# **Antistars in the Galaxy**

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### Content

- Conventional baryogenesis
- OP violation in cosmology and possible antimatter domain
- Ossible discovery of antistars in the Galaxy.
- Antistar dark matter.
- A model of abundant antistar creation in galaxies.
- Conclusion

# Conventional baryogenesis

### A.D. Sakharov, three principles:

- Nonconservation of baryonic number.
- C and CP (explicit) violation.
- Deviation from thermal equilibrium.

"Cosmological constant":

$$\beta = N_B/N_{\gamma} = const$$

In simple versions of the scenario antibaryons are not created by the Sakharov mechanism.

Baryogenesis could proceed without any one of the Sakharov's conditions.

# C and CP violation in cosmology

Antimatter creation depends upon the way of charge symmetry violation:

- 1. Explicit.
- 2. Spontaneous.
- 3. Dynamical.

Explicit is achieved by complex coupling constants or masses.

Spontaneous - by spontaneous symmetry breaking when a complex scalar field acquires non-zero vacuum expectation value. The universe consists of matter and antimatter domain separated by walls which are to be eliminated.

Dynamical one is realized if a classical complex scalar field is somehow realized in the early universe e.g. due to quantum instability of massless fields at de Sitter stage (inflation), but this field disappears later leaving no trace. No domain wall problem. Regions with different values and signs of  $\beta$  can be created.

In versions 2 and 3 antistars are formed but far away from stars, in different galaxies.

# Discovery of antistars?

S. Dupourqué, L. Tibaldo, P. Von Ballmoos, Phys.Rev.D 103 (2021) 8, 083016 • e-Print: 2103.10073 [astro-ph.HE] Possible discovery of several antistars in the Galaxy was recently reported: "Constraints on the antistar fraction in the Solar System neighborhood from the 10-year Fermi Large Area Telescope gamma-ray source catalog." S. Dupourqué, L. Tibaldo, P. Von Ballmoos, Phys.Rev.D 103 (2021) 8, 083016 • e-Print: 2103.10073 [astro-ph.HE] "We identify in the catalog 14 antistar candidates not associated with any objects belonging to established gamma-ray source classes and with a spectrum compatible with baryon-antibaryon annihilation."

### Dark antimatter

In a recent publication a striking idea was put forward that dark matter may consist of compact anti-stars: J. S. Sidhu, R.J. Scherrer, G. Starkman, "Antimatter as Macroscopic Dark Matter", arXiv:2006.01200, astro-ph.CO. Such anti-DM may be easier to spot than other forms of macroscopic DM.

Existence of antistars in the Galaxy was predicted many years ago. Mechanism of massive PBH formation with log-normal mass spectrum: A. Dolgov and J.Silk, PRD 47 (1993) 4244 "Baryon isocurvature fluctuations at small scale and baryonic dark matter. A.Dolgov, M. Kawasaki, N. Kevlishvili, Nucl. Phys. B807 (2009) 229, "Inhomogeneous baryogenesis, cosmic antimatter, and dark matter". Massive PBHs allow to cure multpie inconsistencies with the standard cosmology and astrophysics. Unusual stellar type compact objects are also created, including abundant antistars in the Galaxy. An evidence in favor of the mechanism: The chirp mass distribution of LIGO events very well agrees with theoretical predictions, AD, A.G. Kuranov, N.A. Mitichkin, S. Porey, K.A. Postnov, et al JCAP 12 (2020) 017 - a strong support of the scenario.

# Bounds on antistars in the Galaxy

### As argued in:

- C. Bambi, A.D. Dolgov, "Antimatter in the Milky Way", Nucl.Phys.B 784 (2007) 132-150 e-Print: astro-ph/0702350,
- A.D. Dolgov, S.I. Blinnikov, "Stars and Black Holes from the very Early Universe", Phys.Rev.D 89 (2014) 2, 021301 e-Print: 1309.3395, S.I. Blinnikov, A.D. Dolgov, K.A. Postnov, "Antimatter and antistars in the universe and in the Galaxy", Phys.Rev.D 92 (2015) 2, 023516 e-print: 1409.5736
- an abundant density of compact anti-stars in the universe and even in the Galaxy does not violate existing observational limits. Such anti-DM may be easier to spot than other forms of macroscopic DM.
- Surface annihilation on a compact object is much less efficient than volume annihilation, e.g. inside gas cloud of antimatter.

# Antimatter creation by mirror matter

Antistars or antimatter cores in mirror neutron stars? Zurab Berezhiani (Jun 21, 2021) e-Print: 2106.11203 [astro-ph.HE]

The oscillation of the neutron n into mirror neutron n', its partner from dark mirror sector, can gradually transform an ordinary neutron star into a mixed star consisting in part of mirror dark matter. The implications of the reverse process taking place in the mirror neutron stars depend on the sign of baryon asymmetry in mirror sector. Namely, if it is negative, as predicted by certain baryogenesis scenarios, then  $\bar{n}' - \bar{n}$  transitions create a core of our antimatter gravitationally trapped in the mirror star interior.

The mechanism of massive PBH and antistar formation:

A. Dolgov and J.Silk, PRD 47 (1993) 4244 "Baryon isocurvature fluctuations at small scale and baryonic dark matter.

A.Dolgov, M. Kawasaki, N. Kevlishvili, Nucl.Phys. B807 (2009) 229, "Inhomogeneous baryogenesis, cosmic antimatter, and dark matter". Massive PBHs allow to cure multpie inconsistencies with the standard cosmology and astrophysics.

Unusual stellar type compact objects could also be created.

The model predicts the log-normal mass spectrum of PBH:

$$rac{dN}{dM} = \mu^2 \exp\left[-\gamma \ln^2(M/M_0)
ight],$$

and predicts  $M_0 \approx 10 M_{\odot}$ . A.Dolgov, K.Postnov, Why the mean mass of primordial black hole distribution is close to  $10 M_{\odot}$ . JCAP 07 (2020) 063 • e-Print: 2004.11669 . Very well agrees with observations.

SUSY motivated baryogenesis, Affleck and Dine (AD).

SUSY predicts existence of scalars with  $\mathbf{B} \neq \mathbf{0}$ . Such bosons may condense along flat directions of the quartic potential:

$$U_{\lambda}(\chi) = \lambda |\chi|^4 \left(1 - \cos 4\theta\right)$$

and of the mass term,  $U_m = m^2 \chi^2 + m^{*2} \chi^{*2}$ :

$$U_m(\chi) = m^2 |\chi|^2 [1 - \cos(2\theta + 2\alpha)],$$

where  $\chi=|\chi|\exp{(i\theta)}$  and  $m=|m|e^{\alpha}$ . If  $\alpha\neq 0$ , C and CP are broken. In GUT SUSY baryonic number is naturally non-conserved - non-invariance of  $U(\chi)$  w.r.t. phase rotation.

Initially (after inflation)  $\chi$  is away from origin and, when inflation is over, starts to evolve down to equilibrium point,  $\chi=\mathbf{0}$ , according to Newtonian mechanics:

$$\ddot{\chi} + 3H\dot{\chi} + U'(\chi) = 0.$$

Baryonic charge of  $\chi$ :

$$B_{\chi} = \dot{\theta} |\chi|^2$$

is analogous to mechanical angular momentum.  $\chi$  decays transferred baryonic charge to that of quarks in B-conserving process.

AD baryogenesis could lead to baryon asymmetry of order of unity, much larger than the observed  $10^{-9}\,$ .

If  $m \neq 0$ , the angular momentum, B, is generated by a different direction of the quartic and quadratic valleys at low  $\chi$ . If CP-odd phase  $\alpha$  is small but non-vanishing, both baryonic and antibaryonic domains might be formed with possible dominance of one of them.

Matter and antimatter domains may exist but globally B 
eq 0.

Affleck-Dine field  $\chi$  with CW potential coupled to inflaton  $\Phi$  (AD and Silk; AD, Kawasaki, Kevlishvili):

$$U = g|\chi|^{2}(\Phi - \Phi_{1})^{2} + \lambda|\chi|^{4} \ln(\frac{|\chi|^{2}}{\sigma^{2}}) + \lambda_{1}(\chi^{4} + h.c.) + (m^{2}\chi^{2} + h.c.).$$

Coupling to inflaton is the general renormalizable one.

When the window to the flat direction is open, near  $\Phi=\Phi_1$ , the field  $\chi$  slowly diffuses to large value, according to quantum diffusion equation derived by Starobinsky, generalized to a complex field  $\chi$ .

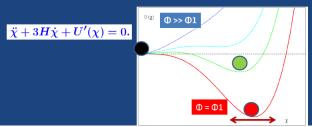
If the window to flat direction, when  $\Phi \approx \Phi_1$  is open only during a short period, cosmologically small but possibly astronomically large bubbles with high  $\beta$  could be created, occupying a small fraction of the universe, while the rest of the universe has normal  $\beta \approx 6 \cdot 10^{-10}$ , created by small  $\chi$ . The mechanism of massive PBH formation quite different from all others. The fundament of PBH creation is build at inflation by making large isocurvature fluctuations at relatively small scales, with practically vanishing density perturbations.

Initial isocurvature perturbations are in chemical content of massless quarks. Density perturbations are generated rather late after the QCD phase transition.

The emerging universe looks like a piece of Swiss cheese, where holes are high baryonic density objects occupying a minor fraction of the universe volume.

### Evolution of AD-field potential

Effective potential of  $\chi$  for different values of the inflaton field  $\Phi$ . The upper blue curve corresponds to a large value  $\Phi >> \Phi 1$  which gradually decreases down to  $\Phi = \Phi 1$ , red curve. Then the potential returns back to the almost initial shape, as  $\Phi 1$  drops down to zero. The evolution of  $\chi$  in such a potential is similar to a motion of a point-like particle (shown as a black ball in the figure) in Newtonian mechanics. First, due to quantum initial fluctuations  $\chi$  left the unstable extremum of the potential at  $\chi = 0$  and "tried" to keep pace with the moving potential minimum and later started to oscillate around it with decreasing amplitude. The decrease of the oscillation amplitude was induced by the cosmological expansion. In mechanical analogy the effect of the expansion is equivalent to the liquid friction term,  $3H\chi$ . When  $\Phi$  dropped below  $\Phi 1$ , the potential recovered its original form with the minimum at  $\chi = 0$  and  $\chi$  ultimately returned to zero but before that it could give rise to a large baryon asymmetry

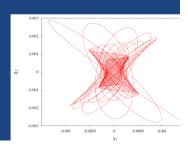


#### Evolution of AD-field

(Dolgov -Kawasaki-Kevlishvili)

Field  $\chi$  "rotates" in this plane with quite large angular momentum, which exactly corresponds to the baryonic number density of  $\chi$ . Later  $\chi$  decayed into quarks and other particles creating a large cosmological baryon asymmetry.

$$B_{\chi} = \dot{\theta} |\chi|^2$$



# Results of (Anti-)Creation

# The outcome, depending on $\beta = n_B/n_{\gamma}$ .

- PBHs with log-normal mass spectrum confirmed by the data!
- Compact stellar-like objects, as e.g. cores of red giants.
- Disperse hydrogen and helium clouds with (much) higher than average  $n_B$  density. Strange stars with unusual chemistry and velocity.
- $oldsymbol{\circ}$  may be negative leading to creation of (compact?) antistars which could survive annihilation with the homogeneous baryonic background. Density is model dependent but can be estimated using existing data on PBHs.

If true, the observation of antistars in the Galaxy is the discovery of the Century.

# More data are badly needed