

# Net Baryon Number Probability Distribution as an Indicator of Phase Transition

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# Outline

- 1 Roberge-Weiss approach.
- 2 Quantities under study:
  - ▶ net-baryon probability distributions  $\mathcal{P}_n$  and  $\mathbf{P}_n$ ;
  - ▶ their moments and cumulants;
  - ▶ their relation to the pressure and grand canonical partition function.
- 3 Evaluation of these quantities in lattice simulations.
- 4 Equation of State (EoS)  $p = f(\rho)$   
at  $T > T_{RW}$  and  $T < T_c$ .
- 5 Asymptotic behavior of  $\mathbf{P}_n$  at  $n \rightarrow \infty$ .
- 6 Phenomenological issues.

In collaboration with V.A.Goy

Roberge-Weiss approach in QCD at  $\mu_B \neq 0$ :

Fock space includes  
only colorless states  
at all  $T$  and  $\mu_B$ .

$$\theta \equiv \frac{\mu_B}{T} = \theta_R + i\theta_I$$

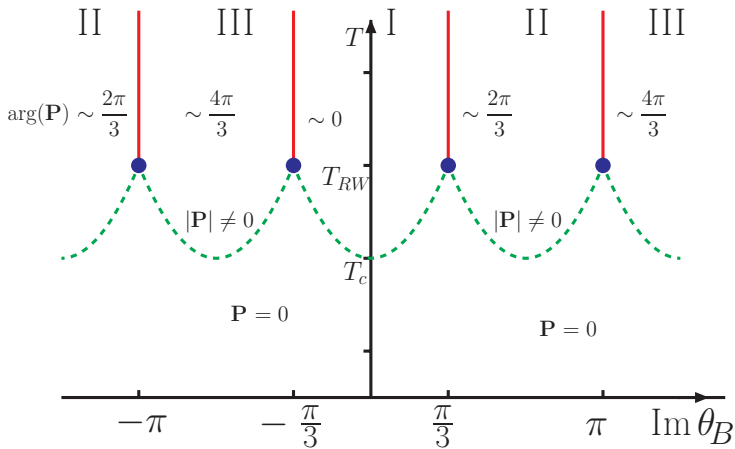
$$Z_{GC}(\theta_I) = Z_{GC}(\theta_I + 2\pi/N_c)$$



Quark number  $Q$  is a multiple of  $N_c$

Grand canonical partition function

$$Z_{GC}(\theta, T, V) = \sum_j \langle j | \exp \left( \frac{-\hat{H} + \mu \hat{Q}}{T} \right) | j \rangle$$



## The quantities under study:

- probability  $\mathcal{P}_n$  that the net baryon charge| of the fireball at a given  $\mu_B$  equals  $n$
- and the corresponding
  - ▶ moments  $\mu_k = \sum_{n=-\infty}^{+\infty} \mathcal{P}_n n^k$  and the respective moments generating function

$$M(t) = 1 + \sum_{k=1}^{\infty} \frac{\mu_k}{k!} t^k$$

- ▶ the cumulant generating function

$$K(t) = \ln M(t) = \sum_{k=1}^{\infty} \frac{\kappa_k}{k!} t^k$$

The probabilities  $\mathcal{P}_n$  can be determined from

- experimental data

$$\begin{aligned} N_{events}(\text{Net-Baryon Number} = n) &= \\ &= N_{events}(\text{Net-Proton Number} = 0.4n) \end{aligned}$$

- lattice simulations  
(of the net-baryon density at imaginary  $\mu_B$ )
- models of strong-interaction matter  
— Hadron Resonance Gas (HRG) model —

Grand canonical partition function

$Z_{GC}(\theta, T, V) \equiv Z_{GC}(\theta)$  can be expanded as follows:

$$Z_{GC}(\theta) = \exp\left(\frac{p(\theta)V}{T}\right) = \sum_{n=-\infty}^{\infty} z_C(n) e^{n\theta},$$

The inverse transform:

$$z_C(n) = \int_{-\pi}^{\pi} \frac{d\theta_I}{2\pi} e^{-in\theta_I} Z_{GC}(\theta) \Big|_{\theta_R=0}.$$

$$\theta = \mu_B/T = \theta_R + i\theta_I, \quad \rho = \frac{1}{T} \frac{\partial p}{\partial \theta}$$

$\mathcal{P}_n(\theta) = \frac{Z_C(n)e^{n\theta}}{Z_{GC}(\theta)}$  - is the probability that

the baryon charge at the given  $T$  and  $\mu_B$  equals  $n$ .

C-parity conservation implies  $Z_C(n) = Z_C(-n)$

$$\implies \frac{\mathcal{P}_n}{\mathcal{P}_{-n}} = \xi^{2n} \implies \mu_B = \frac{T}{2n} \ln \left( \frac{\mathcal{P}_n}{\mathcal{P}_{-n}} \right)$$

- possible procedure of measurement of  $\mu_B$   
[A.Nakamura, K.Nagato 2013]
- criterion of thermodynamical equilibrium:  
 $\mu_B$  measured for different  $n$  coincide



## Net-baryon probability distribution at $\mu_B = 0$

$\mathbf{P}_n \equiv \mathcal{P}_n(\theta = 0)$  involve all info on  $\theta$ -dependence:

$$\mathcal{P}_n(\theta) = \frac{Z_C(n)e^{n\theta}}{Z_{GC}(\theta)} = \mathbf{P}_n e^{n\theta} \frac{Z_{GC}(0)}{Z_{GC}(\theta)}$$

$$M_\theta(t) = \frac{Z_{GC}(t + \theta)}{Z_{GC}(\theta)} \longrightarrow \mathfrak{M}(t) = \frac{Z_{GC}(t)}{Z_{GC}(0)}$$

$$K_\theta(t) = \longrightarrow \mathfrak{K}(t) = \frac{(p(t) - p(0))V}{T}$$

$$\mathbf{P}_n = \frac{Z_C(n)}{Z_{GC}(0)}$$

and the respective cumulants  
in contrast to  $\theta$ -dependent cumulants  $\kappa_n(\theta)$   
coincide with the coefficients  
of the Taylor expansion of the pressure in  $\theta$ :

$$p(\theta) = p(0) + \sum_{n=1}^{\infty} \frac{\kappa_{2n}}{(2n)!} \theta^{2n}$$

Main attention is focused on

EXP.:  $\varkappa_n(\theta)$  at small  $n$  instead of  $\mathcal{P}_n(\theta)$

THEOR.:  $\kappa_n$  at small  $n$  instead of  $\mathbf{P}_n(\theta) \quad \forall n$

because  $\kappa_n = \varkappa_n(\mathbf{0})$

are related to the Taylor expansion of the pressure.

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**We argue that**

Asymptotic behavior of  $\mathbf{P}_n$  at  $n \rightarrow \infty$

may become an indicator of the chiral phase transition

Problem: Given  $\kappa_n$  find  $\mathbf{P}_n$

In lattice QCD at  $\text{Re}\mu_B = 0$ ,  $\text{Im}\mu_B \neq 0$   
we employ the formula

$$\mathbf{Z}_{GC}(\theta) = \int \mathbf{D}U e^{-S_G} (\det \mathcal{D}(\mu_B))^{N_f}$$

to find the net baryon number density  $\rho$   
and  $\implies$  the grand canonical partition function

$$\rho(\theta) = \frac{1}{V} \frac{\partial (T \ln \mathbf{Z}_{GC})}{\partial \mu_B} \implies$$
$$\mathbf{Z}_{GC}(\theta_I) |_{\theta_R=0} = \exp \left( V \int_0^{\theta_I} \rho(x) dx \right)$$

## Results of lattice simulations

$T = 1.35T_c > T_{RW}$  :  $\text{Im}\rho(\theta_I)$  is  $\frac{2\pi}{3}$ -periodic function  
with discontinuities at  $\theta_I = \frac{(2n+1)\pi}{3}$ ;

at  $|\theta_I| < \frac{\pi}{3}$  is well fitted by the polynomial

$$\text{Im}\rho(\theta_I) \simeq a_1\theta_I - a_3\theta_I^3$$

$T = 0.93T_c$  :  $\text{Im}\rho(\theta_I)$  is well fitted by the sine

$$\text{Im}\rho(\theta_I) \simeq f_1 \sin(\theta_I)$$

## Results of lattice simulations

$T > T_{RW}$  :  $\text{Im}\rho(\theta_I)$  is a periodic function  
fitted by the polynomial of the type

$$\text{Im}\rho(\theta_I) \simeq a_1\theta_I - a_3\theta_I^3 + \dots + \simeq a_n\theta_I^n$$

over each segment  $\theta_I^{(n-1)} < \theta_I < \theta_I^{(n)}$ ,

$$\text{where } \theta_I^{(n)} = \frac{(2n+1)\pi}{3};$$

$T \sim T_c$  :  $\text{Im}\rho(\theta_I)$  should be fitted by

$$\text{Im}\rho(\theta_I) \simeq f_1 \sin(\theta_I) + f_2 \sin(2\theta) + \dots + f_n \sin(n\theta) + \dots$$

where  $\{f_n\}$  rapidly decreases with  $n$ .

## Equation of State

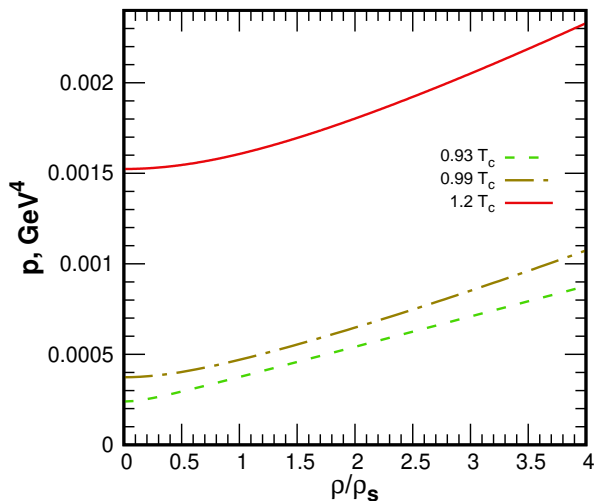
$$T = 1.35T_c > T_{RW} :$$

$$\begin{aligned}\frac{\rho}{T^3} &= a_1\theta + a_3\theta^3 \\ \frac{p}{T^4} &= \frac{a_1}{2}\theta^2 + \frac{a_3}{4}\theta^4 + \hat{p}_0,\end{aligned}$$

$$T = 0.93T_c :$$

$$\frac{p(\rho)}{T^4} = \hat{p}_0 + \left( \sqrt{\frac{\rho^2}{T^6} + f_1^2} - f_1 \right) .$$

here  $\hat{p}_0 = \left( \text{the pressure}/T^4 \right)$  at  $\theta = 0$ .



$\rho_s = 0.153 \text{ fm}^3$ ; data for  $\hat{p}_0$  are taken from HotQCD Collab.

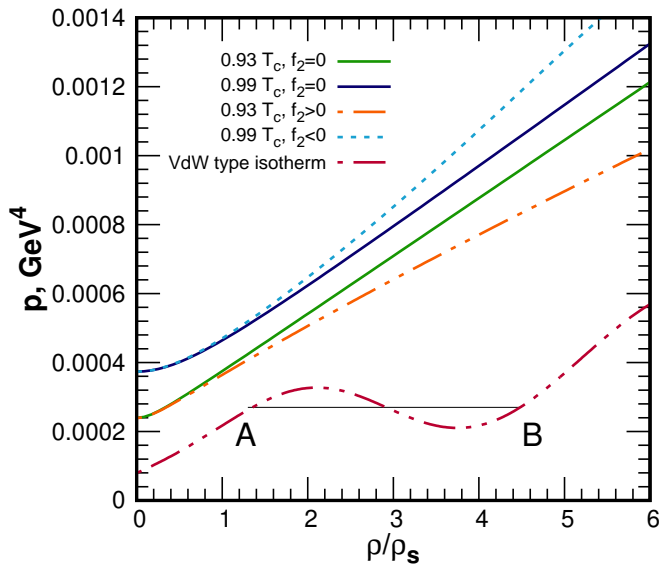


## Equation of State at $T \sim T_c$

$$T = 0.99T_c : \text{Im}\rho(\theta_I) \simeq f_1 \sin(\theta_I) + f_2 \sin(2\theta)$$
$$f_1 = 0.2541(8), \quad f_2 = -0.0053(7)$$

$$\hat{p} = f_1 s + 2f_2 s \sqrt{s^2 + 1};$$
$$\hat{p} = f_1(\sqrt{s^2 + 1} - 1) + f_2 s^2 + \hat{p}_0.$$

here  $\hat{p}_0 = p/T^4$  at  $\theta = 0$ ;  $s = \sinh(\theta)$ ;  $f_2 < 0$ .



$$T > T_{RW}: \quad \mathbf{P}_n \simeq \exp\left(-\frac{n^2}{2a_1 VT^3}\right), \quad n \ll VT^3$$

$$\mathbf{P}_n \simeq \exp\left(-\frac{3}{4} \sqrt[3]{\frac{3}{a_3}} \left(\frac{n}{VT^3}\right)^{4/3}\right), \quad \text{when } n \gg VT^3$$

$T < T_c$ : coincidence with the HRG,

$$\mathbf{P}_n \simeq e^{-A} I_n(A)^\dagger \quad \Longrightarrow \quad A = 2\sqrt{b\bar{b}}$$

$(\bar{b})b$  is the average number of the (anti)baryons in the fireball

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<sup>†</sup> [Bornyakov et al., 1611.04229]

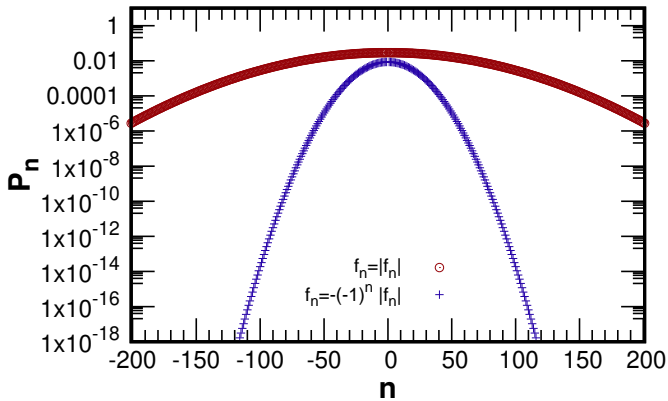
$$\text{Im}\rho(\theta_I) \simeq a_1\theta_I + \dots + a_{2J+1}\theta_I^{2J+1},$$

$$\text{sign } a_{2J+1} = (-1)^J$$

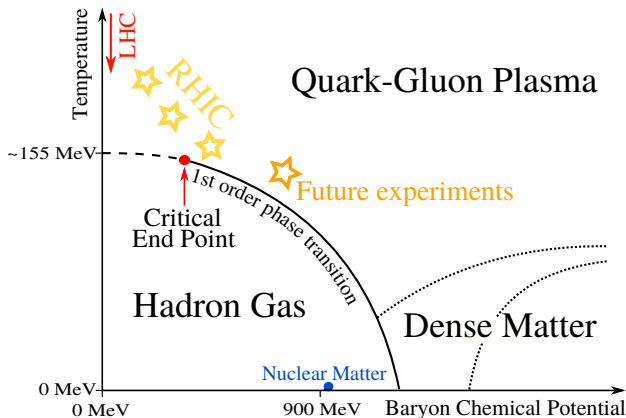
$$\mathbf{P}_n \sim \exp\left(-\frac{J}{J+1} \sqrt[J]{\frac{n^{J+1}}{\nu a_J}}\right) \quad \nu = VT^3.$$

$$\text{Im}\rho(\theta_I) \simeq f_1 \sin(\theta_I) \dots + f_J \sin(J\theta), \quad f_J > 0 \quad \forall J$$

$$\mathbf{P}_n \sim \frac{(\nu f_J)^{n/J}}{\Gamma\left(\frac{n}{J} + 1\right)}, \quad \nu = VT^3$$



# Hypothetical QCD phase diagram



Fireball evolution:  $T_{ini}, \mu_B^{(ini)} \longrightarrow T_F, \mu_B^{(F)}$

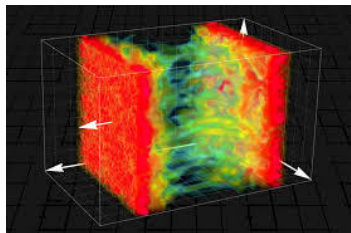
The Krein criterion states that the problem of moments becomes indeterminate when

$$\int d\mathbf{x} \frac{\ln \varphi(\mathbf{x})}{(1 + \mathbf{x}^2)} > -\infty, \quad (1)$$

where  $\varphi(\mathbf{x})$  is the probability density function.

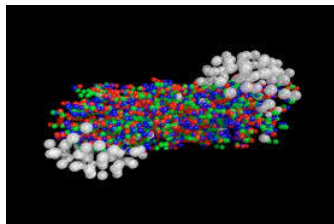
The rate of decrease in  $\mathbf{P}_n$  at low temperatures is very close to the line of demarcation between probability mass functions generating determinate and indeterminate moment problems

## Two scenarios of thermalization



1. **Exchange** of conserved charges ( $B, Q, S$ ) proceeds during the fireball expansion.

Grand canonical approach works down to  $T_{freezeout}$



2. **The fireball** after formation at an early stage is **isolated** from the remnants of colliding nuclei.

Evolution starts with the  $Z_{GC}(\mu_{ini}, T, V)$  and proceeds with  $Z_C(n, T, V)$ .



## Conclusions:

- Net-baryon number distribution  $\mathbf{P}_n$  is evaluated on a lattice at  $T > T_{RW}$  (it is similar to but doesn't coincide with the free theory) and at  $T < T_c$  (coincides with the HRG predictions).
- Reconstruction of  $\mathbf{P}_n$  from cumulants is either ambiguous or highly sensitive to small variations in higher-order cumulants. The analysis of experimental data based on the net-baryon number distribution involves additional information compared to that extracted from the set of cumulants.
- The dependence of the EoS on  $T$  and fit parameters has been used to formulate a possible scenario of emergence of the van der Waals isotherms corresponding to the first-order chiral phase transition.