

# Emergent Planck mass and dark energy from affine gravity

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*I. Kharuk, INR RAS, [ivan.kharuk@phystech.edu](mailto:ivan.kharuk@phystech.edu)*

# Affine gravity

Types of gravity, based on independent variables:

- ▶ Metric:

*GR*

$$g_{\mu\nu}, \quad \Gamma_{\mu\nu}^{\lambda} = \{\lambda_{\mu\nu}\}.$$

- ▶ Metric-affine:

*Palatini approach*

$$g_{\mu\nu}, \quad \Gamma_{\mu\nu}^{\lambda}.$$

- ▶ Affine:

*Einstein, Eddington,  
Schrodinger, ...*

$$\cancel{g_{\mu\nu}}, \quad \Gamma_{\mu\nu}^{\lambda}.$$

# Affine gravity

Lagrangian of the affine gravity:

$$\mathcal{L} = \sqrt{-\det (R_{(\mu\nu)}(\Gamma))}$$

Affine gravity is equivalent to GR with  $\Lambda \neq 0$ :

$$\sqrt{-g}(M_{pl}^2 R + 2\Lambda) \longleftrightarrow 2\frac{M_{pl}^4}{\Lambda} \sqrt{-\det R_{\mu\nu}}$$

## Affine gravity

Metric is the momentum, canonically conjugated to the connection:

$$\mathfrak{g}_{\lambda}^{\mu\nu\rho} = \frac{\delta\mathcal{L}}{\delta\partial_{\rho}\Gamma_{\mu\nu}^{\lambda}}, \quad \mathfrak{g}^{\mu\nu} = \frac{\delta\mathcal{L}}{\delta R_{(\mu\nu)}}.$$

Lagrange formulation:  $\mathcal{L}$ ,  $d\mathcal{L} = J_{\lambda}^{\mu\nu} d\Gamma_{\mu\nu}^{\lambda} + \mathfrak{g}^{\mu\nu} dR_{(\mu\nu)}$

Hamiltonian formulation:  $L = \mathcal{L} - \mathfrak{g}^{\mu\nu} R_{(\mu\nu)}$ ,

$$dL = J_{\lambda}^{\mu\nu} d\Gamma_{\mu\nu}^{\lambda} - R_{(\mu\nu)} d\mathfrak{g}^{\mu\nu}, \quad R_{(\mu\nu)} = -\frac{\delta L}{\delta\mathfrak{g}^{\mu\nu}}$$

## Pros and cons of affine gravity

- Natural from geometrical viewpoint
- Allows for a wide range of modifications
- Unified description of SM (**fermions**) and gravity

## Scalar-affine gravity

$$\varphi: \quad \varphi \rightarrow J\varphi, \quad \nabla_\mu \varphi = \partial_\mu \varphi - \Gamma_{\mu\sigma}^\sigma \varphi \quad (w = 1).$$

$$\text{Lagrangian:} \quad \mathcal{L} = \sqrt{-\det L_{\mu\nu}} + \varphi, \quad L_{\mu\nu} = R_{(\mu\nu)} - \frac{3}{2} \frac{\nabla_\mu \varphi \nabla_\nu \varphi}{\varphi^2}$$

Model is scale invariant (no dimensionfull parameters)

Spontaneous symmetry breaking:

$$-\nabla_\lambda \left( g^{\mu\nu} \frac{\delta L_{\mu\nu}}{\delta \nabla_\lambda \varphi} \right) + g^{\mu\nu} \frac{\delta L_{\mu\nu}}{\delta \varphi} + 1 = 0$$

⇓

$$\varphi = \varphi_0^4 e^\pi, \quad g^{\mu\nu} = -\varphi_0^2 \sqrt{-g} g^{\mu\nu}, \quad \varphi_0 \neq 0, \quad g_{\mu\nu} \neq 0.$$

# Scalar–affine gravity

Vacuum solution:

$$g_{\mu\nu} = \text{diag}(-1, a^2(t), a^2(t), a^2(t))$$

$$\Gamma_{ij}^0 = a^2(H + 3v)\gamma_{ij}, \quad \Gamma_{0j}^i = (H + v)\delta_j^i, \quad \Gamma_{00}^0 = -v, \quad H \equiv \frac{a'}{a},$$

$$6(3Hv + v') = -\kappa\varphi_0^2.$$

Master equation:

$$\left(\frac{H}{M_{pl}}\right)^2 = \frac{\kappa + 4}{6}.$$

# Conclusion

Scalar–affine gravity:

- ▶ realizes the **no–scale scenario**
- ▶ yields the emergent Planck mass,
- ▶ scalar  $\varphi$  acting as dark energy
- ▶ is phenomenologically viable (Newton's law and Linearized limit)