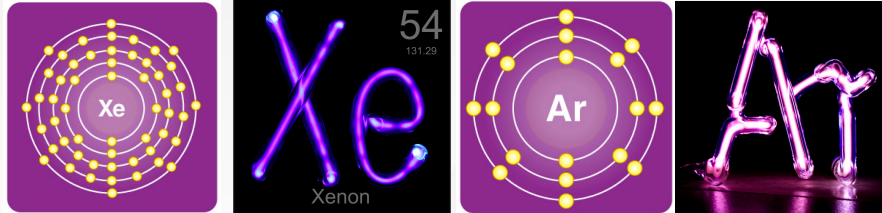
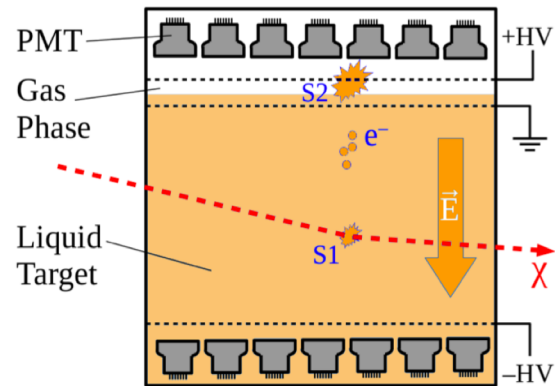
**03**

**~GeV Searches**

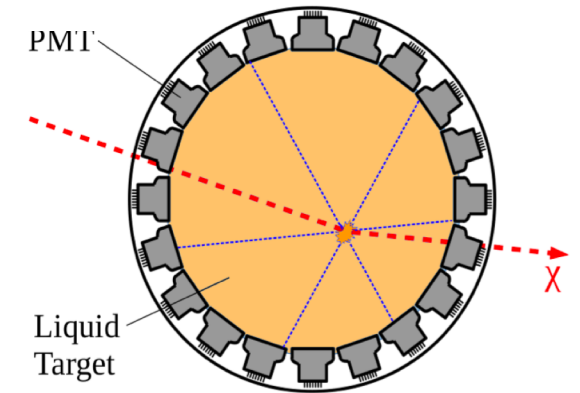
**\*\*Disclaimer: Apology for omittance!\*\***

**\*\* Most materials are borrowed from public talks, especially from Prof J. Liu, Prof. Q. Yue's, and talks from IDM2024.\*\***

# Noble Liquid



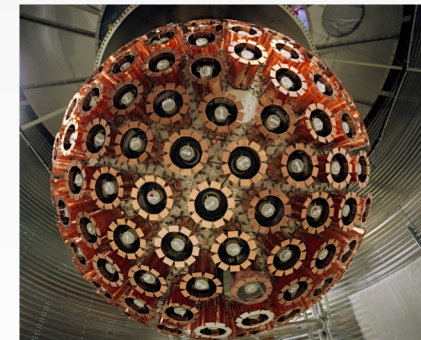
Scalability	Good	Good
Density	3 kg/L	1.4 kg/L
A	133	40
Long-lived bkg	Xe136 Negligible	Ar39 (1 Bq/kg in nat Ar, $T_{\text{half}}=269\text{yr}$ )
Bkg rej.	S2/S1, 3D pos.	S2/S1, Pulse shape, 3D pos.
Scintillation light	178 nm	128 nm (need WLS)
Cost	¥15000/kg	¥20/kg for nat Ar



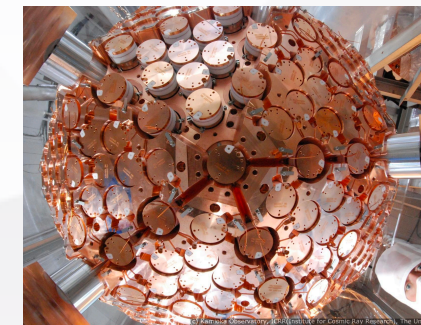
DarkSide-50  
(DarkSide-20k  
ARGO)



PandaX-4T  
XENONnT  
LZ  
(XLZD  
PandaX-xT)



DEAP-3600

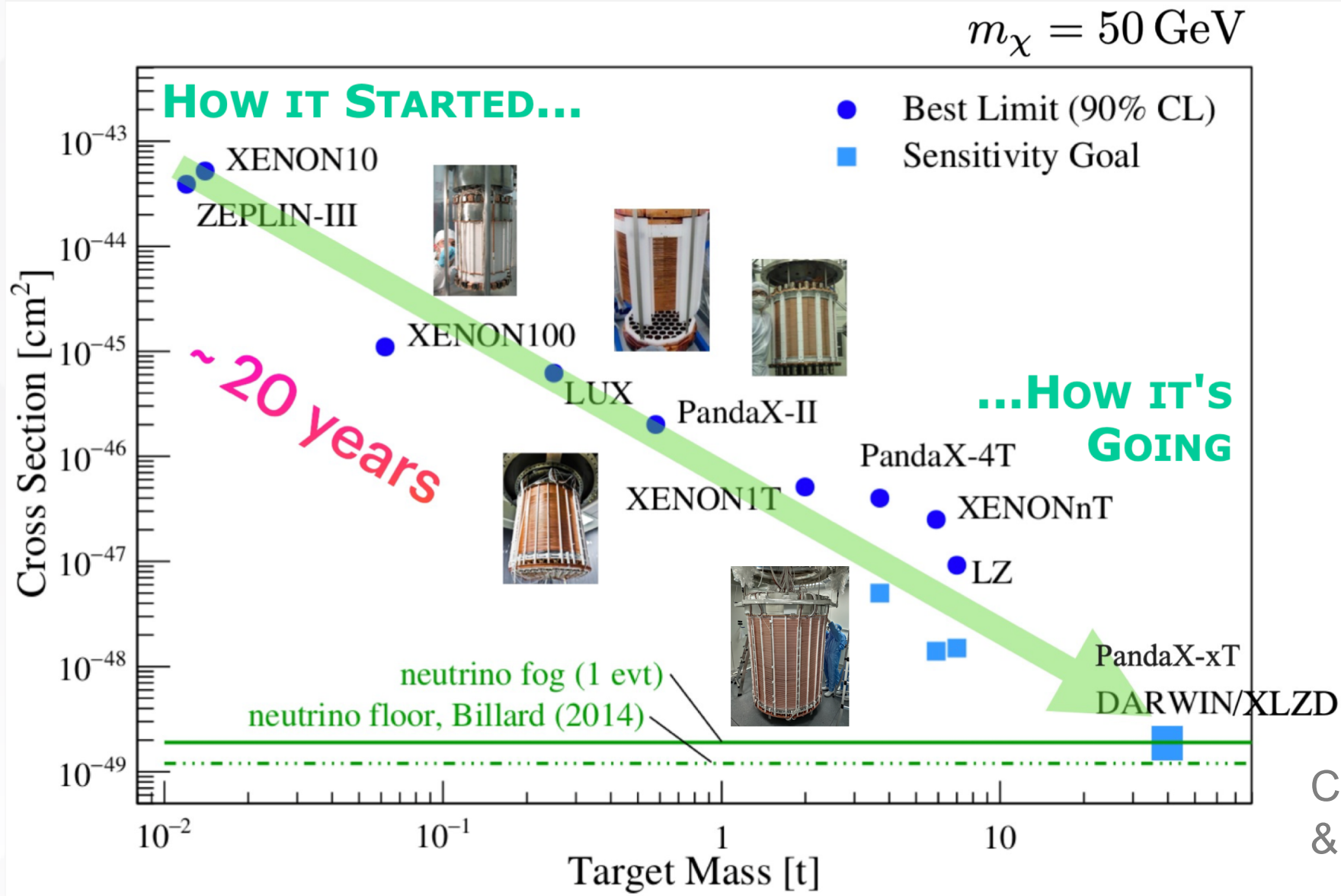


XMASS

- ❑ High sensitivity to  $>10\text{GeV}$  DM due to its good scalability and bkg discrimination;
- ❑ Liquid xenon dual-phase TPC currently has the world-leading sensitivity to  $>10\text{GeV}$  DM.



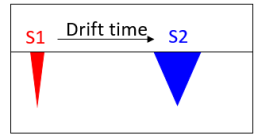
# Evolution of Noble Liquid Sensitivities



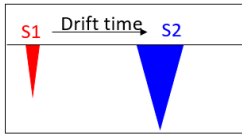
Courtesy to P. Ullio  
& M.E. Monzani

# Liquid Xenon Dual phase Time Projection Chamber

Dark matter: nuclear recoil (NR)

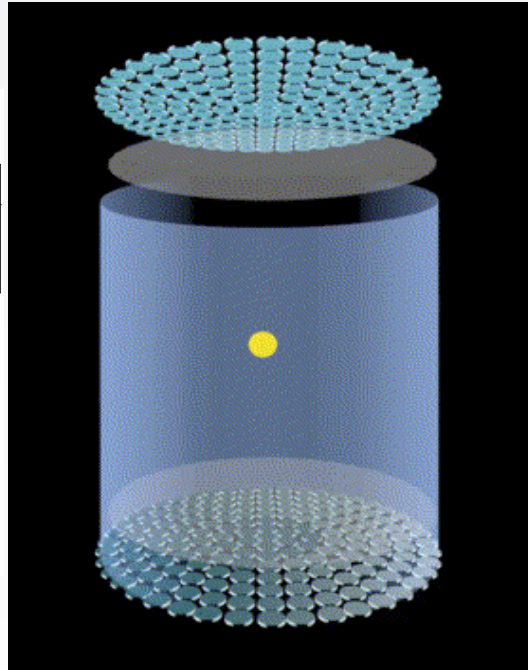
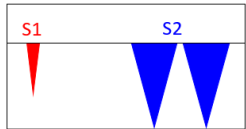


$\gamma$  background: electron recoil (ER)

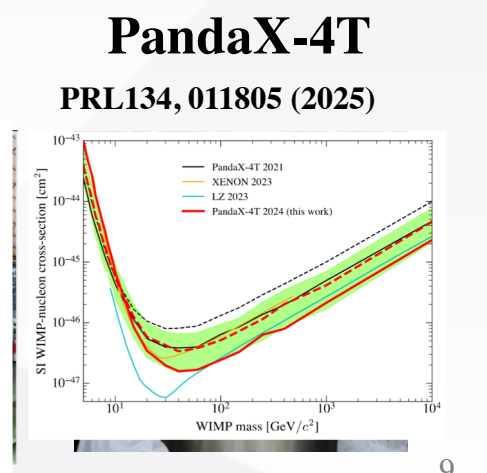
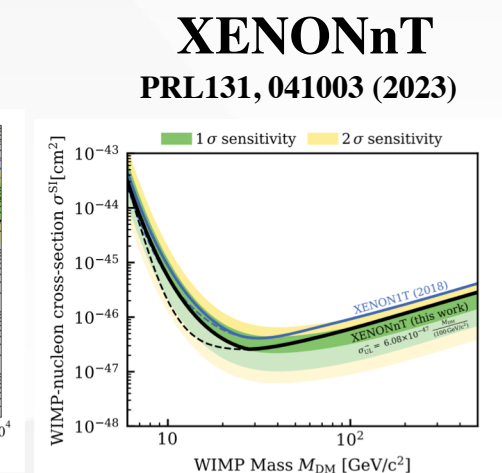
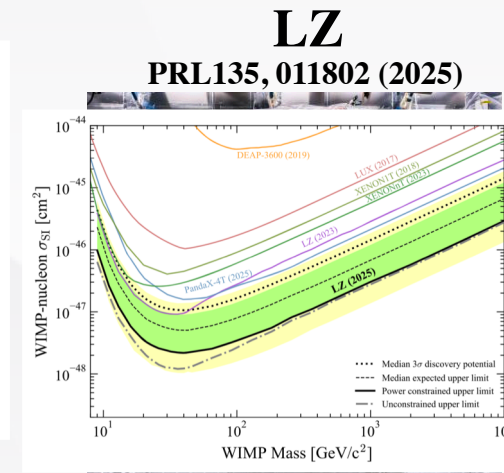
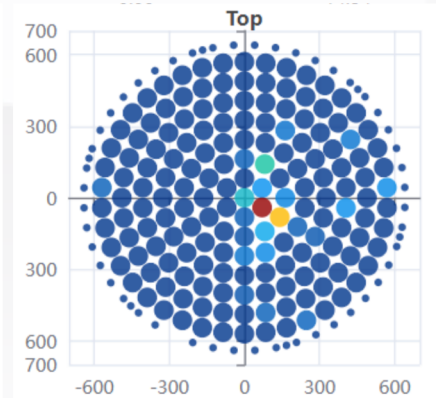
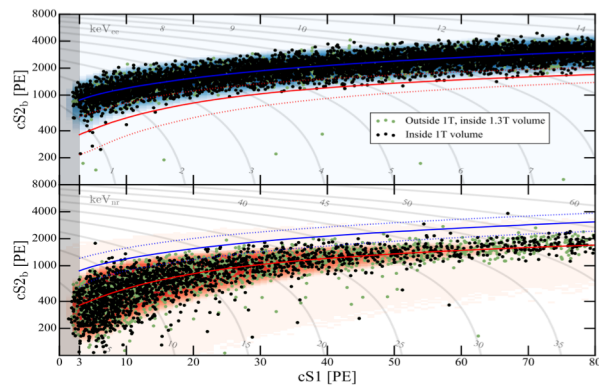
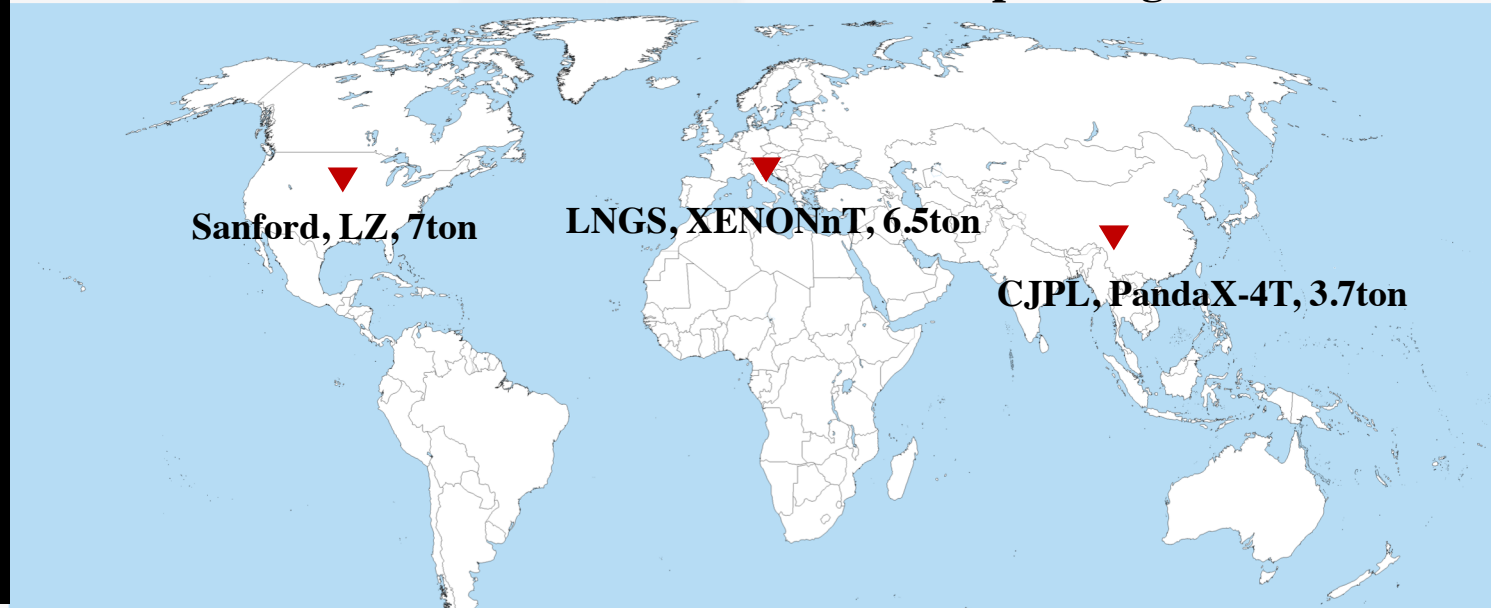


$$(S2/S1)_{NR} \ll (S2/S1)_{ER}$$

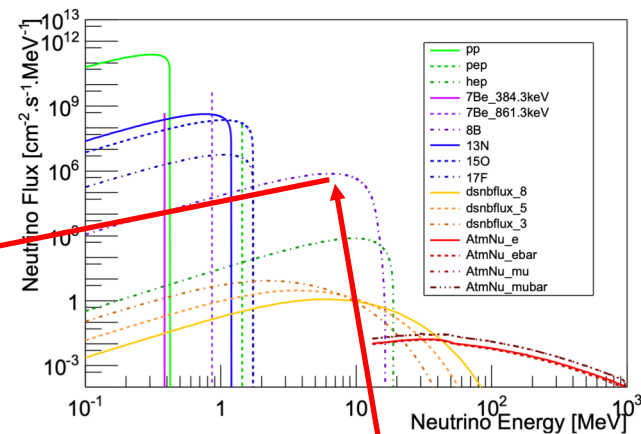
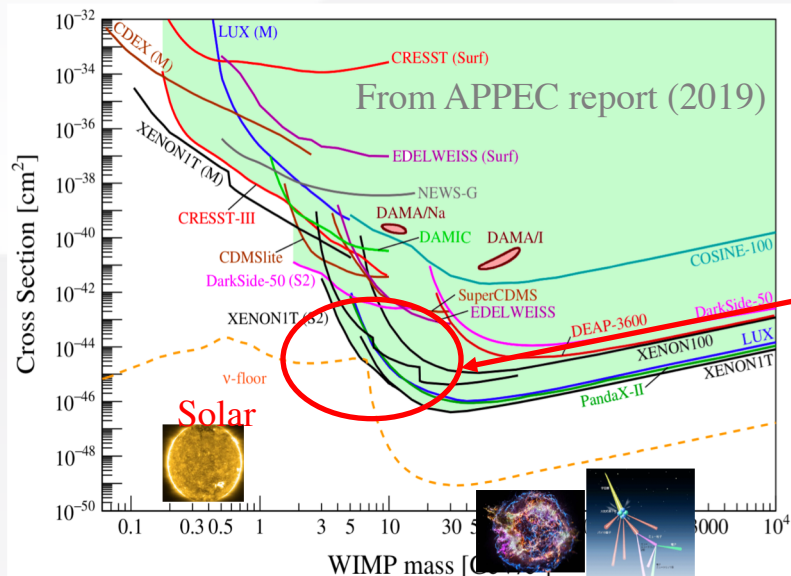
Multi-site scattering background (ER or NR)



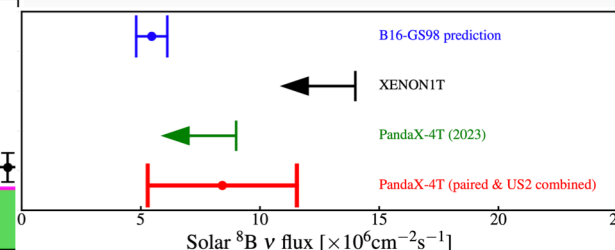
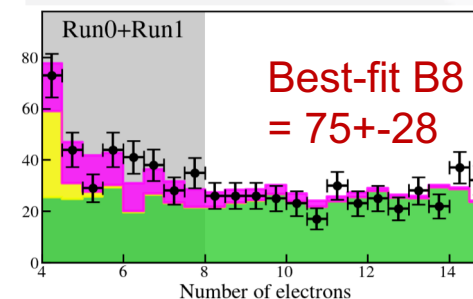
Three multi-tonne detectors operating.



# Touching neutrino fog



**PandaX-4T (1.05 t-yr; 2.64 $\sigma$ )**

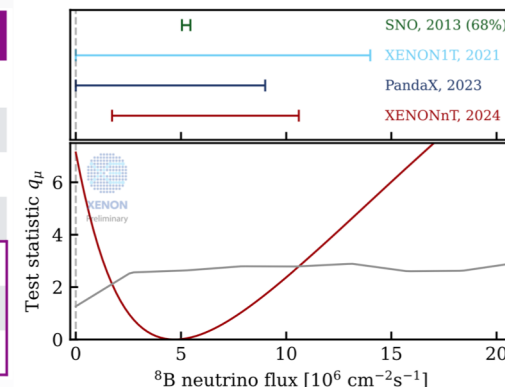


Q. Lin, IDM2024

PRL 133, 191001 (2024)

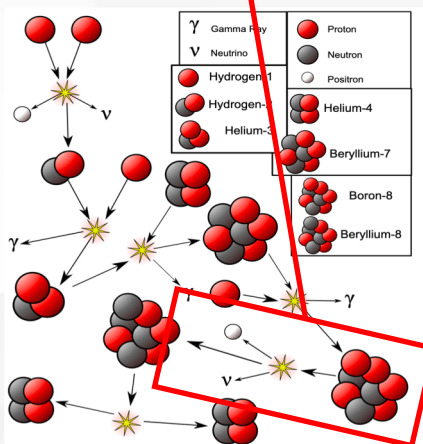
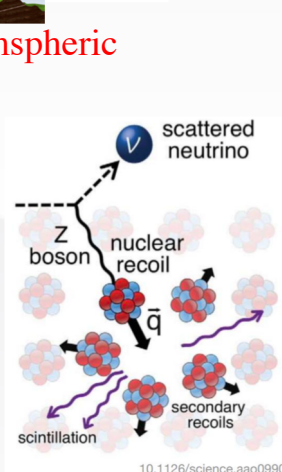
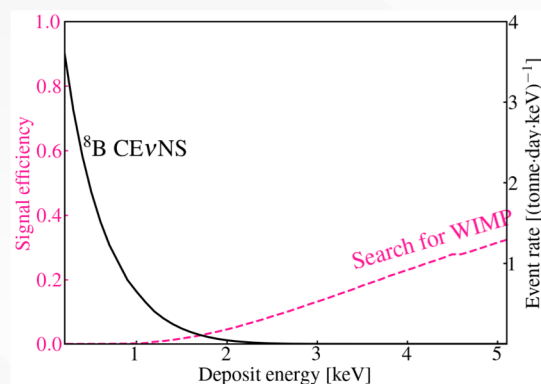
**XENONnT (3.51 t-yr; 2.73 $\sigma$ )**

Component	Background only fit	Background + $^8\text{B}$ fit	Nominal Expectation
AC - SR0	7.55	7.36	$7.48 \pm 0.52$
AC - SR1	18.26	17.90	$17.77 \pm 1.23$
ER	0.74	0.54	$0.68 \pm 0.68$
NR	0.50	0.45	$0.47 \pm 0.32$
Total Background	27.05	26.24	$26.4 \pm 1.5$
$^8\text{B}$	-	10.71	$11.9 \pm 3.1$
Observed		37	



F. Gao, IDM2024

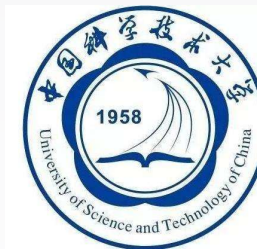
PRL 133, 191002 (2024)



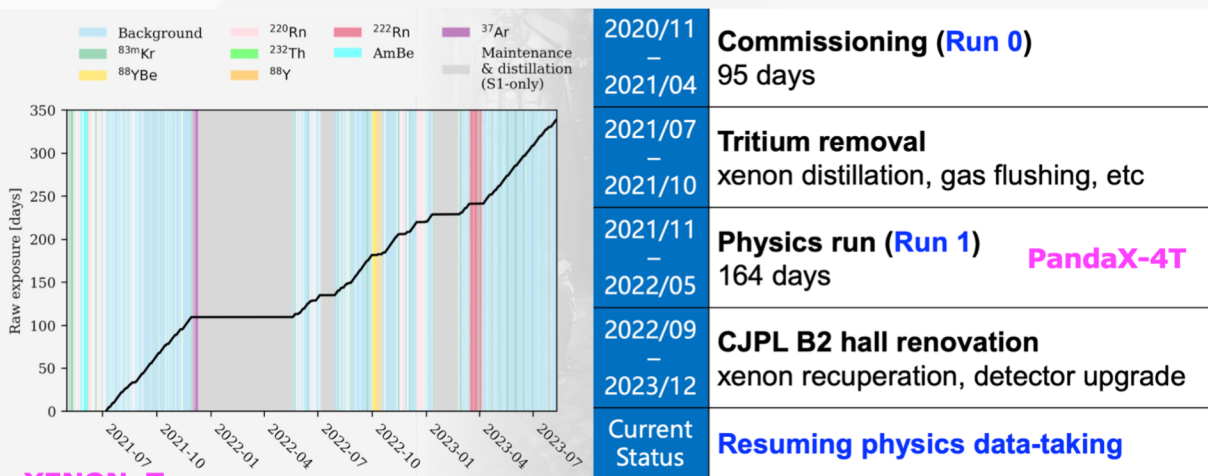
Lower the threshold!!!



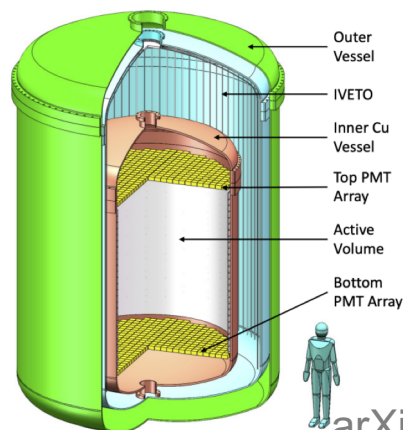
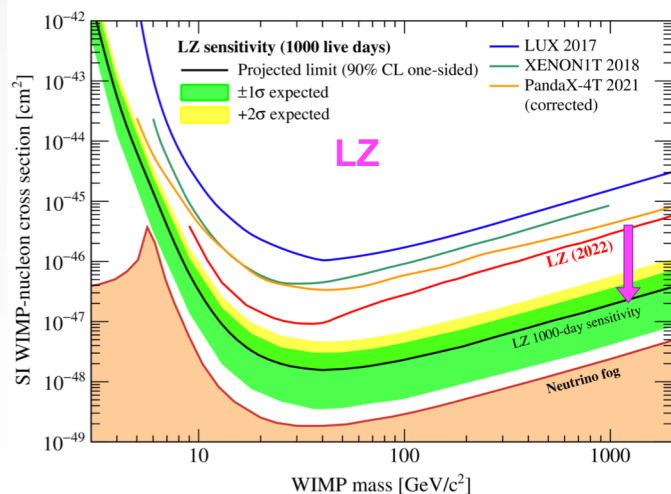
# Future LXeTPC



## 2<sup>nd</sup>-generation LXe-TPCs continue taking data!



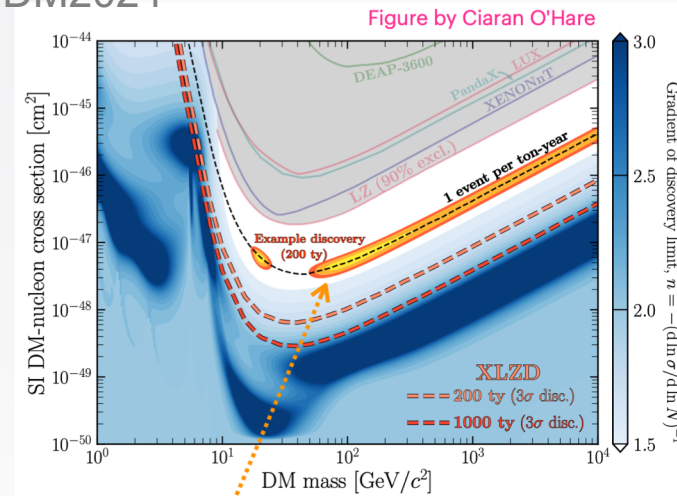
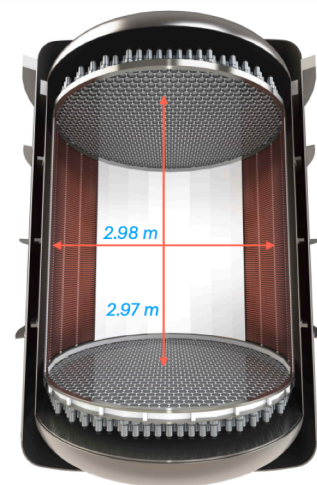
Courtesy to M. E. Monzani, IDM2024



arXiv: 2402.03596

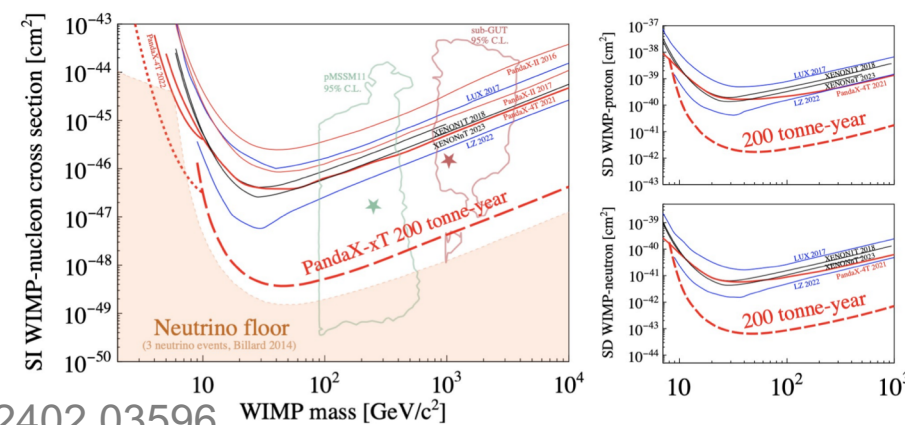
## XLZD (XENON-LZ-DARWIN); > 60t

Courtesy to L. Baudis, IDM2024



Confidence intervals for 200 tonne-yr (1-, 2-, 3- $\sigma$ )

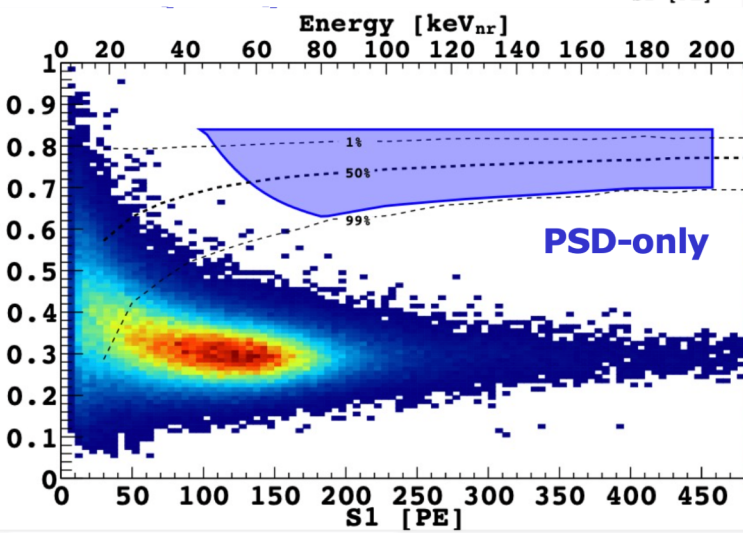
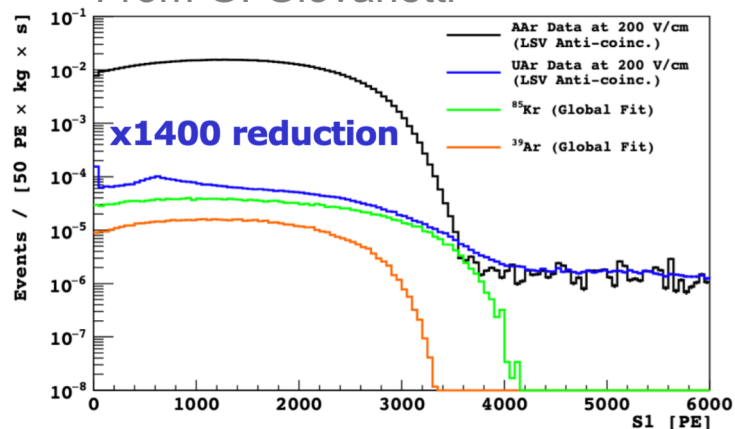
## PandaX-xT; ~43 tonne



# LAr-TPC work force



From G. Giovanetti



**Higher intrinsic bkg with natural Ar,  
but with PSD discrimination!**

Global Argon Dark  
Matter Collaboration

MiniCLEAN  
SNOLAB



ArDM  
Canfranc

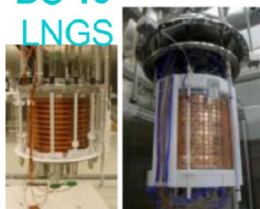


DEAP-3600  
SNOLAB



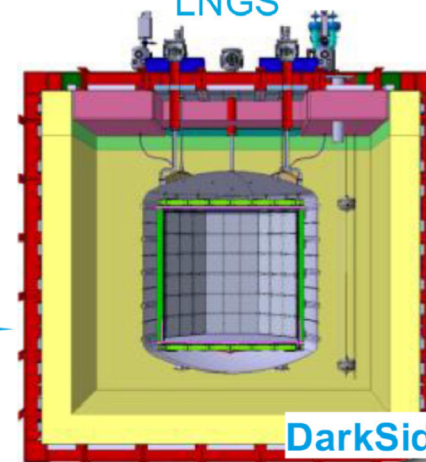
DarkSide50

DS-10  
LNGS



From M. E. Monzani, IDM2024

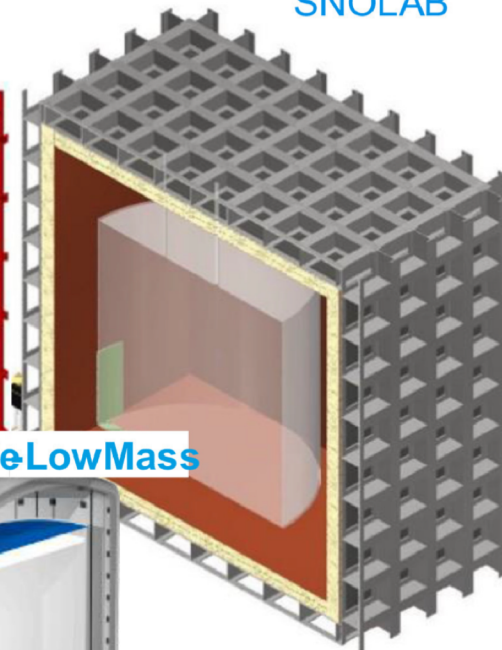
DarkSide20k  
LNGS



DarkSideLowMass



Argo  
SNOLAB



2011

2013

2016

2026

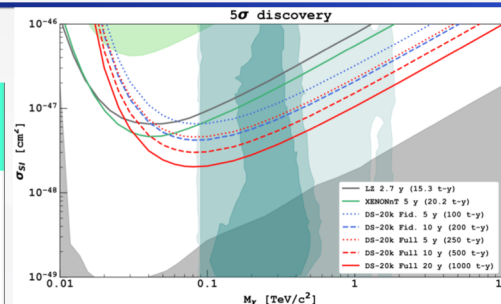
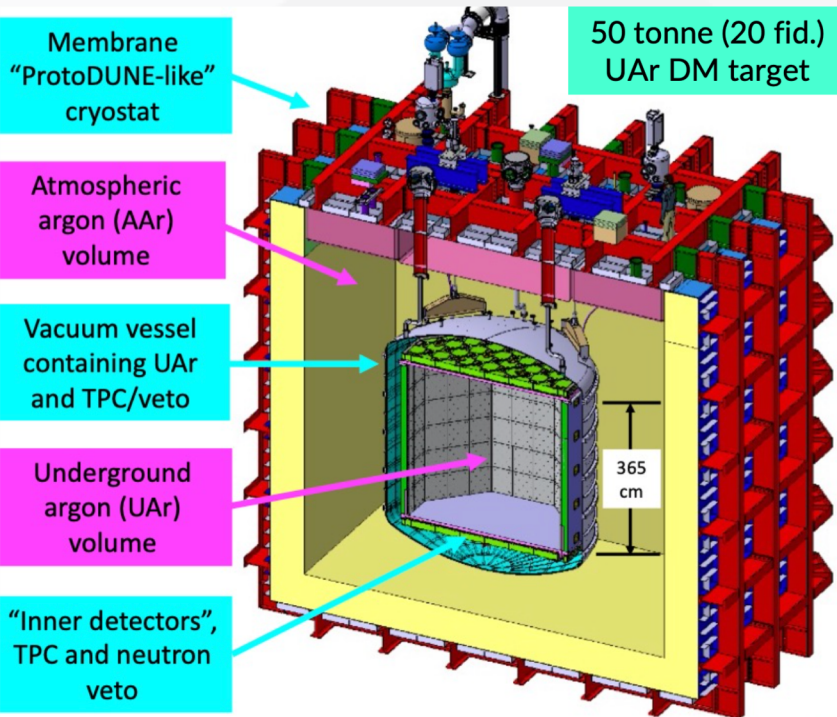
2030+



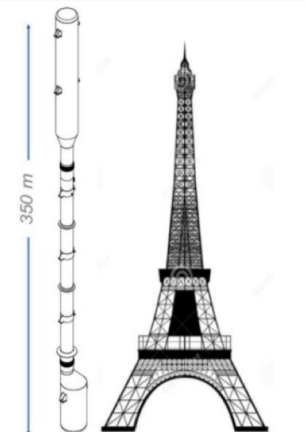
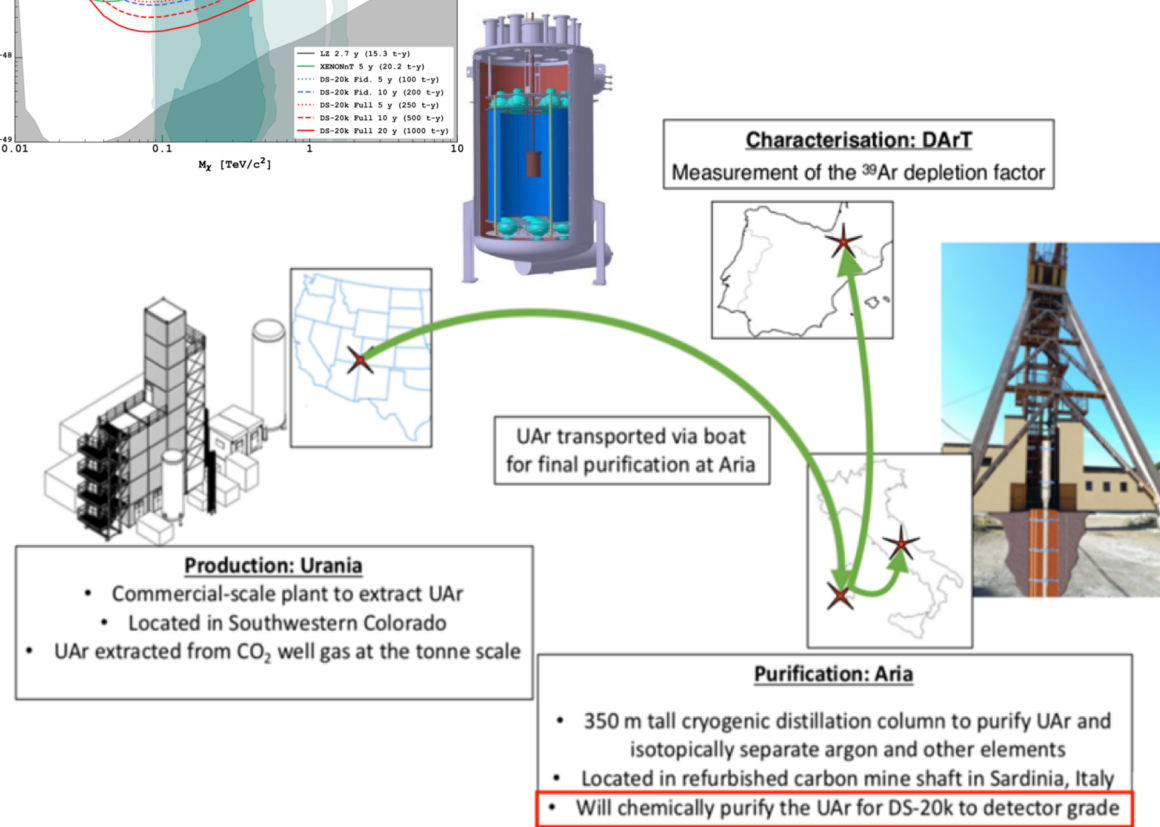
# DarkSide-20k



From M.E. Montani, IDM2024



From Roberto Santorelli - La Thuile 2022



- ☐ Under construction,;
- ☐ 50-tonne total mass;
- ☐ 20t fiducial mass for 10-year exposure;
- ☐ SiPM 4pi coverage;

Courtesy to J. Monroe, IDM2024

- ☐ UAr mined from Urania;
- ☐ Further distilled at Aria facility;



# CONTENTS

---



**01**

**DM direct search**

**02**

**>10GeV Searches**

**03**

**~GeV Searches**

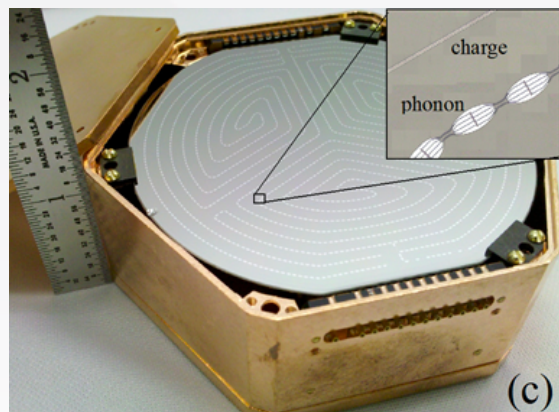
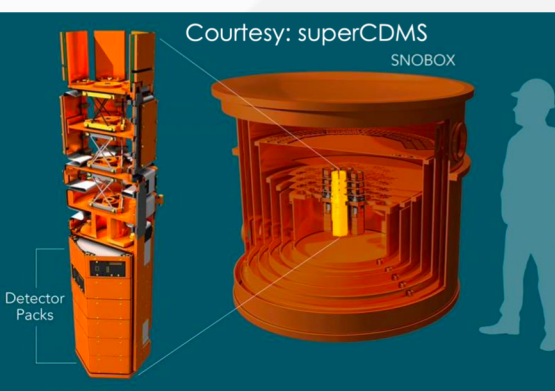
**\*\*Disclaimer: Apology for omittance!\*\***

**\*\* Most materials are borrowed from public talks, especially from Prof J. Liu, Prof. Q. Yue's, and talks from IDM2024.\*\***

# Ultra-low temperature: Bolometer

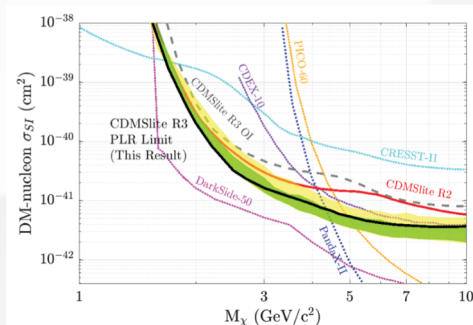
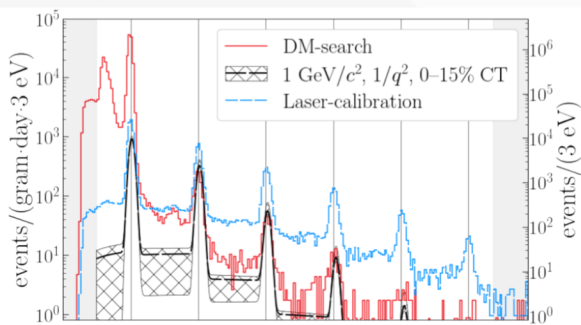


## SuperCDMS @ SNOLAB



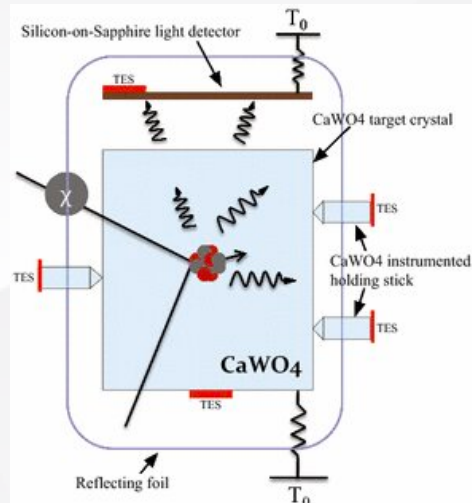
PHYSICAL REVIEW LETTERS 121, 051301 (2018)

Phys. Rev. D 102, 091101 (2020) PHYS. REV. D 99, 062001 (2019)

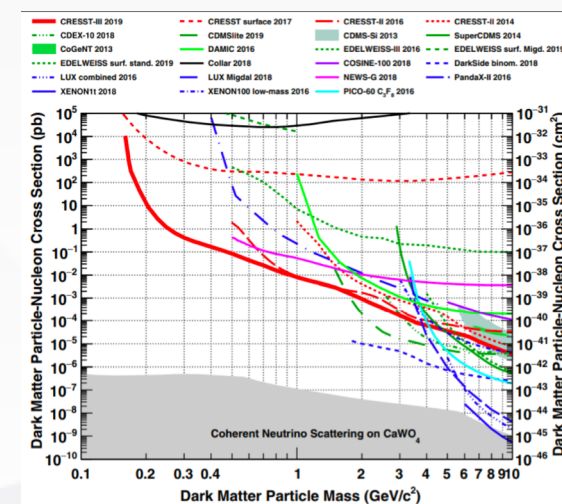
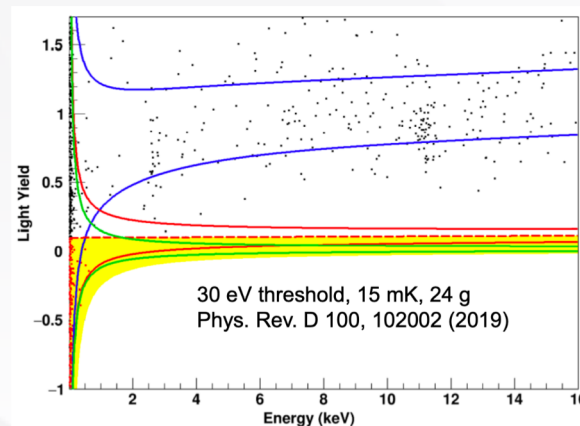
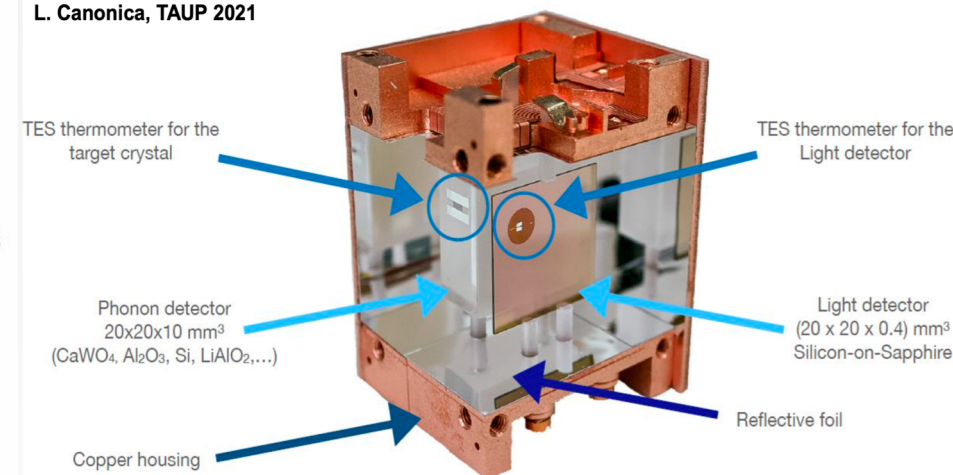


- Super-high resolution, 3% to single e-h pair;
- High sensitivity to GeV WIMP.
- **Current status: Detector under construction**

## M. Mancuso, Journal of Low Temperature Physics 199, 547 (2022) CRESST @ LNGS



L. Canonica, TAUP 2021

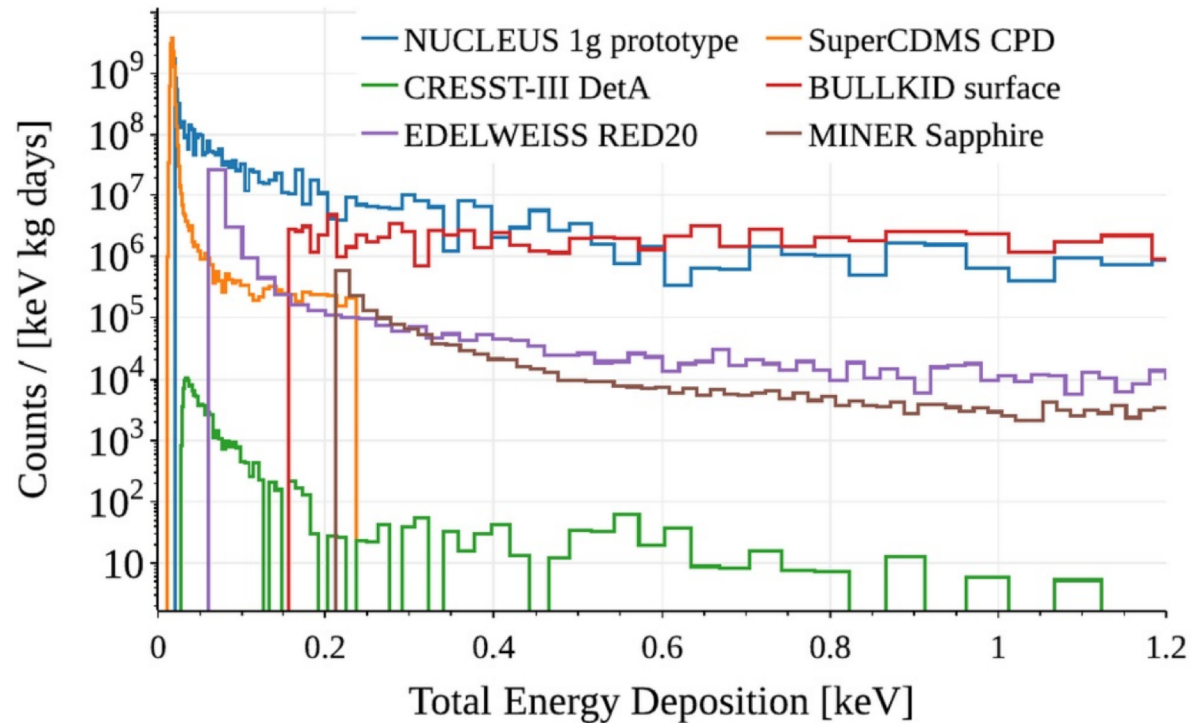


- Great background rejection power.
- **Status: Upgrade detector and readout**

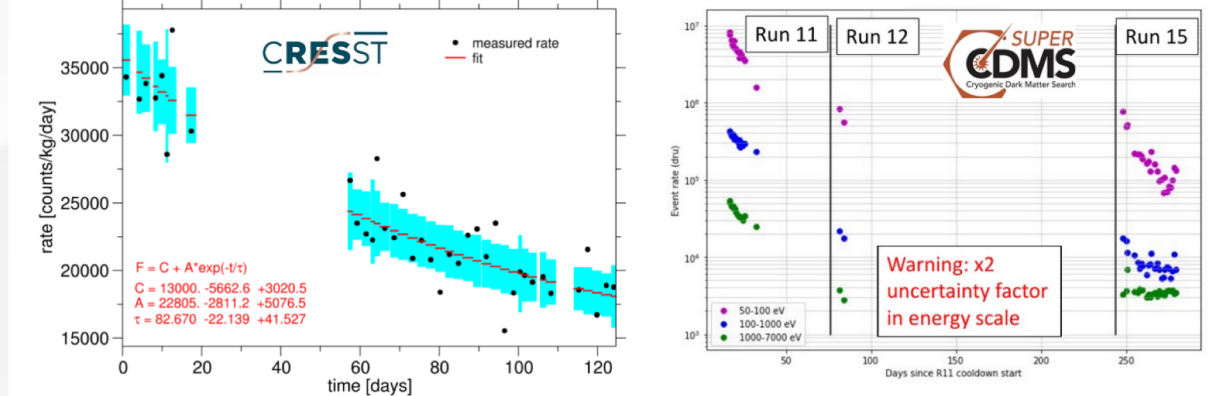
# Low energy excess in Cryogenic Crystal Detectors

Distinct excess rates among experiments!

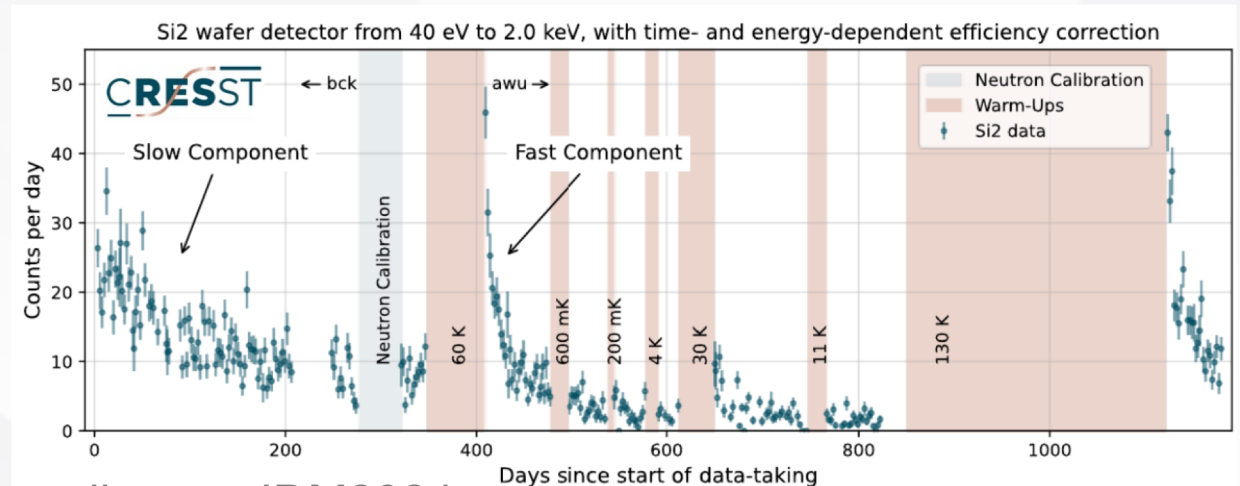
NOT DM since it decreases over time!



<https://github.com/fewagner/excess>



Also activated with thermal cycles!

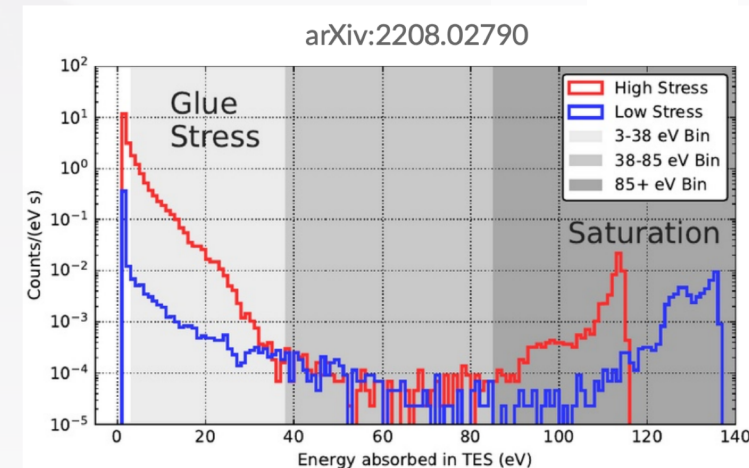
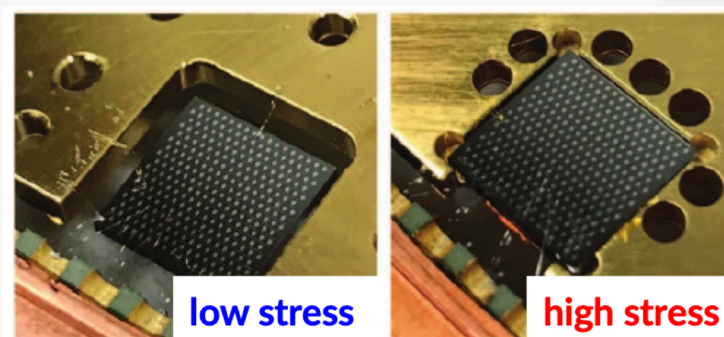
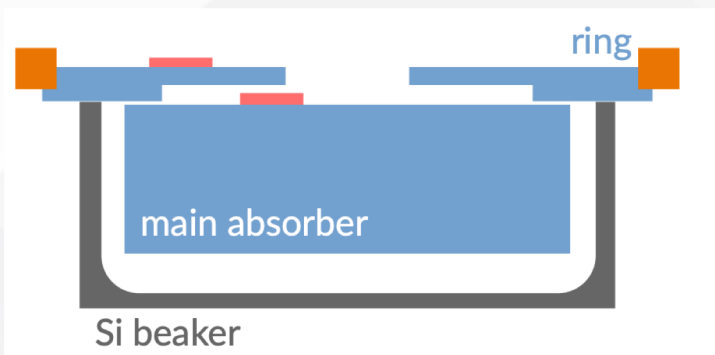




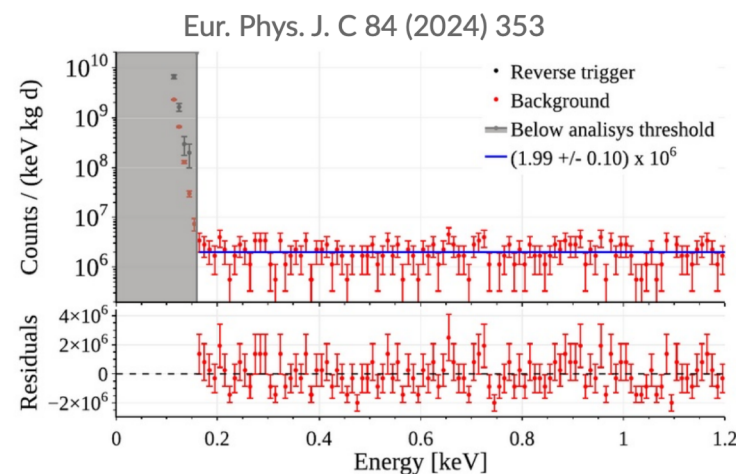
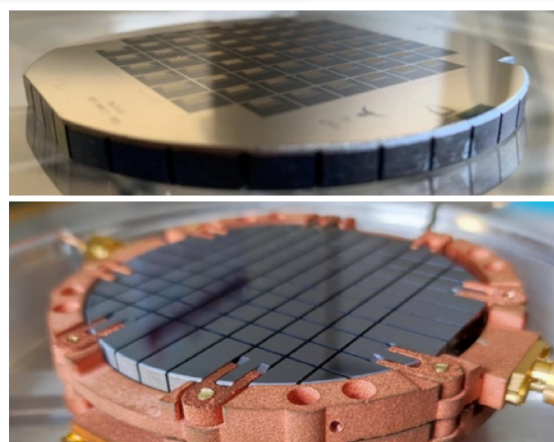
# Understanding the LEE



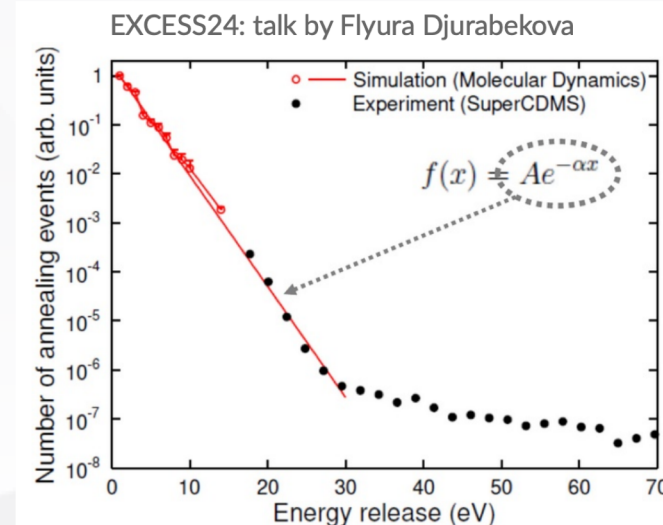
## Stress due to holders!



## Pixelized -> less LEE!

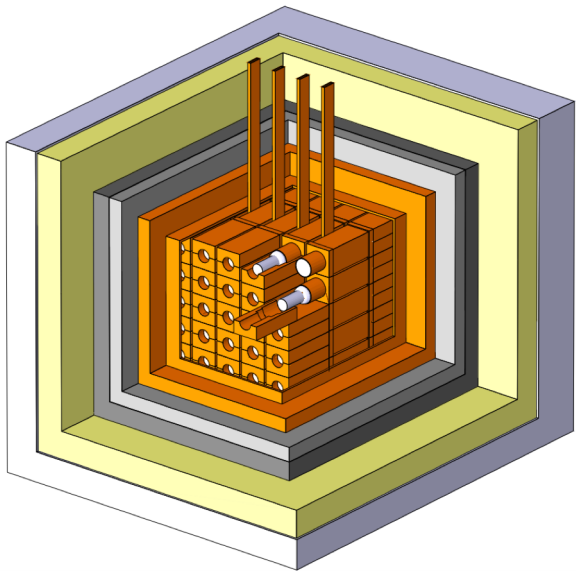


## Sim of internal defects match data!



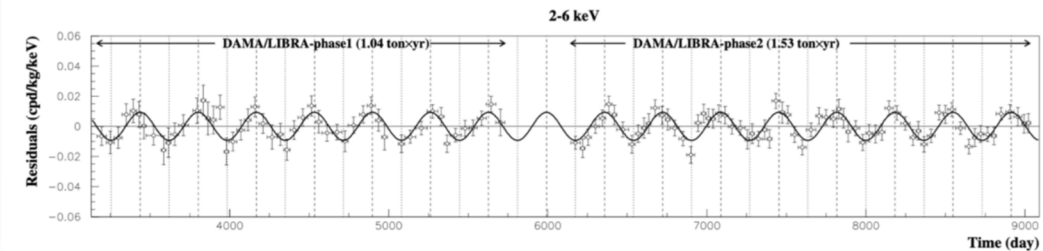
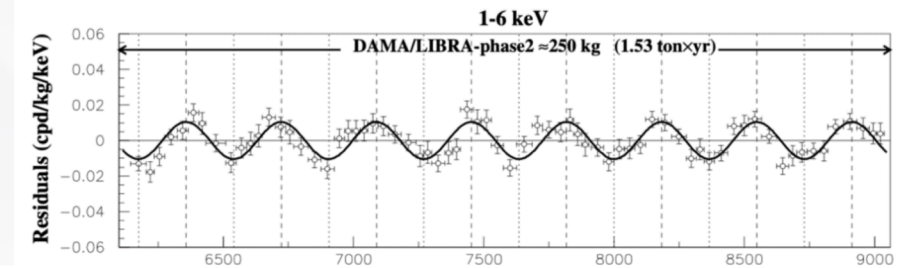
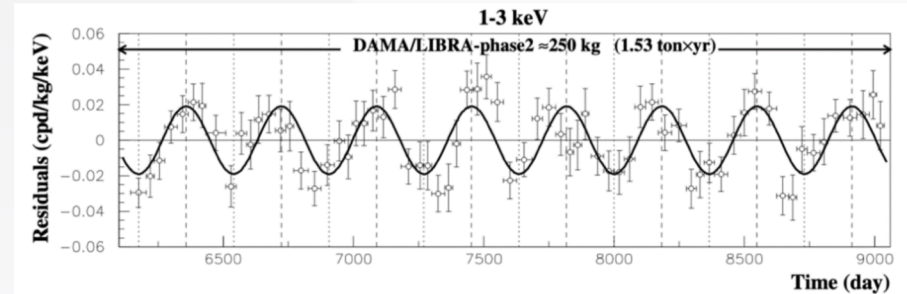
Courtesy to C. Strandhagen, IDM2024

# DAMA/LIBRA (Scintillator)

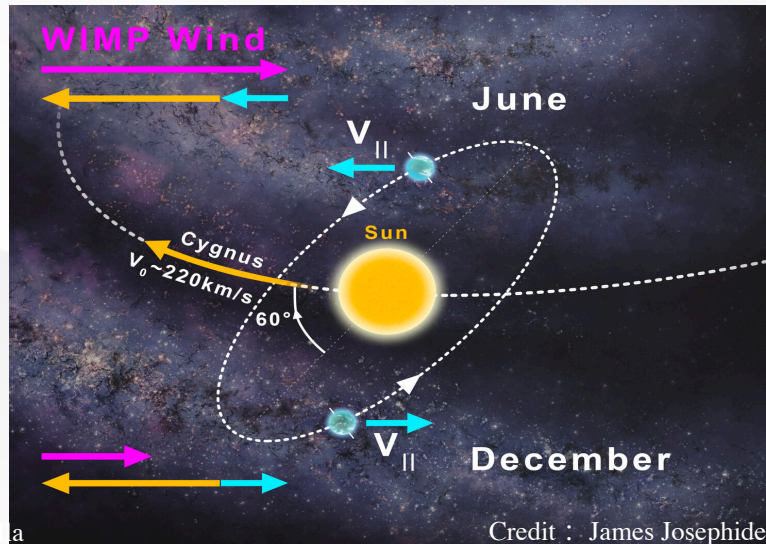


- NaI crystal coupled with PMT at both ends.
- Crystal array embedded in copper cage, then shielded by PE and lead.
- No PID, bkg high;
- ✓ Annual Modulation Signal

Full exposure  $13.7\sigma$



*Nuclear Physics and Atomic Energy.*

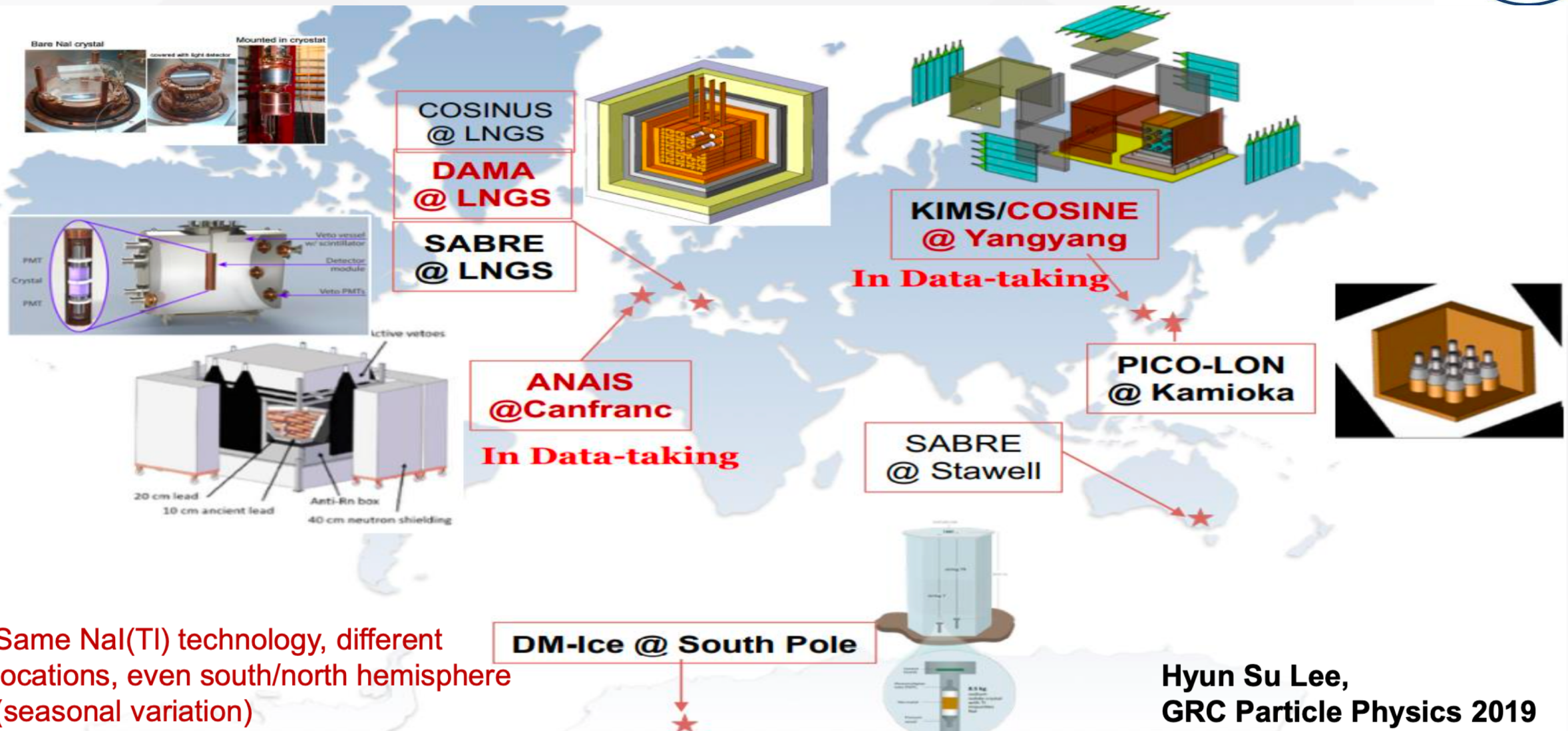


Credit : Fabio Cappella

Credit : James Josephides



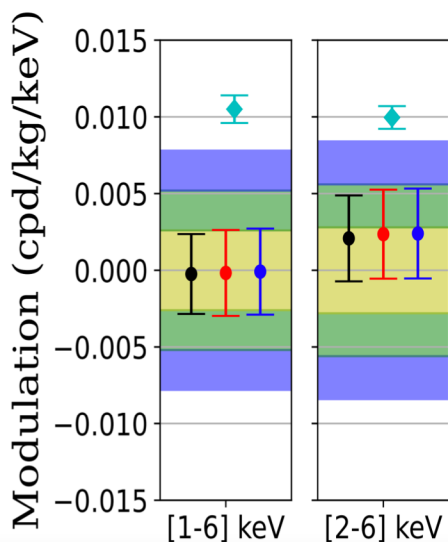
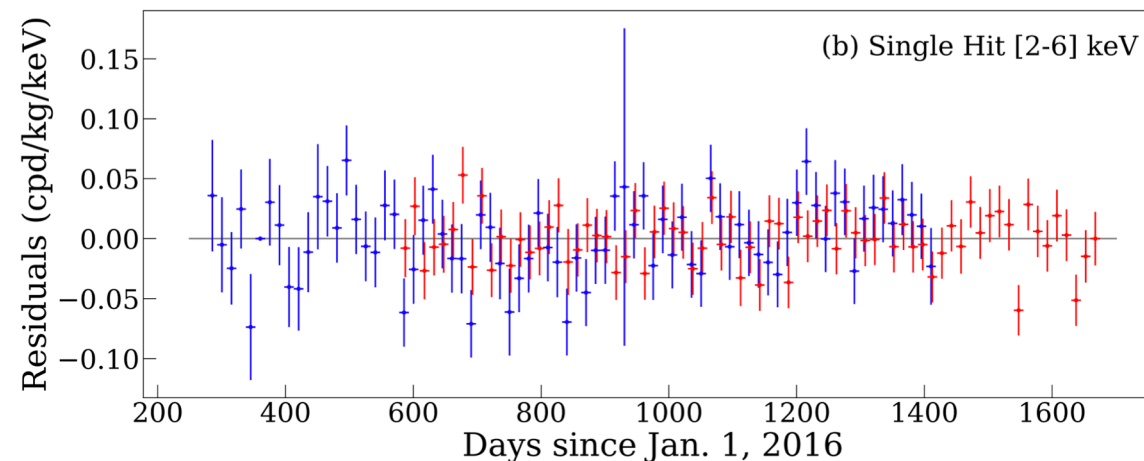
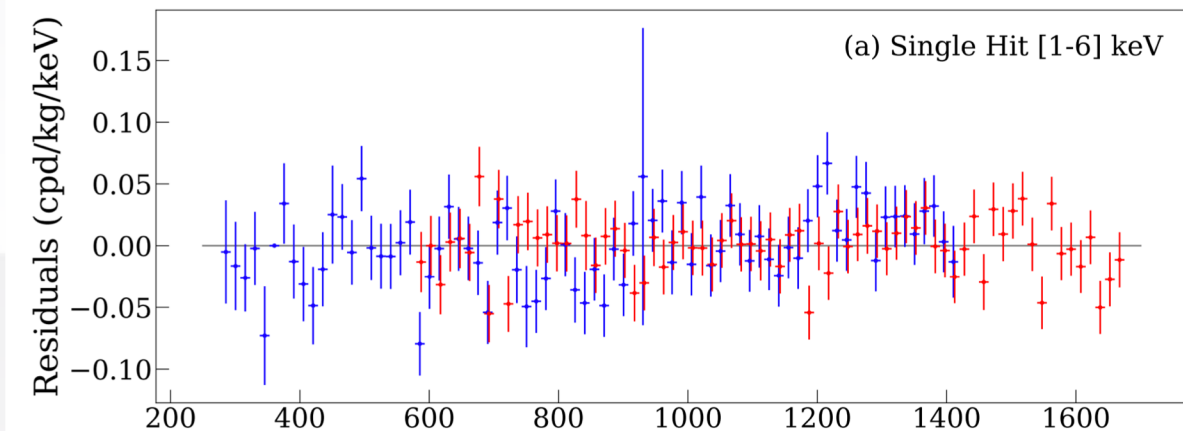
# Annual modulation validation working force



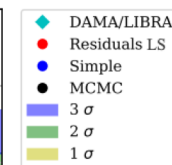
Hyun Su Lee,  
GRC Particle Physics 2019



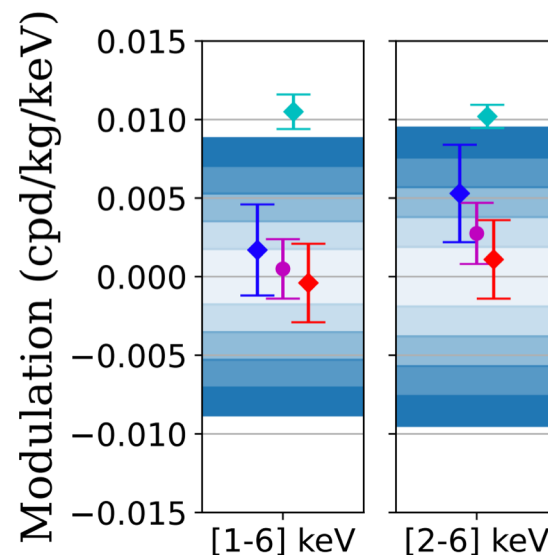
# Joint analysis between COSINE and ANAIS



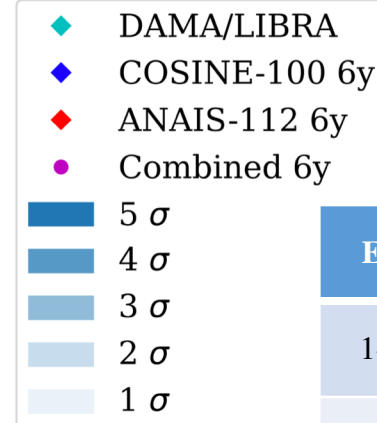
3-yr results



Energy	Amplitude [dru]	DAMA exclusion
1-6 keV	$-0.0002 \pm 0.0026$	$3.7\sigma$
2-6 keV	$0.0021 \pm 0.0028$	$2.6\sigma$



6-yr results

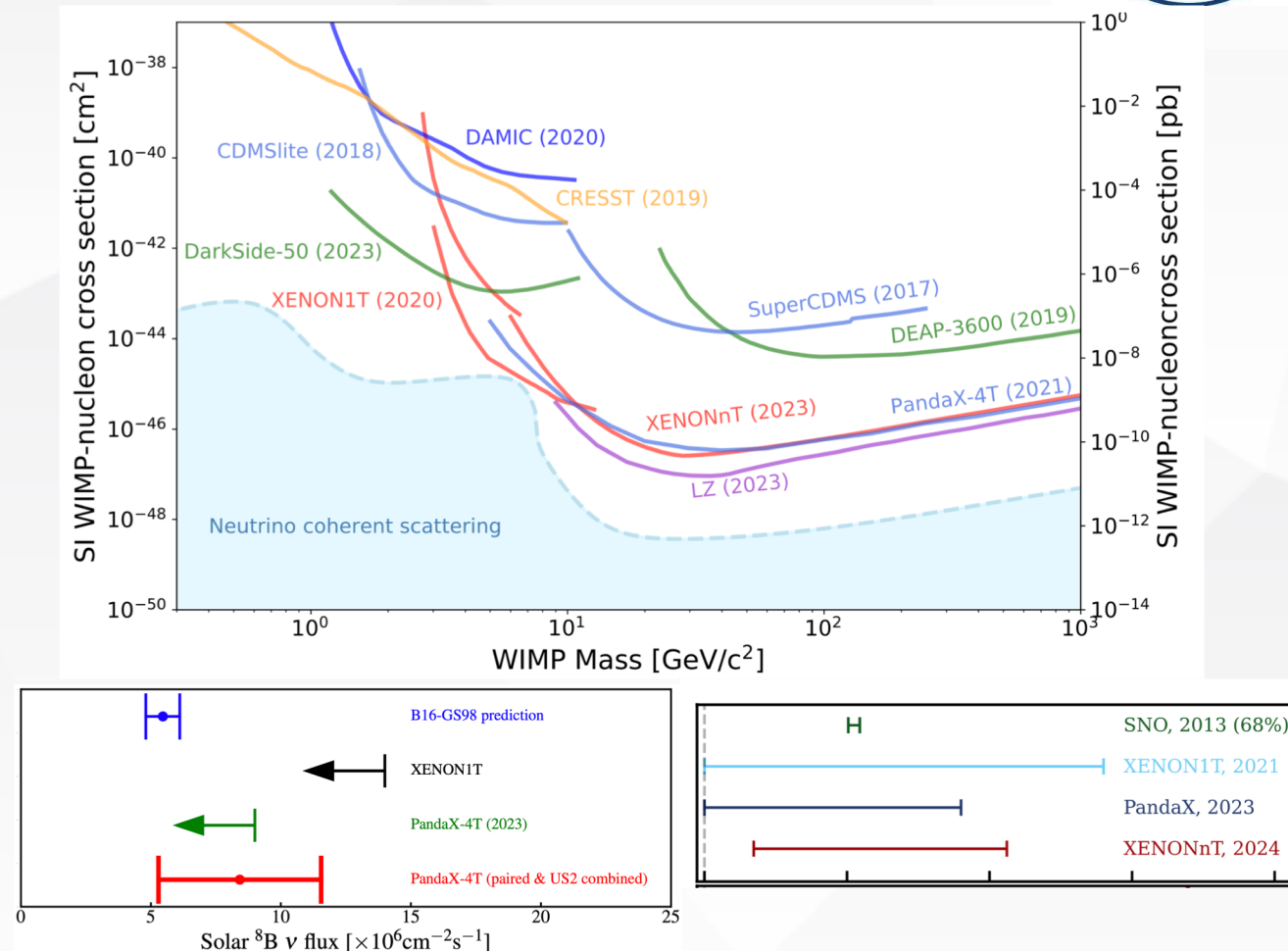


Energy	Amplitude [dru]	DAMA exclusion
1-6 keV	$0.0005 \pm 0.0019$	$4.7\sigma$
2-6 keV	$0.0027 \pm 0.0021$	$3.5\sigma$

# Summary



- Liquid xenon TPCs currently are leading the DM search with mass  $>10\text{GeV}$ ;
- Neutrino floor has been touched with a significance of  $\sim 2.6\sigma$ ;
- Semi-conductors and bolometers are the main forces for  $\sim\text{GeV}$  and sub- $\text{GeV}$  DM searches;
- DAMA/LIBRA's annual modulation DM signals have been almost vetoed by COSINE/ANAIS joint data, with significance of  $4.7\sigma$ ;



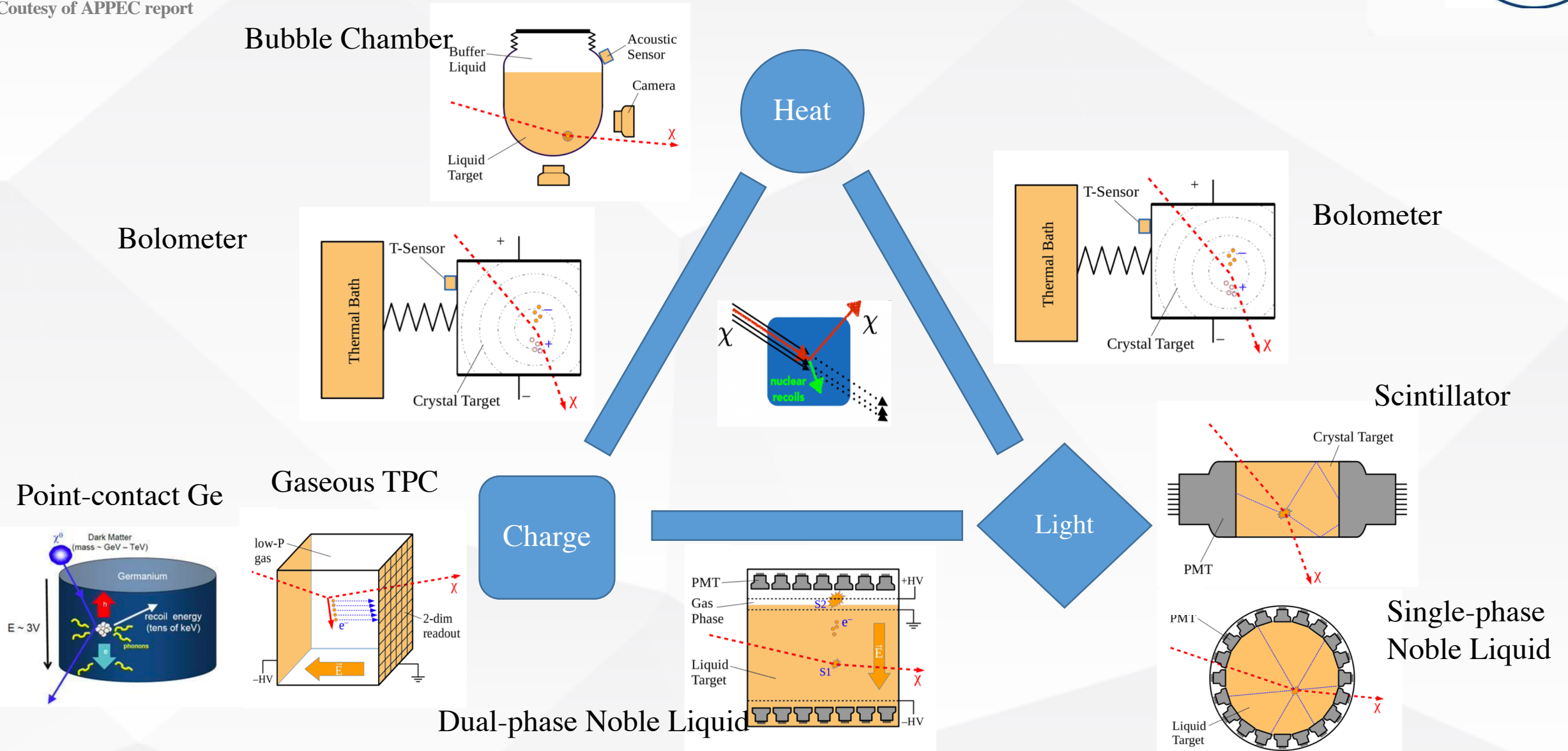
**Thank you for your attention!!**

# Backup



# Detection Techniques

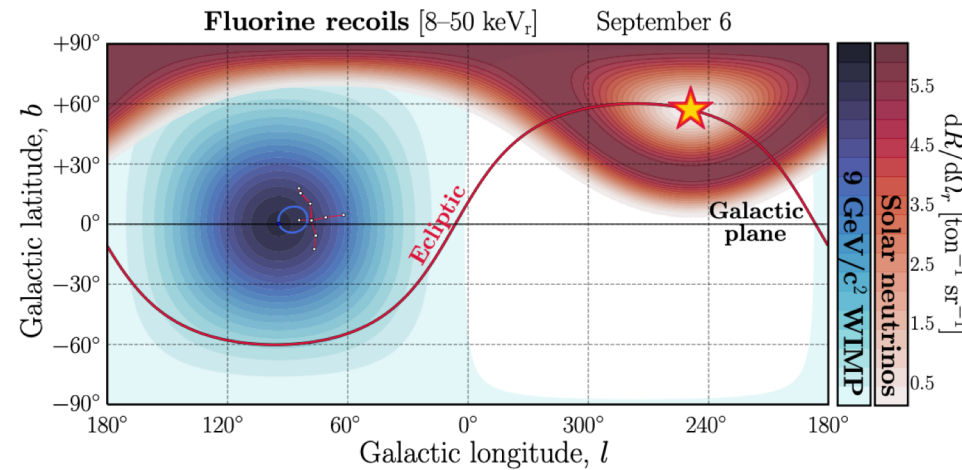
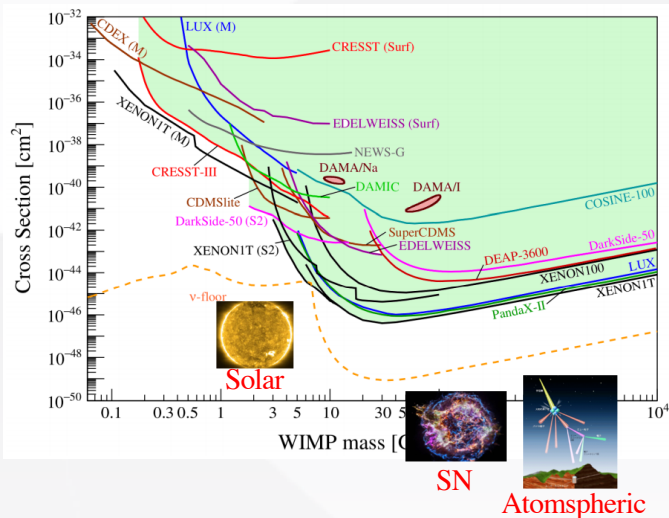
Courtesy of APPEC report



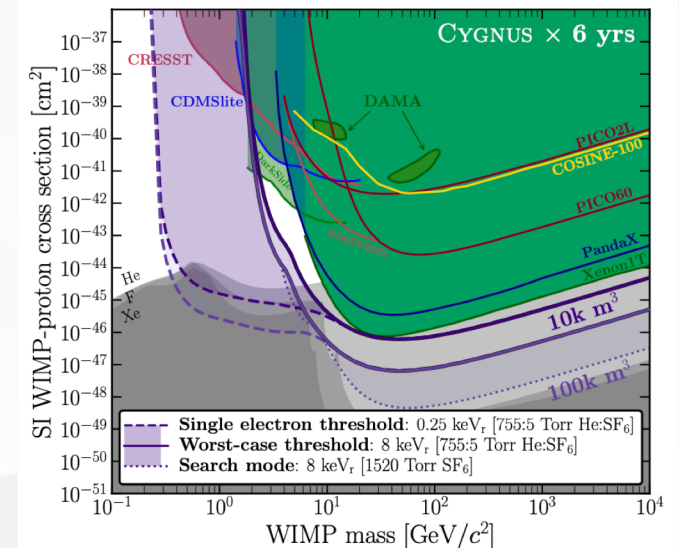
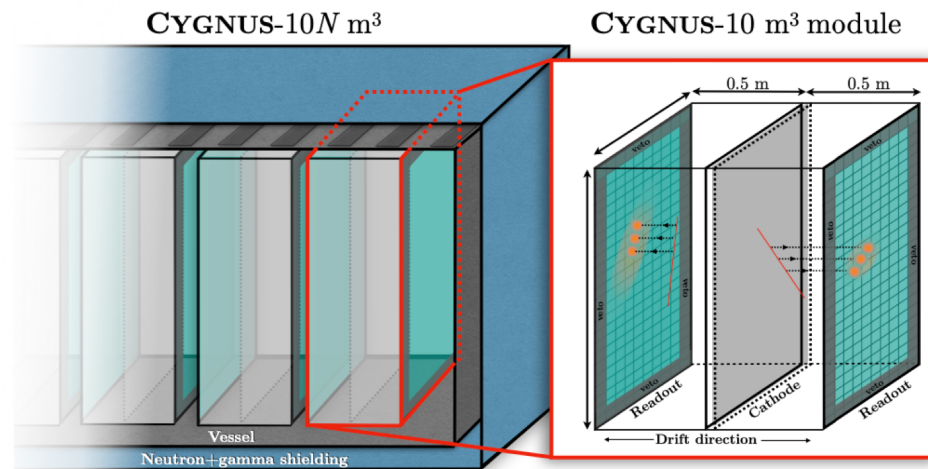
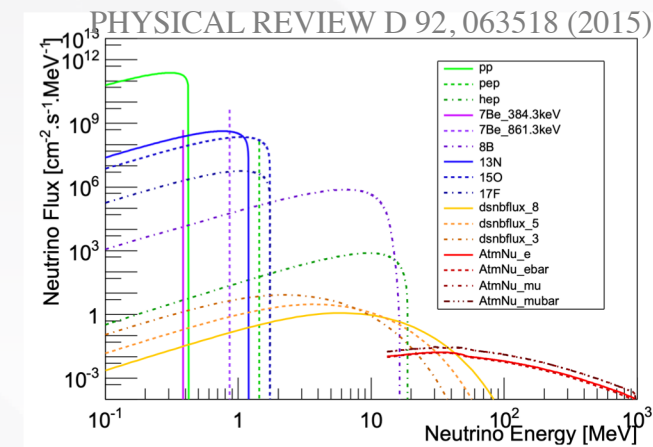


# Directional Detector: Gaseous TPC

SE. Vahsen et al., arXiv:2008.12587



- Low-pressure gaseous TPC;
- Highly pixelized charge readout for high track resolution;
- Directionality can help go beyond neutrino floor to some extent.



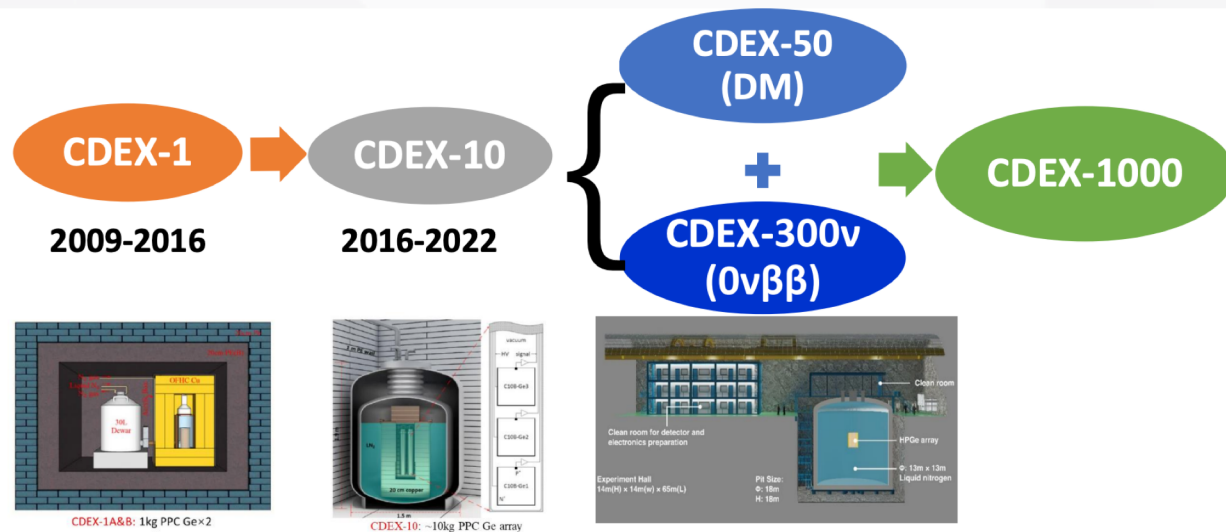
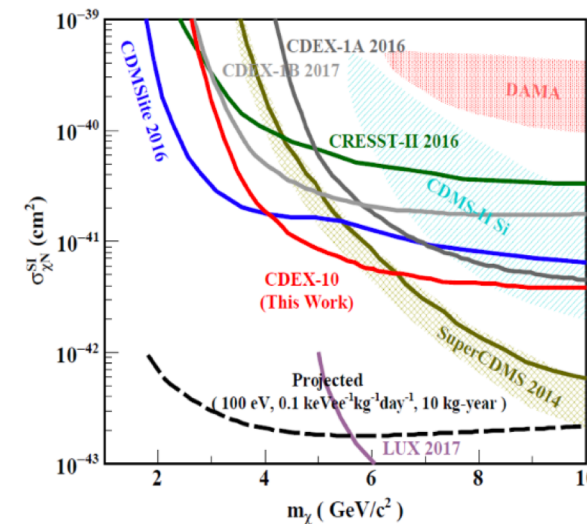
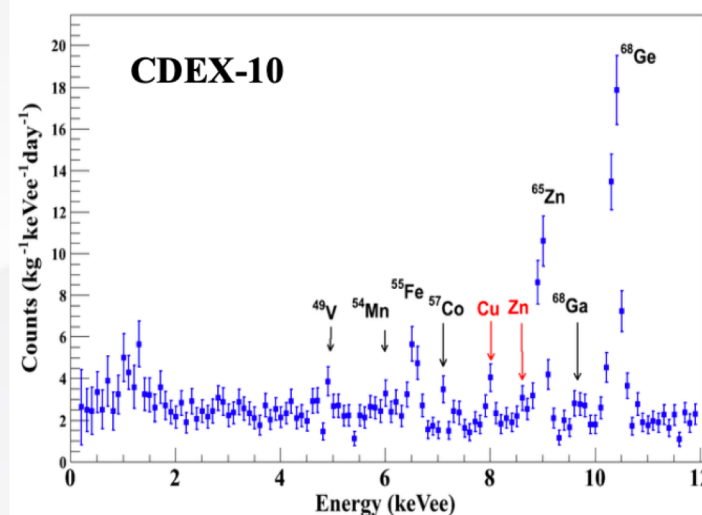
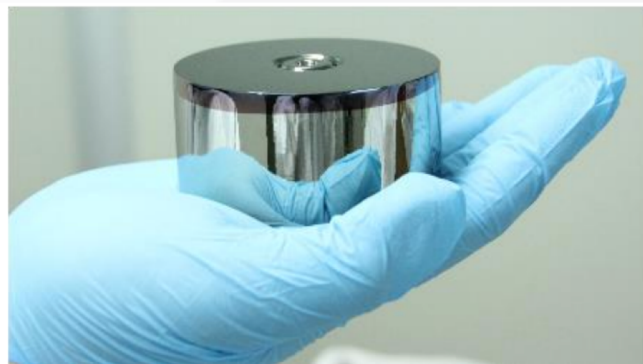
Ultimate background: neutrinos

# CDEX @ CJPL



From Q. Yue

PRL120, 241301 (2018)

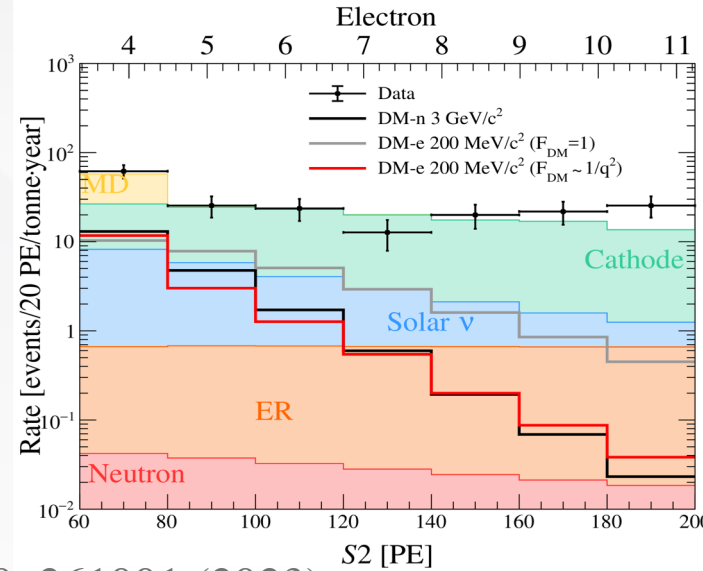


- LN2 temperature; HPGe collects pure ionization signal;
- Point-contact detector (small cap. ~pF; low electronic noise);
- Background level ~2 counts/kg/keV/day (2-4keV);
- CDEX-10, threshold ~160eV;

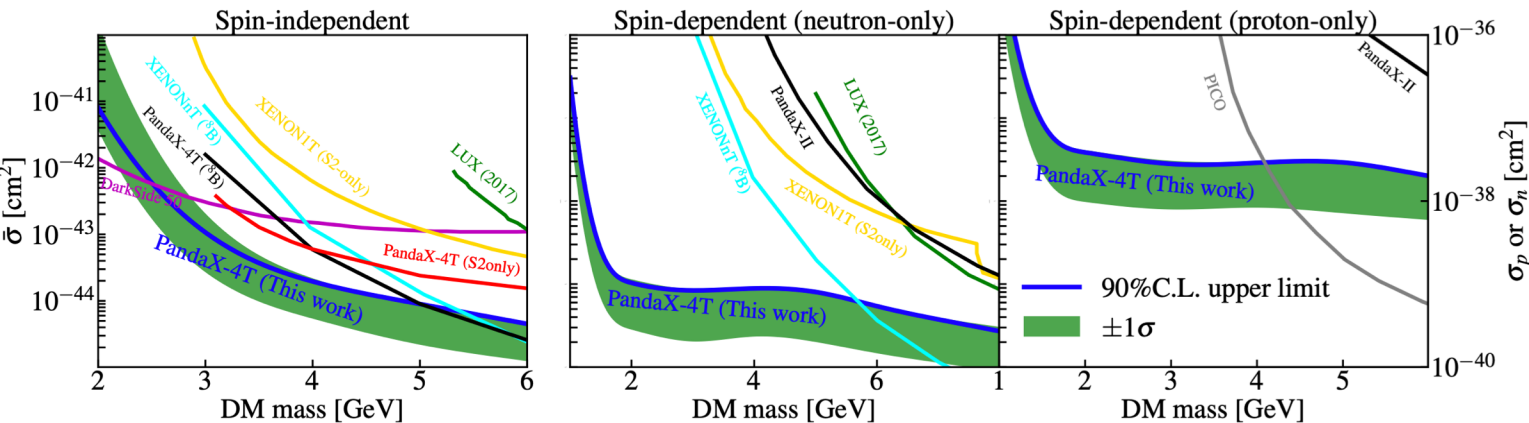


# S2-only in Noble Liquid Detectors

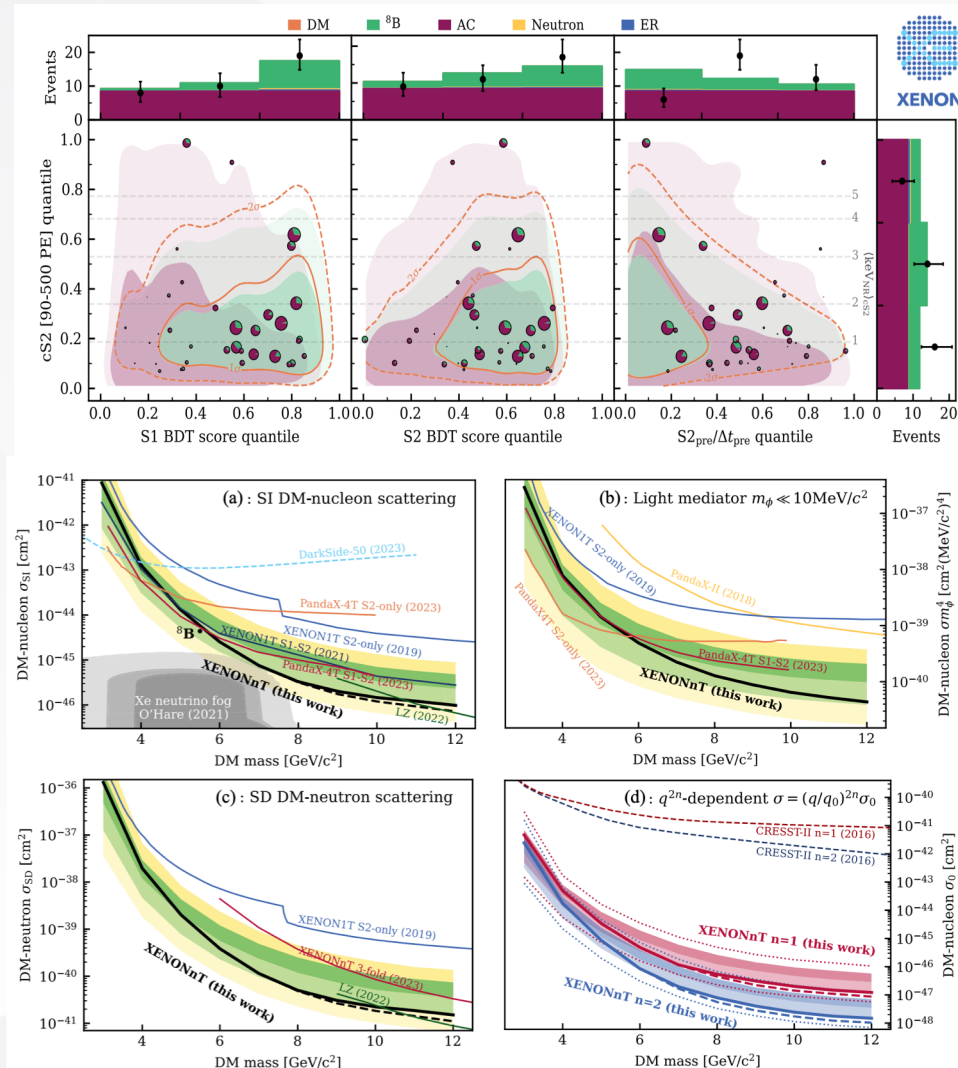
- Greatly lower the threshold for Noble Liquid TPCs;
- Combine with large exposures, can reach leading sensitivities for  $\sim 100$  MeV LDM (WIMP-e scattering);



PandaX-4T, PRL 130, 261001 (2023)  
arXiv: 2507.11930 (2025)



XENONnT, PRL 134, 111802 (2025)

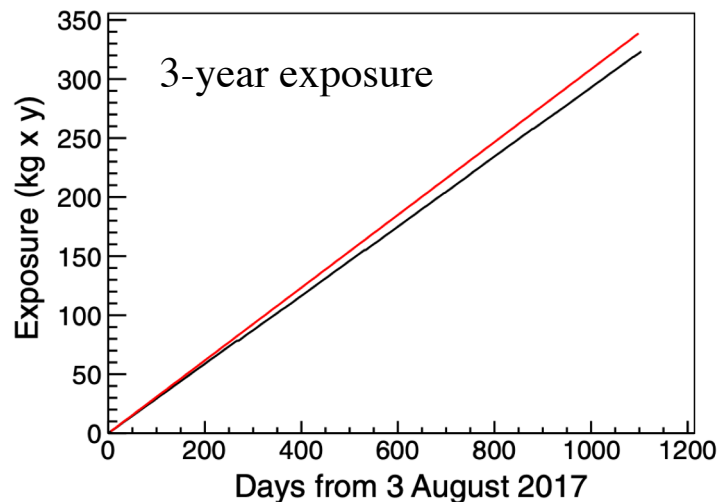
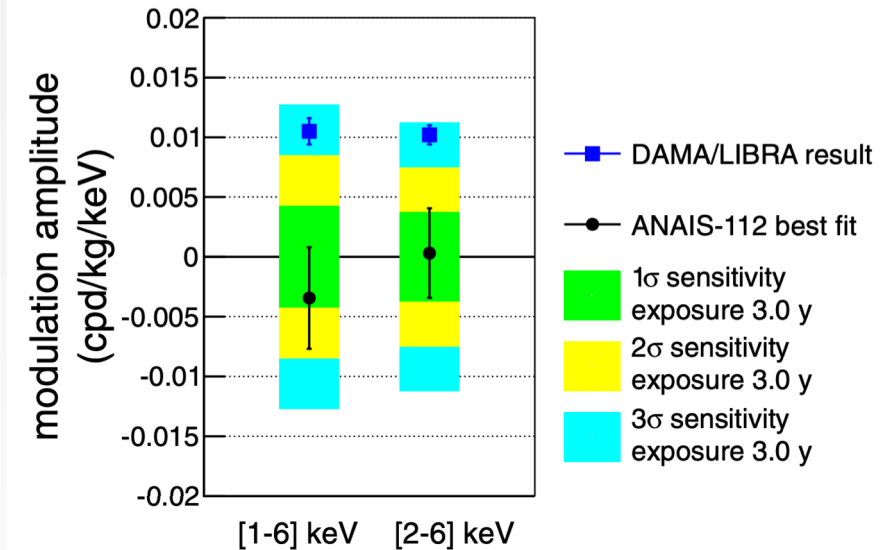
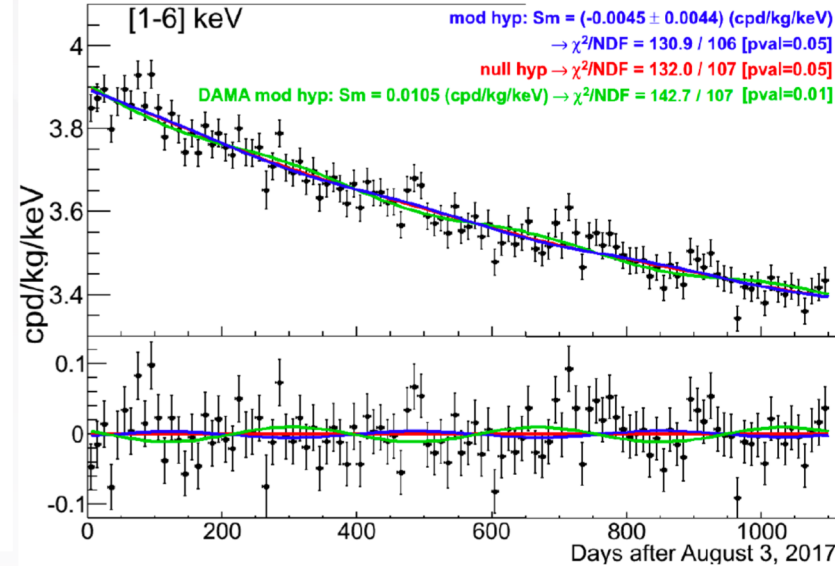
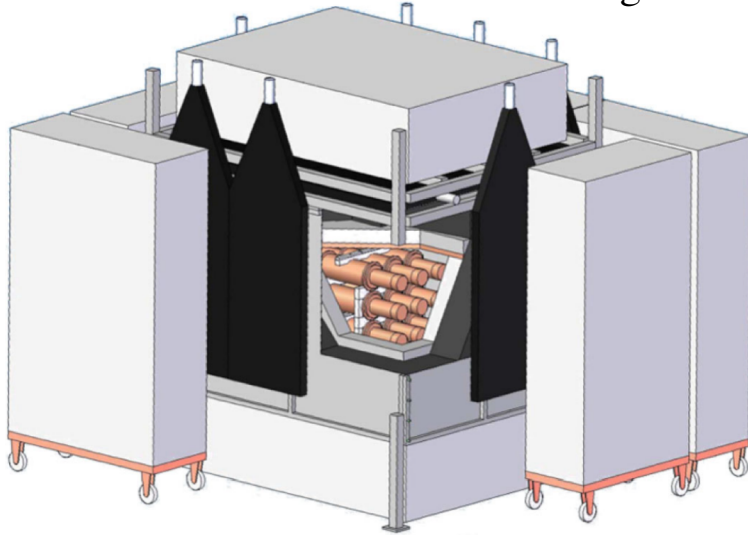


# ANAIS-112



Phys. Rev. D 103, 102005 (2021)

112.5-kg mass

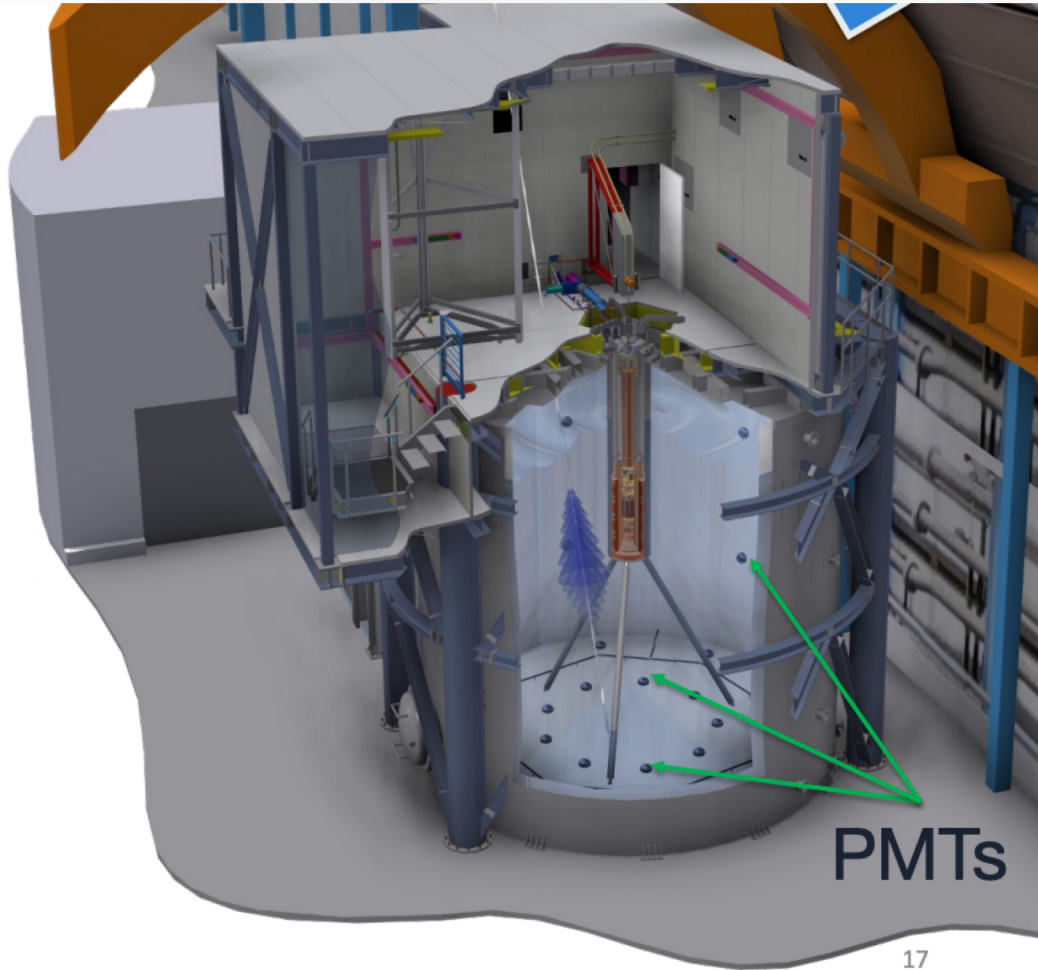


- ANAIS-112's 3-year results support the absence of annual modulation;
- The results are incompatible with DAMA/LIBRA at **3.3 (2.6) σ** for energy window of 1-6 (2-6) keV.

# COSINUS Experiment @ LNGS



Courtesy to F. Reindl, IDM2024



17

Water tank  
2021

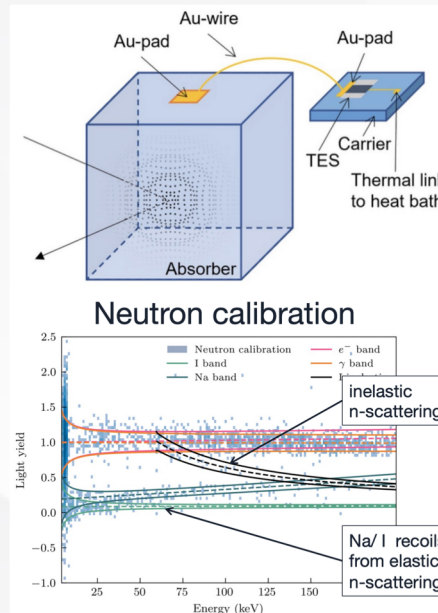


Cryostat delivered  
2023



Buildings finished  
2022

Cryostat installed  
and cooled to 9mK  
2024



- Start data taking in 2025;
- Cryogenic NaI detector with light and phonon readout;
- Event-by-event discrimination;
- Counting & Modulation searches;



# CONTENTS

---



**01**

**DM direct search**

**02**

**>GeV Searches**

**03**

**~GeV Searches**

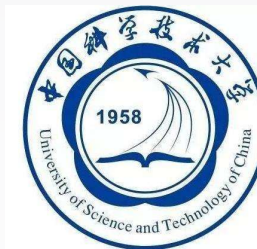
**04**

**<GeV Searches**

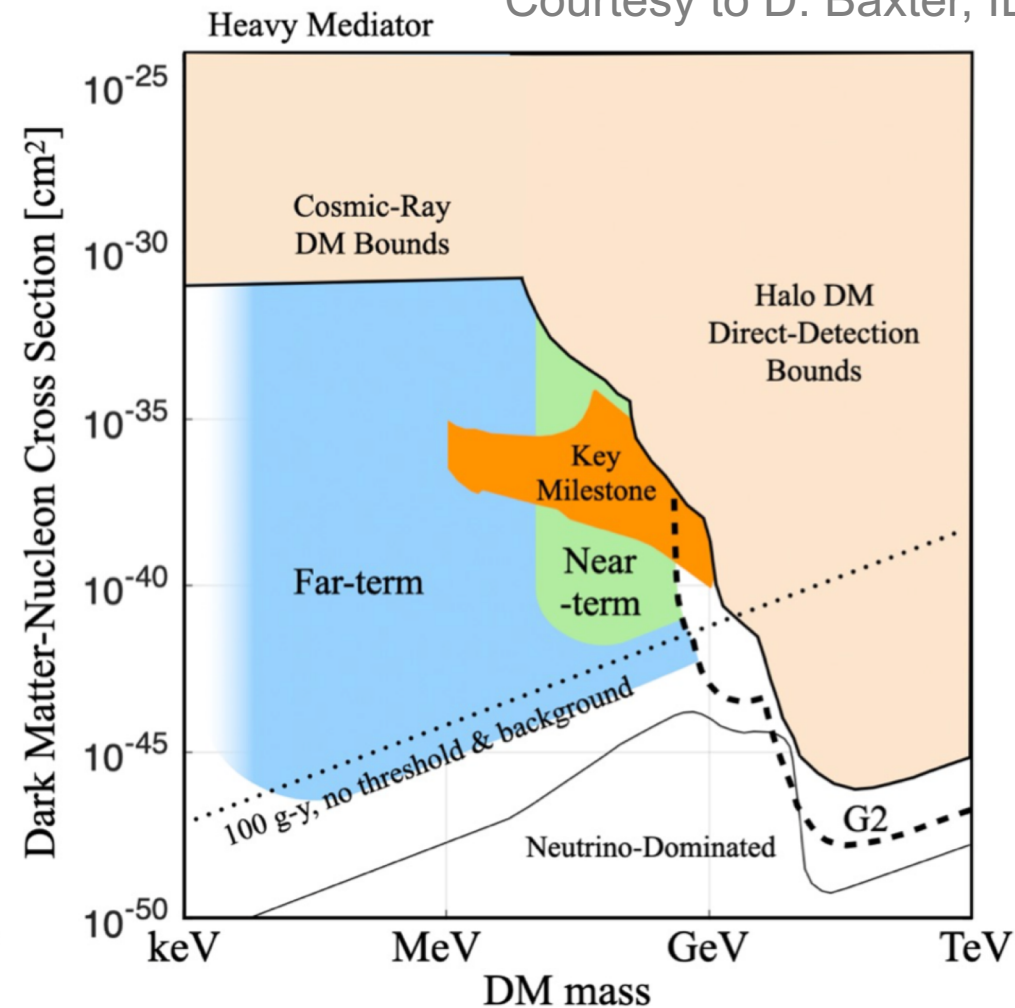
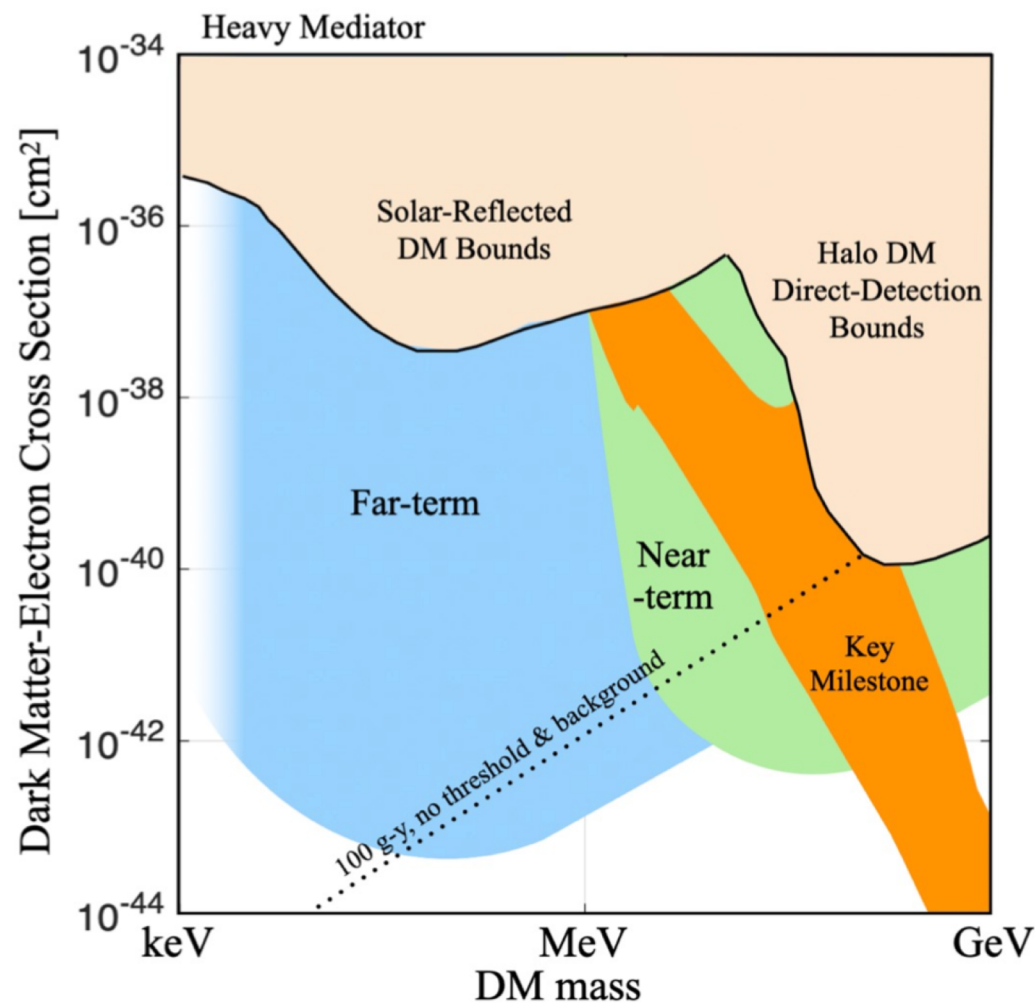
**\*\*Disclaimer: Apology for omittance!\*\***

**\*\* Most materials are borrowed from public talks, especially from Prof J. Liu, Prof. Q. Yue's, and talks from IDM2024.\*\***

# Low energy regimes



Courtesy to D. Baxter, IDM2024

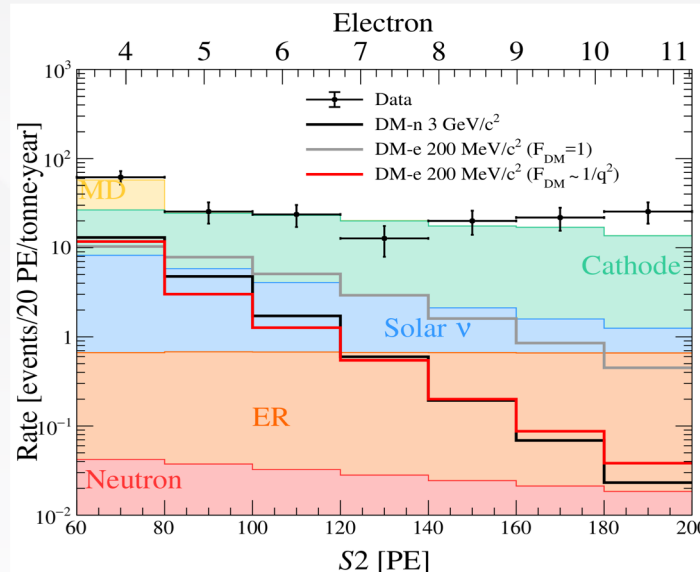


Essig et al, Snowmass CF1 WP2 (2022) [arXiv:2203.08297]

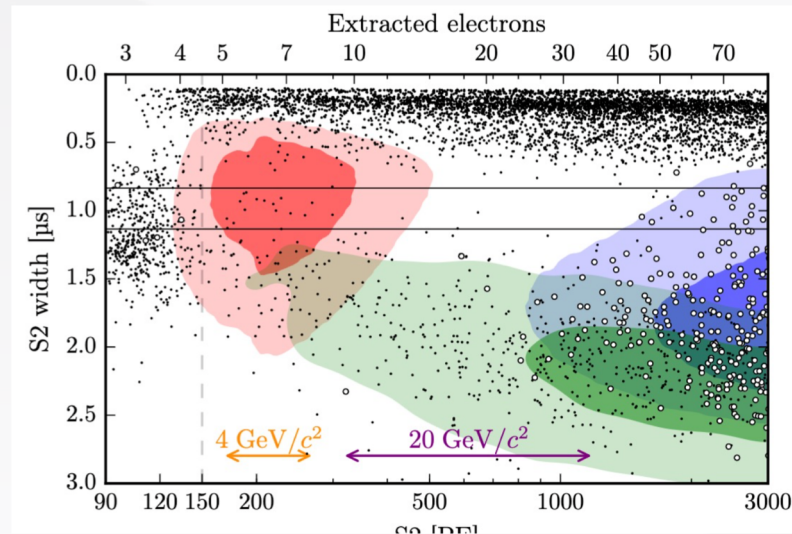
# S2-only in Noble Liquid Detectors



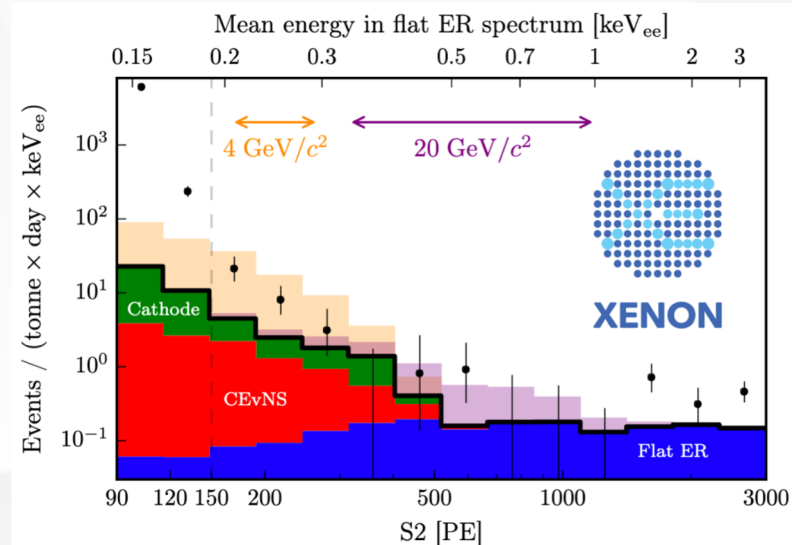
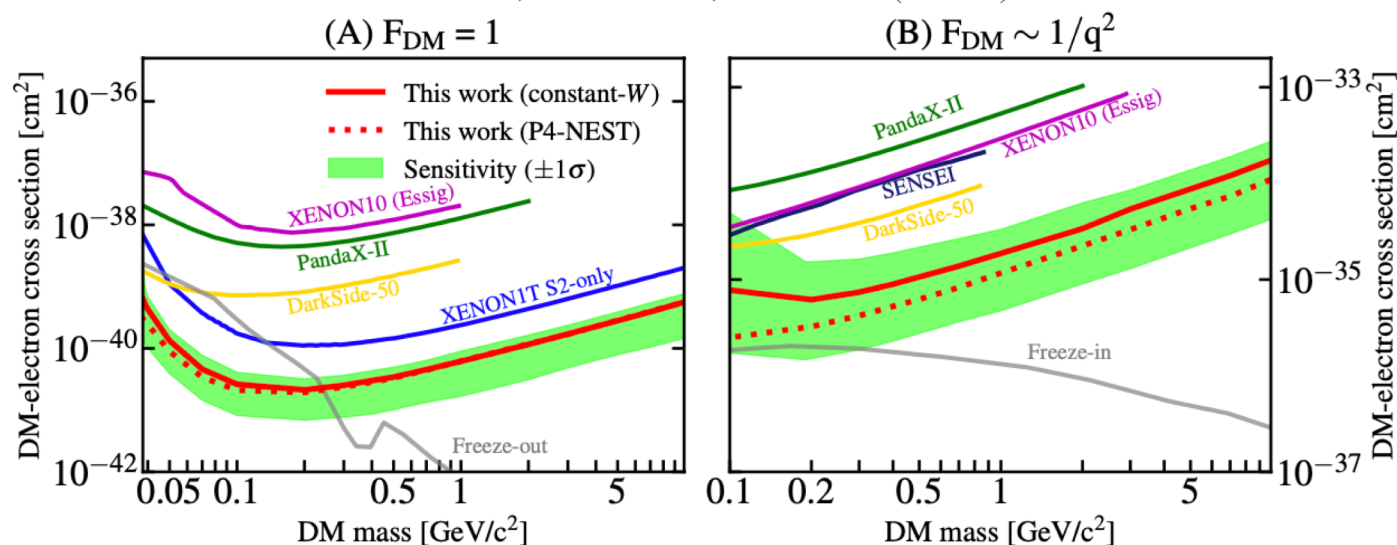
- Greatly lower the threshold for Noble Liquid TPCs;
- Combine with large exposures, can reach leading sensitivities for  $\sim 100$  MeV LDM (WIMP-e scattering);



XENON1T, PRL 123, 251801 (2019)

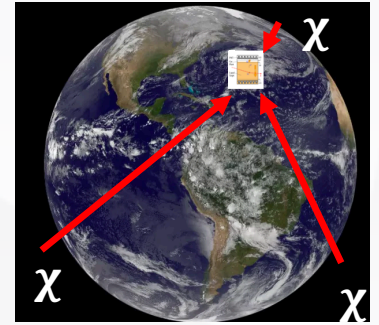
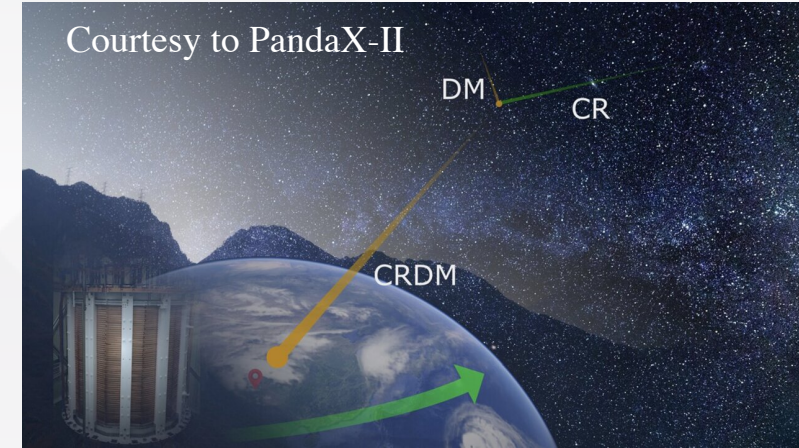
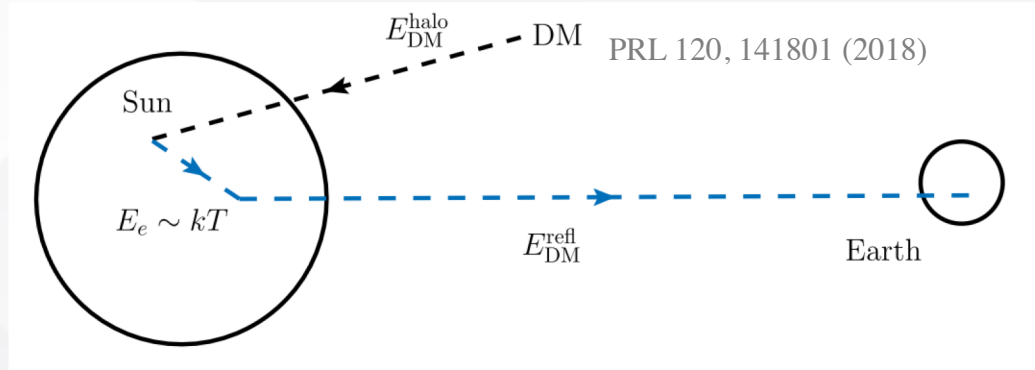


PandaX-4T, PRL 130, 261001 (2023)

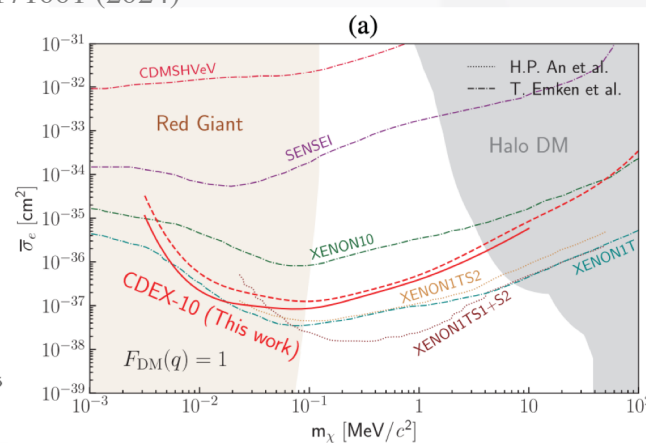
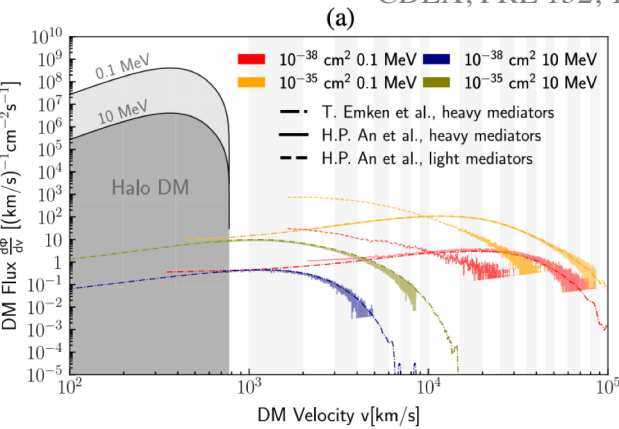




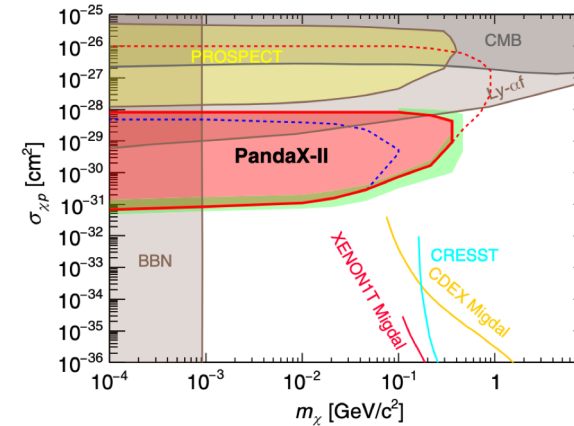
# Solar reflected and cosmic ray boosted DMs



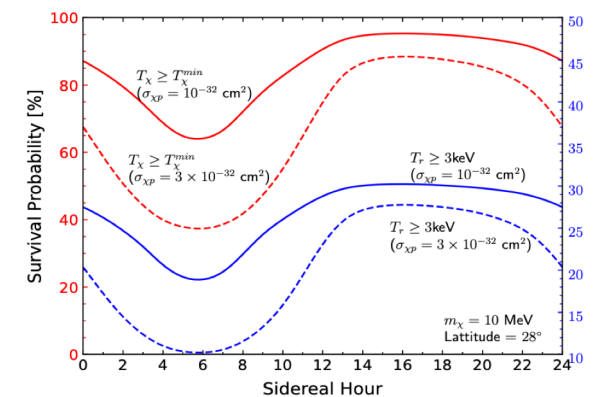
CDEX, PRL 132, 171001 (2024)



PandaX-II, PRL 128, 171801 (2022)



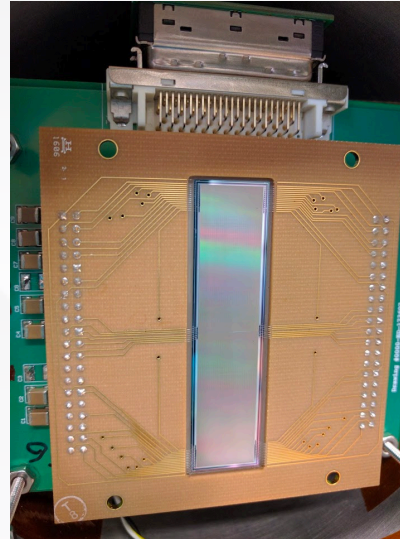
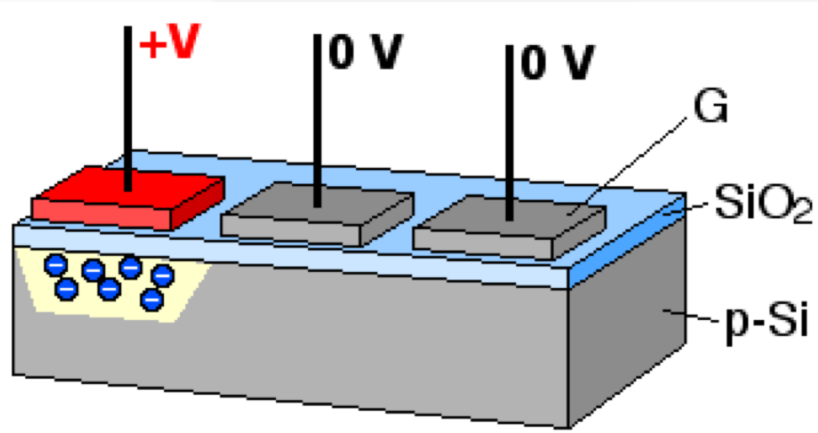
Ge, Shao-Feng, et al. PRL 126, 091804 (2021)



Diurnal Effect

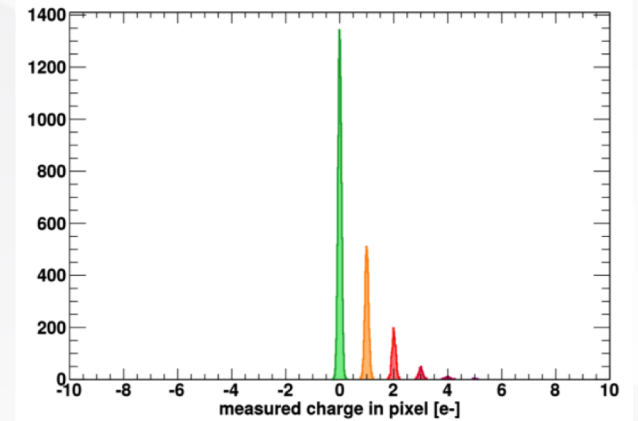
Accelerated DM by solar electrons or “heated” cosmic rays in center of MW surpass thermal speed!

# CCD detectors

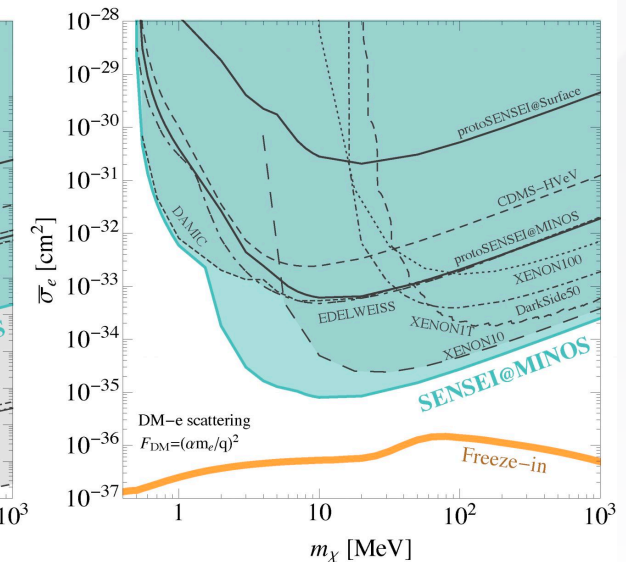
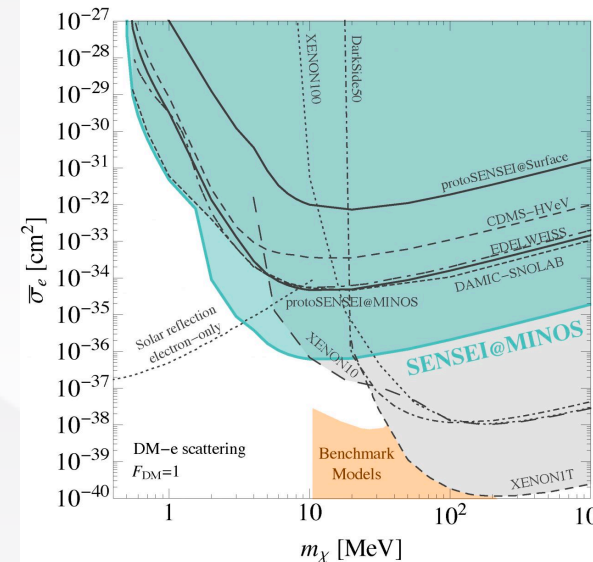


<https://news.fnal.gov/2020/06/sensei-gets-quiet/>

Readout-noise: 0.067 e RMS



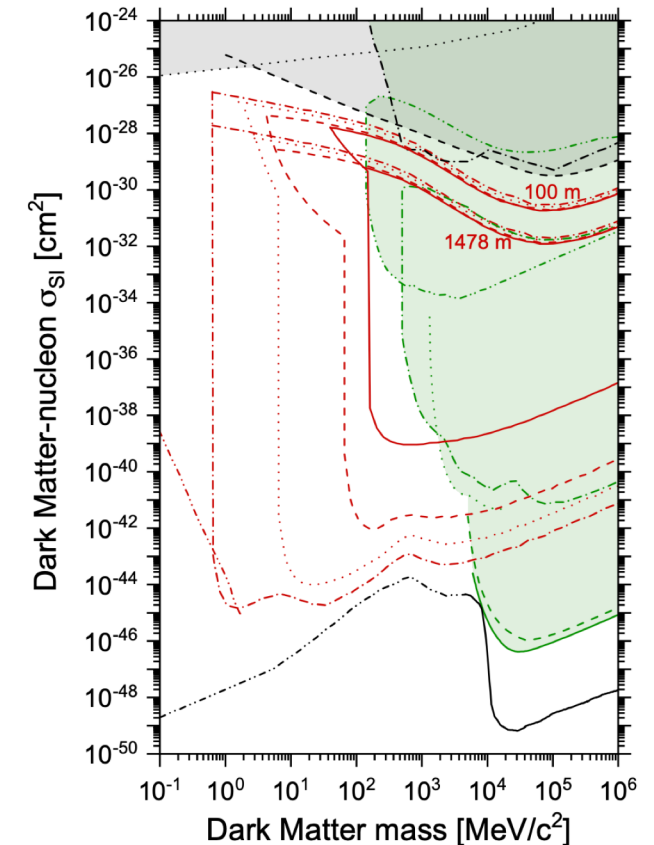
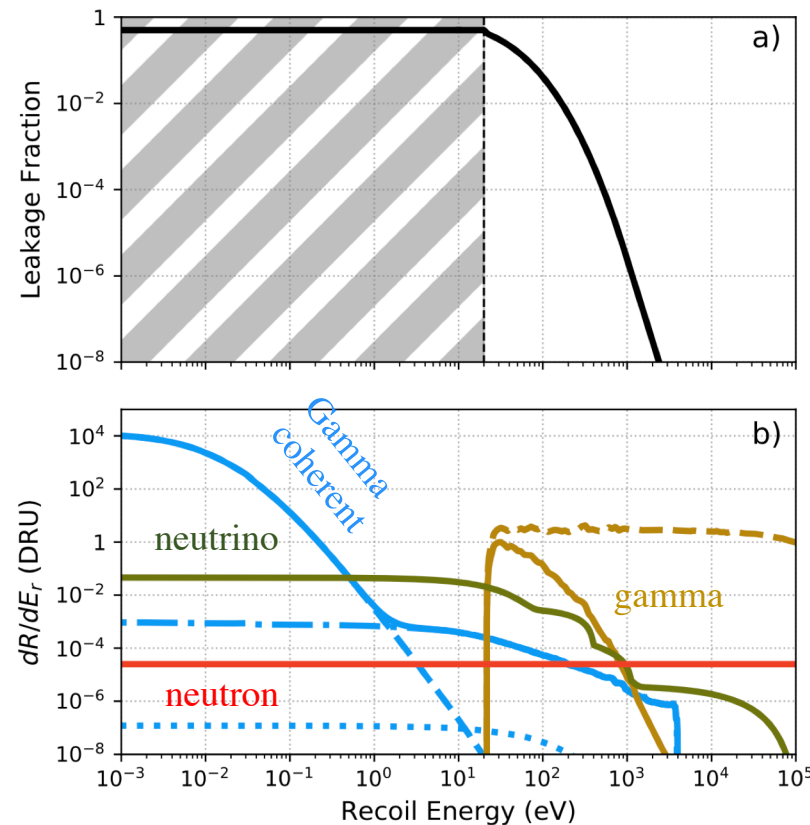
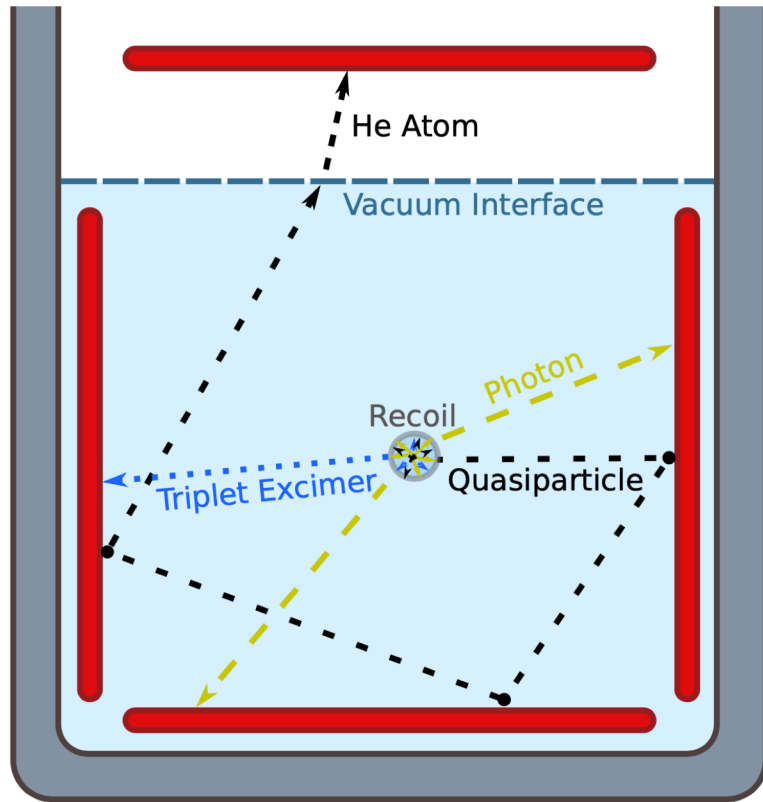
- 0.01 g Si Skipper CCD;
- Reduce noise through non-destructively measure charge multiple times;
- Single electron sensitivity with resolution to 6.7%.



# Superfluid Helium Detector



Phys. Rev. D **100**, 092007 (2019)

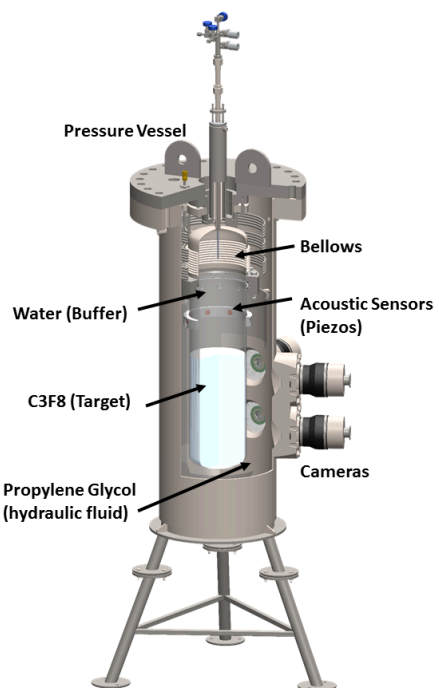


- Very low  $W$  value for roton and phonon;
- Quasiparticle propagate ballistically, including directionality;

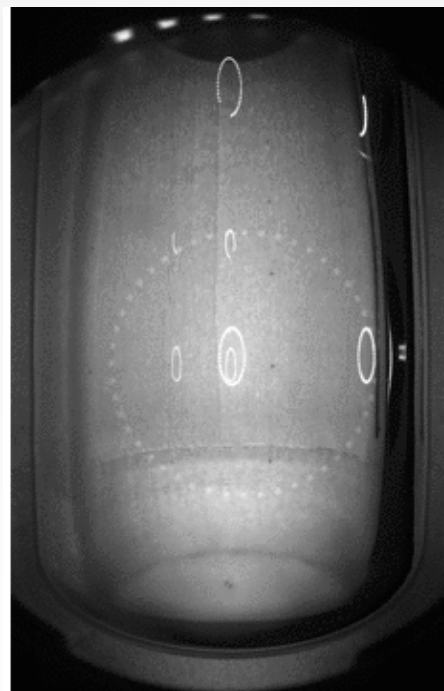
- Background dominated by gamma and neutrino coherent scattering;



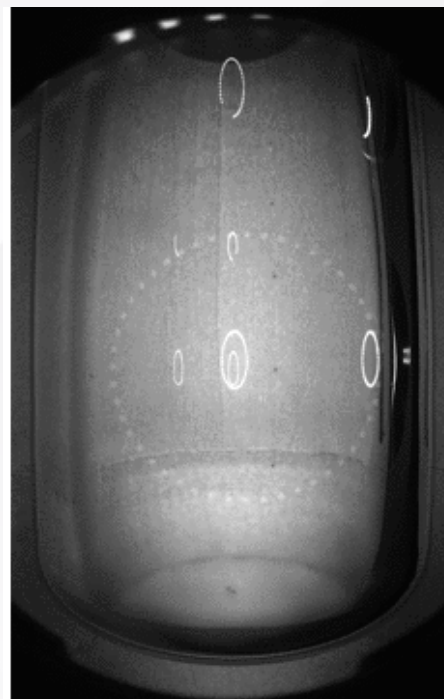
# Bubble Chamber: PICO



Single Scatter

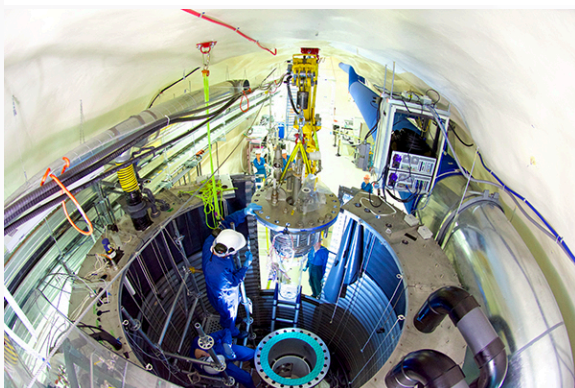
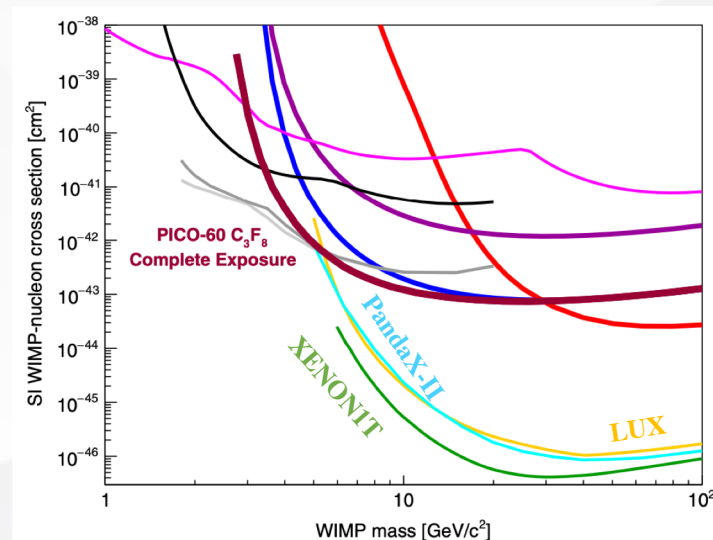
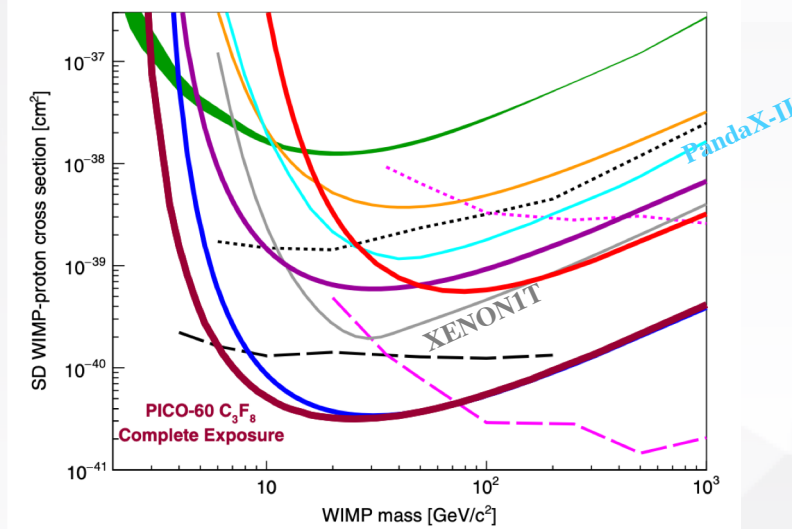


Multi-Scatter

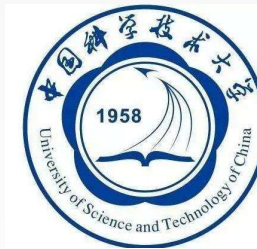


<https://www.picoexperiment.com/pico-60/>

- Largest bubble chamber with superheated  $C_3F_8$ ;
- No bkg from gamma, e and muons;
- Discrimination against neutron through single and multiple scatter images;

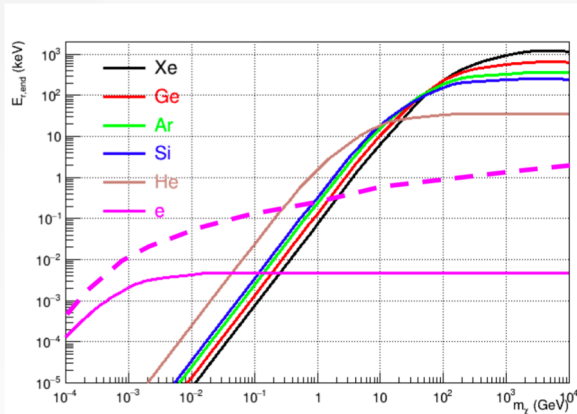
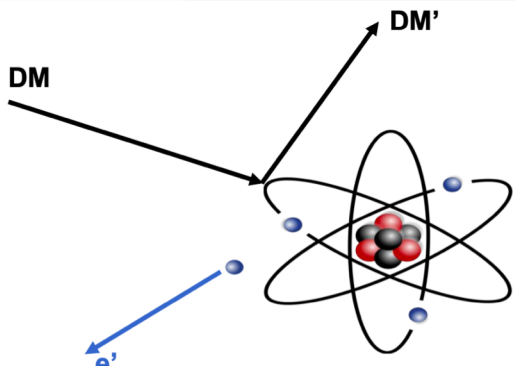


# DM-e scattering and Migdal effect

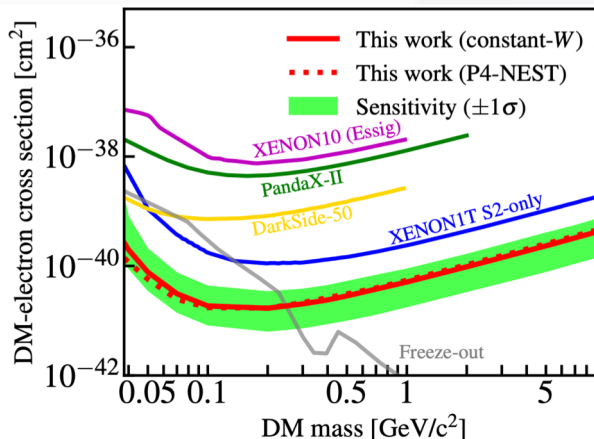
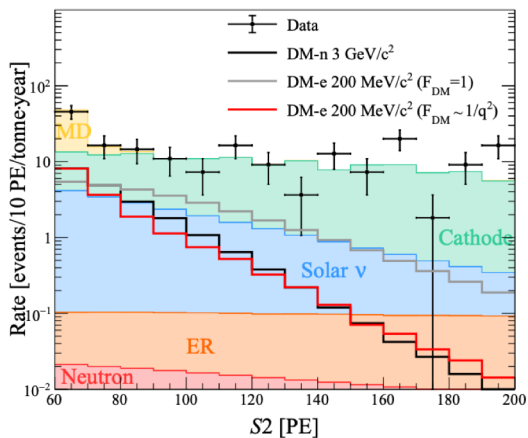


## DM-electron scattering

Essig, Mardon, Volansky, PRD 85, 076007 (2012)

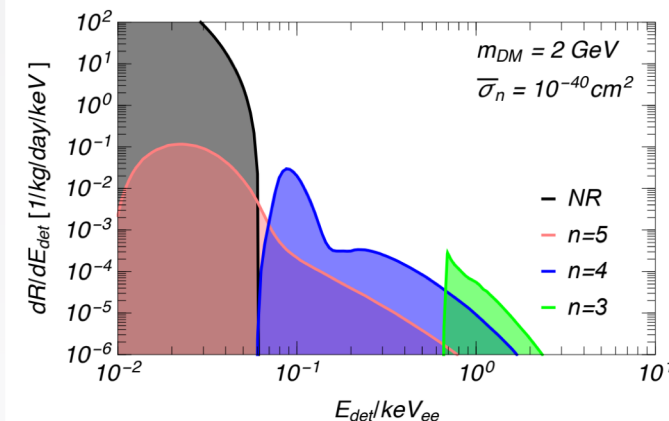
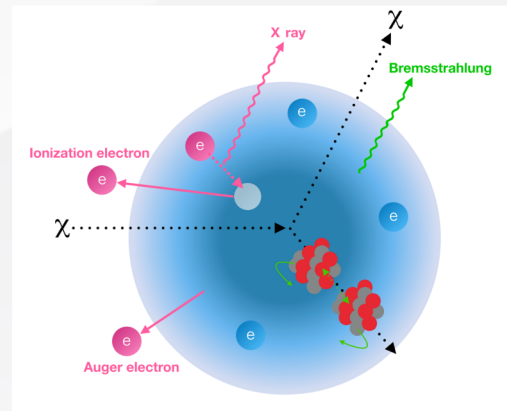


PandaX collaboration, arXiv: 2212.10067 (2022)

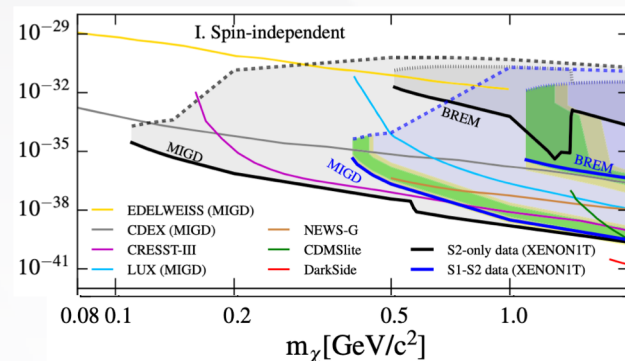


## Migdal Effect

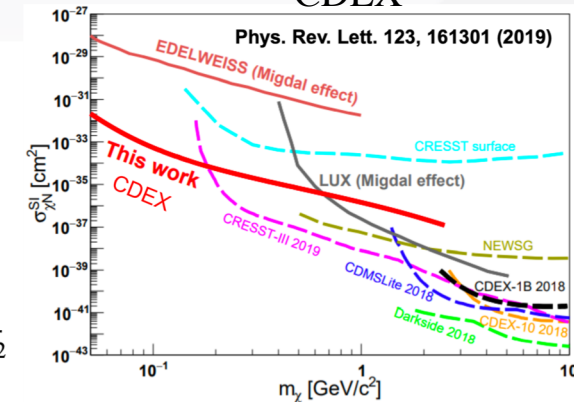
M. Ibe et al., Journal of High Energy Physics 2018.3 (2018)  
XENON collaboration, PRL 123, 241803 (2019)



## XENON1T



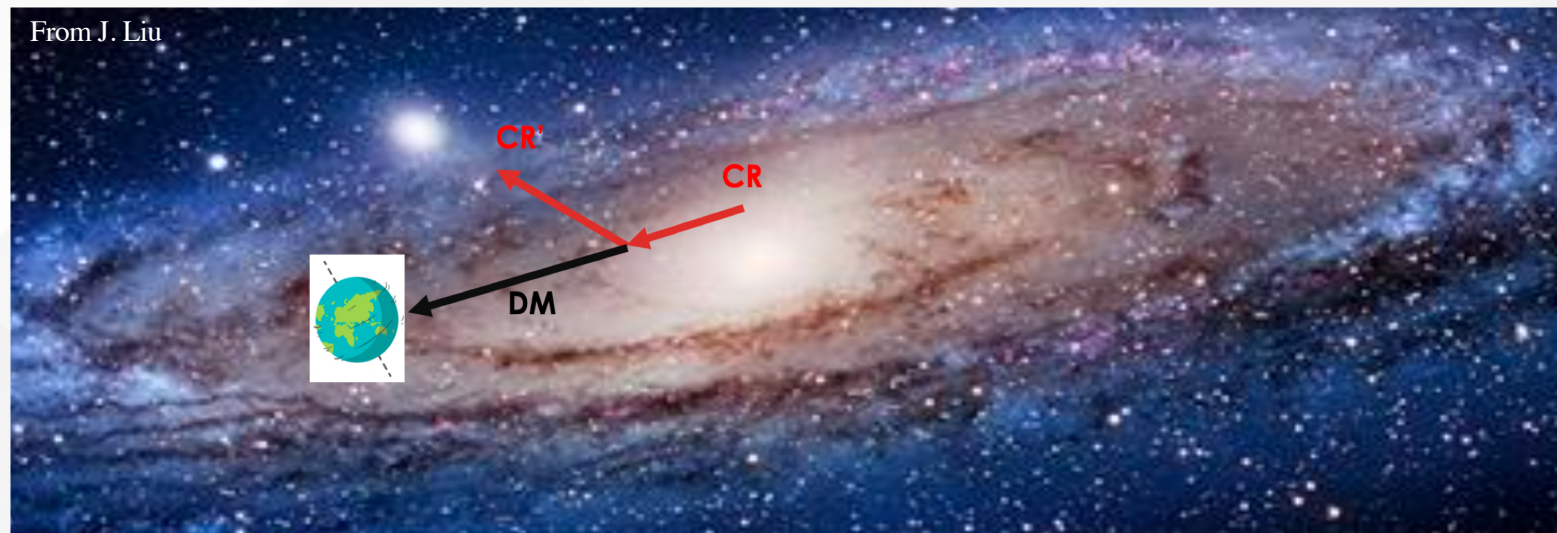
## CDEX



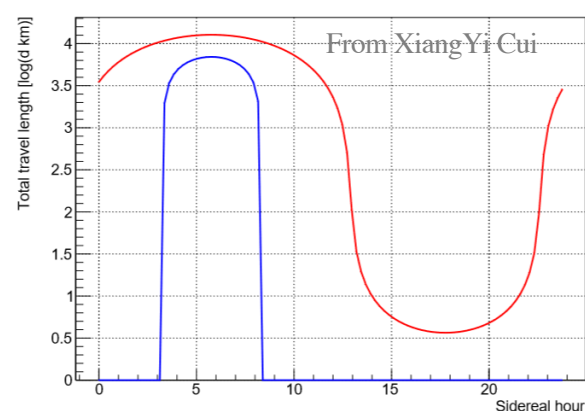
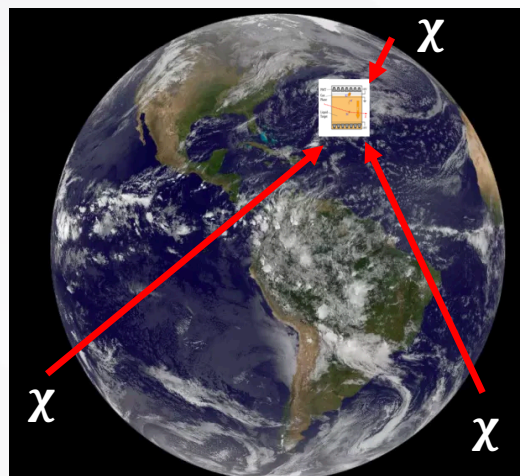
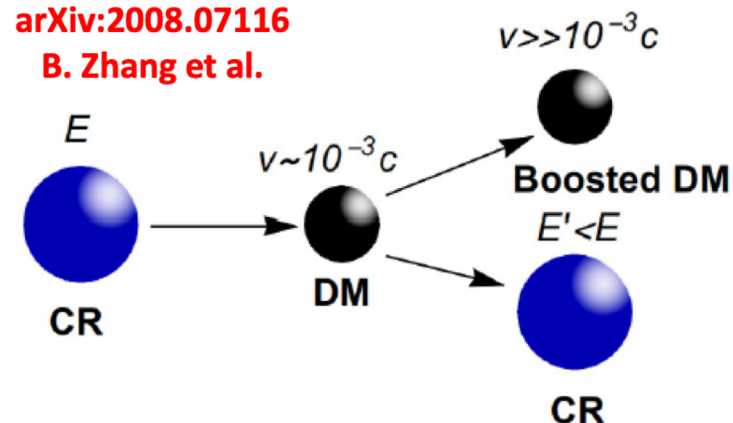
Capable of Probing light dark matter due to high momentum transfer



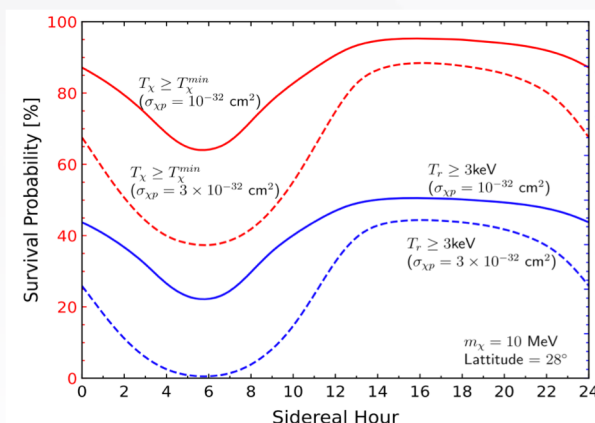
# Diurnal Modulation



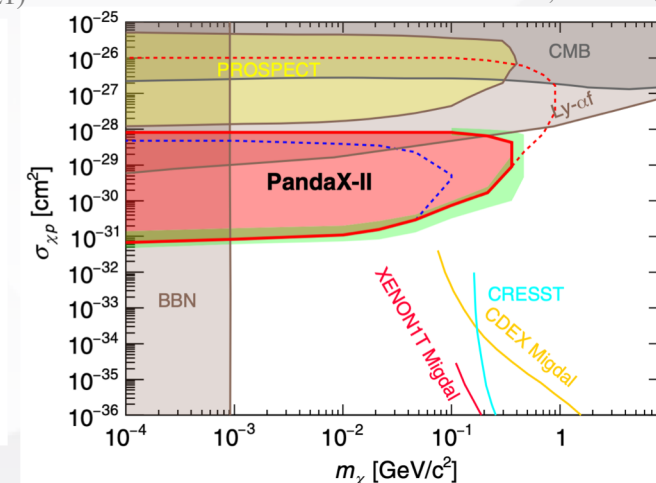
arXiv:2008.07116  
B. Zhang et al.



PHYSICAL REVIEW LETTERS 126, 091804 (2021)



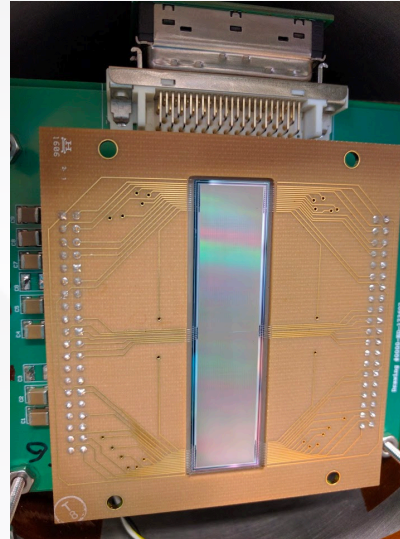
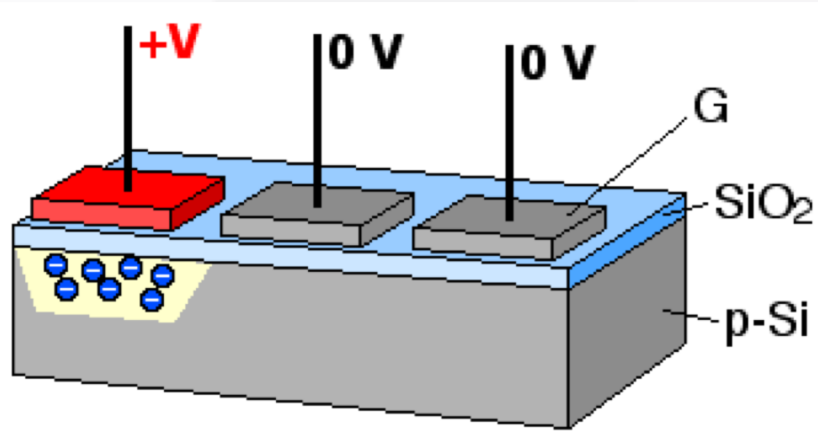
PHYSICAL REVIEW LETTERS 128, 171801 (2022)



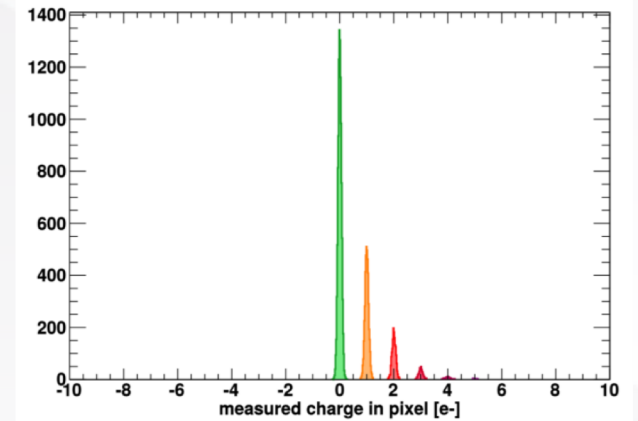
Expected diurnal modulation due to traveling length's day-night difference for transpassing boosted DM.



# Skipper CCD technique: SENSEI

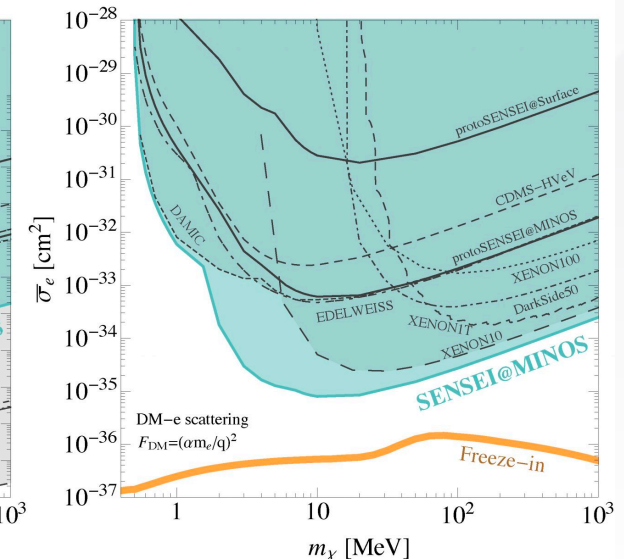
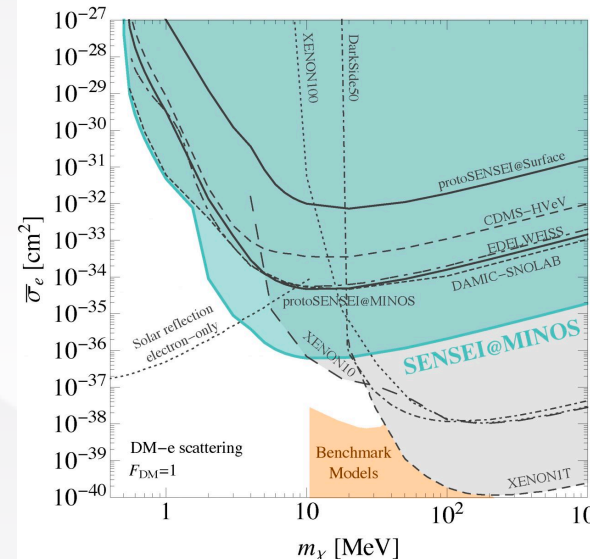


Readout-noise: 0.067 e RMS

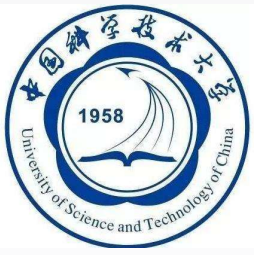


- 0.01 g Si Skipper CCD;
- Reduce noise through non-destructively measure charge multiple times;
- Single electron sensitivity with resolution to 6.7%.

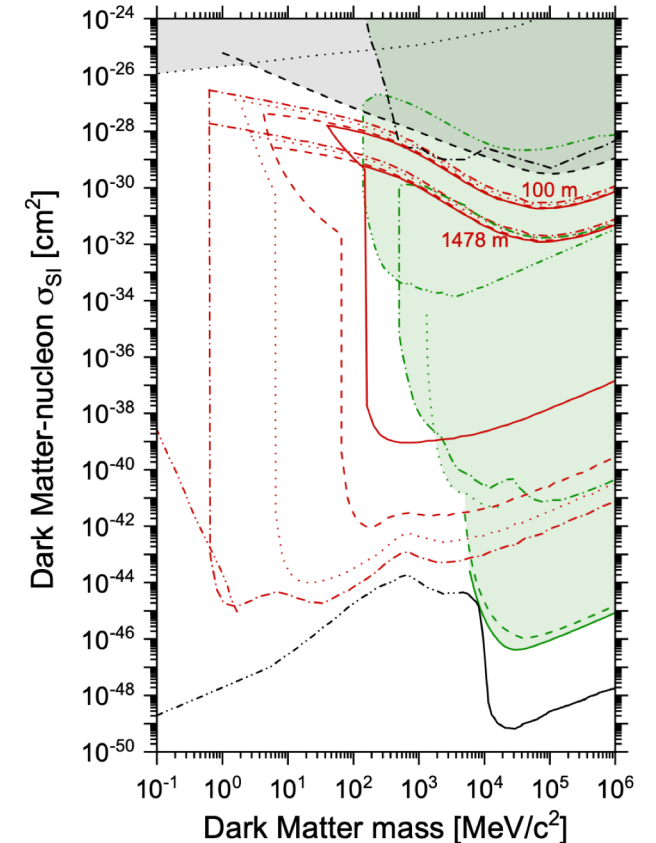
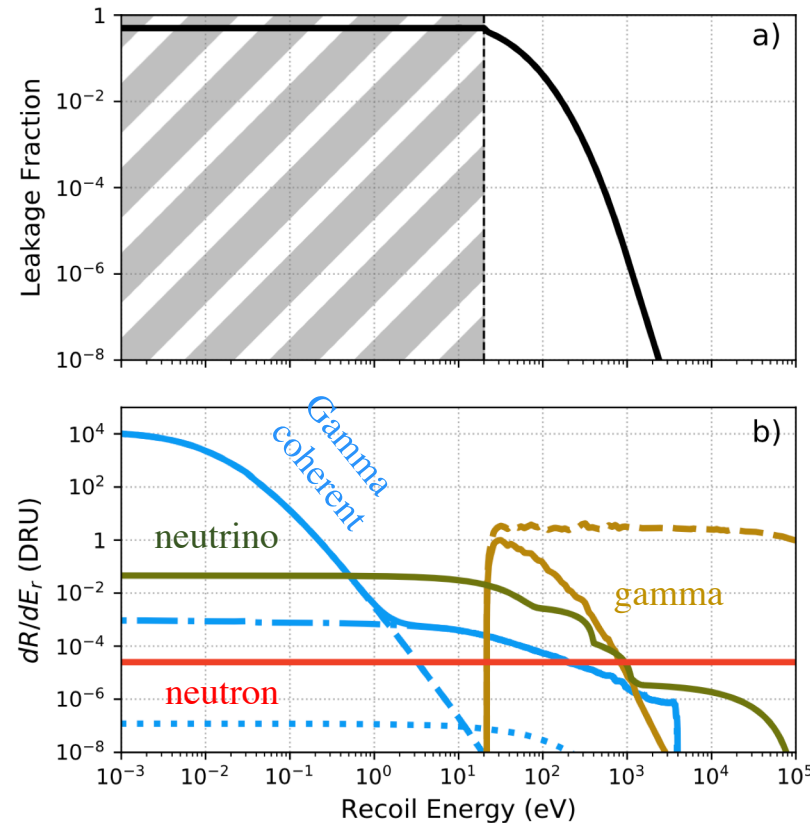
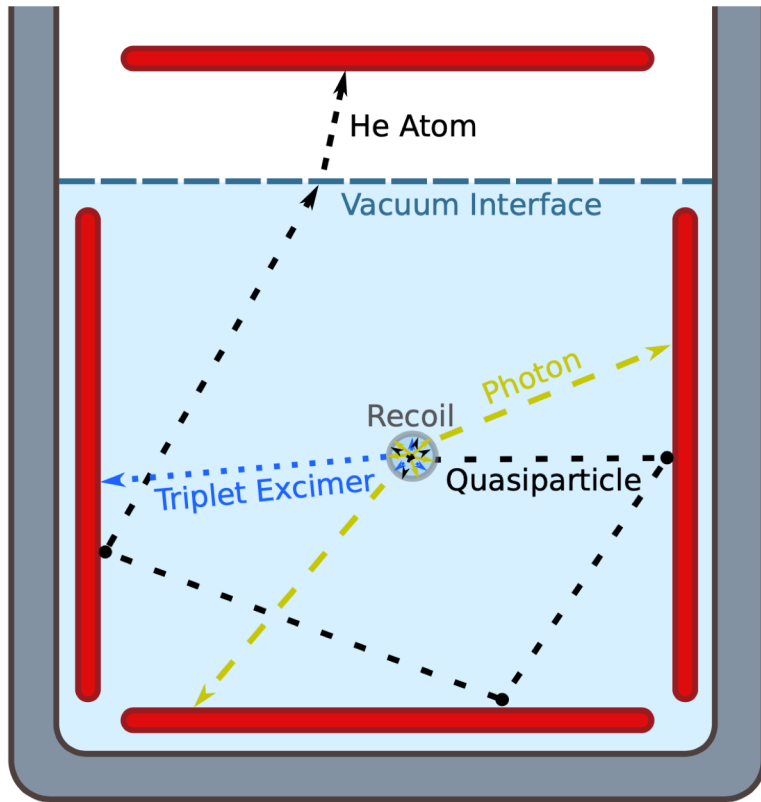
<https://news.fnal.gov/2020/06/sensei-gets-quiet/>



# Helium Detector



Phys. Rev. D **100**, 092007 (2019)



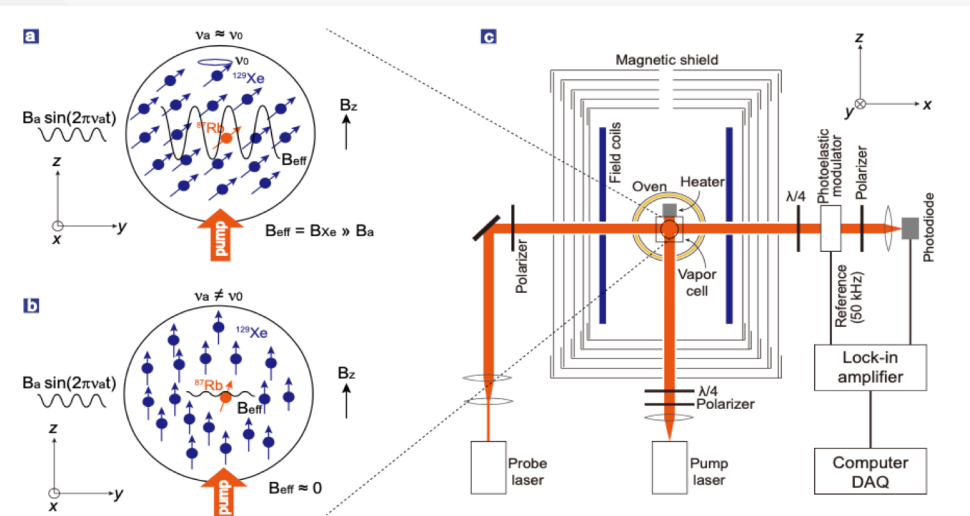
- Very low  $W$  value for roton and phonon;
- Quasiparticle propagate ballistically, including directionality;

- Background dominated by gamma and neutrino coherent scattering;

# Quantum sensors

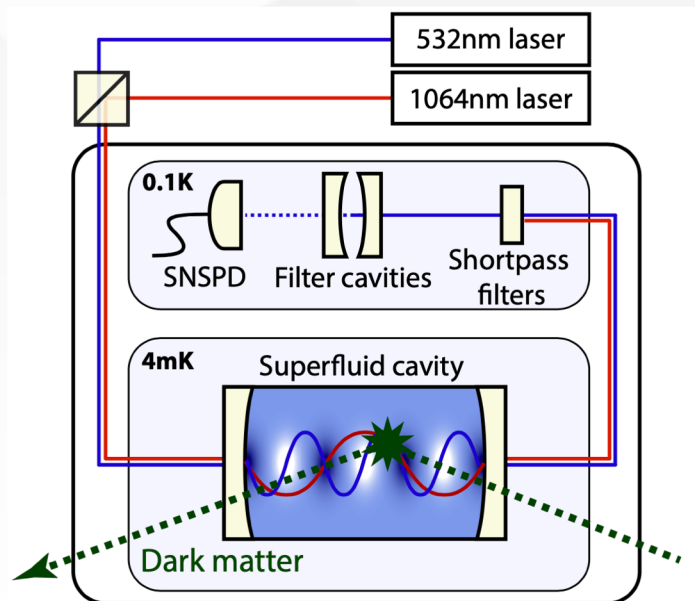


## Magnetometer



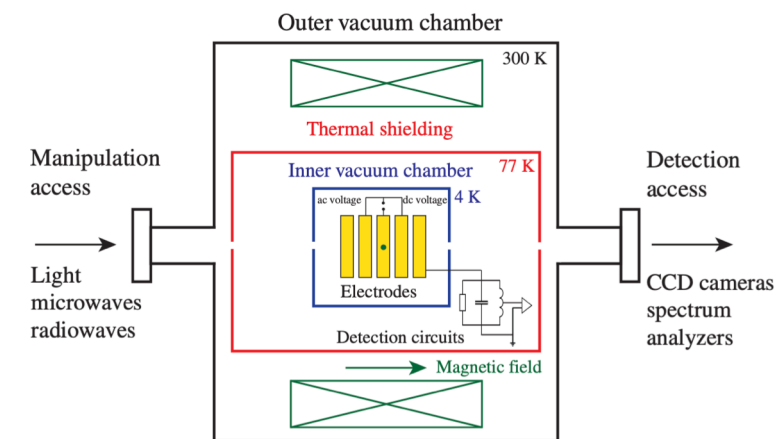
Nature Phys. 17, 1402–1407 (2021).

## Optomechanical



arXiv: 2306.09726

## Ion Trap



PRX QUANTUM 3, 010330 (2022)